

Interaction of partially ionized plasmas with electromagnetic fields

D Kremp¹, D Semkat¹, Th Bornath¹, M Bonitz², M Schlanges³ and P Hilse³

¹ Universität Rostock, Institut für Physik, D-18051 Rostock

² Christian-Albrechts-Universität zu Kiel, Institut für Theoretische Physik und Astrophysik, D-24098 Kiel

³ Ernst-Moritz-Arndt-Universität Greifswald, Institut für Physik, D-17489 Greifswald

E-mail: dietrich.kremp@uni-rostock.de

Abstract. The interaction of partially ionized plasmas with an electromagnetic field is investigated using quantum statistical methods. A general statistical expression for the current density of a plasma in an electromagnetic field is presented and considered in the high field regime. Expressions for the collisional absorption are derived and discussed. Further, partially ionized plasmas are considered. Plasma Bloch equations for the description of bound-free transitions are given and the absorption coefficient as well as rate coefficients for multiphoton ionization are derived and numerical results are presented.

1. Introduction

Due to the enormous progress in short-pulse laser physics, the interaction of matter with electromagnetic fields is a topic of rapidly growing interest. In this paper we consider a partially ionized plasma under the influence of an external laser field.

A partially ionized plasma is characterized by the densities of its constituents: electrons (n_e), ions (n_i), and atoms (n_a). The temporal change of these densities due to ionization and recombination is determined by rate equations. In the field-free case, these processes are caused by three-particle collisions (impact ionization). Then the rate equation for the electron density has the following form

$$\frac{dn_e}{dt} = \alpha n_e n_a + \beta n_e n_e n_i$$

with the ionization α and recombination β coefficients. In dense plasmas, these coefficients are essentially determined by the interaction. The resulting density dependence of α can be seen in Fig. 1 [1].

In the stationary case, from the rate equation a mass action law follows which determines the equilibrium composition of the partially ionized plasma. This is shown in Fig. 2 [1].

Under the action of an external electromagnetic field given by, e.g.,

$$\mathbf{A}(t) = \mathbf{A}_0 \sin(\omega t), \quad \mathbf{E}(t) = -\frac{1}{c} \frac{\partial}{\partial t} \mathbf{A}(t) = \mathbf{E}_0 \cos(\omega t), \quad \mathbf{E}_0 = -\frac{\omega \mathbf{A}_0}{c}$$

the situation is more complicated. Several additional interesting processes occur, e.g. (i) radiational induced transitions between bound and scattering states, i.e. radiation ionization

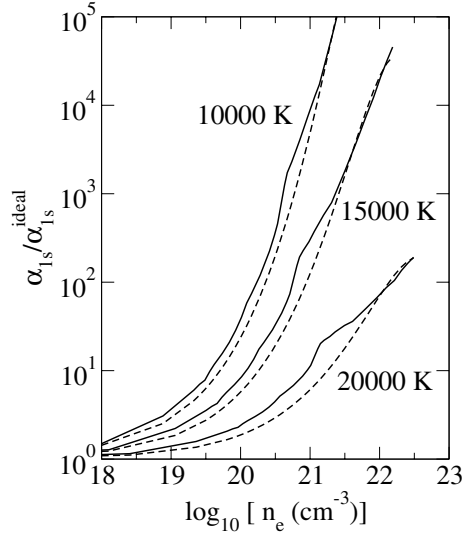


Figure 1. Ionization coefficient vs. density for several temperatures

and recombination, (ii) excitation and deexcitation processes of atoms in the plasma, and (iii) processes between scattering states like (inverse) bremsstrahlung (collisional absorption).

All these processes are connected with an energy exchange between plasma and field which is determined by the electrical current \mathbf{j} ,

$$\frac{dW^{\text{kin}}}{dt} + \frac{dW^{\text{pot}}}{dt} = \mathbf{j} \cdot \mathbf{E}, \quad (1)$$

i.e., the change of the total energy of the system of particles is equal to $\mathbf{j} \cdot \mathbf{E}$ which is in turn the energy loss of the electromagnetic field due to Poynting's theorem. From the quantity \mathbf{j} we get, as is well known, further important quantities like the polarization $d\mathbf{P}/dt = \mathbf{j}$, the absorption, reflection and refraction coefficients.

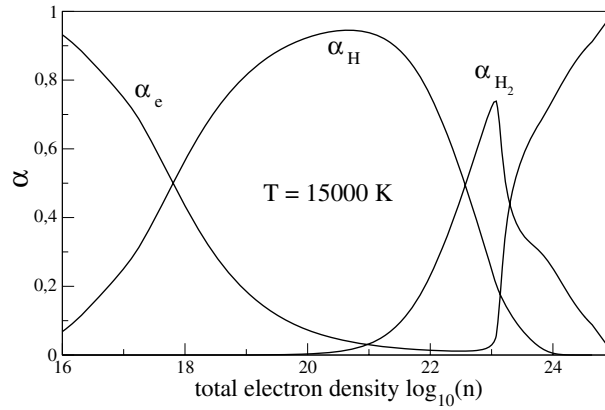


Figure 2. Equilibrium composition of a hydrogen plasma at $T = 15000$ K

A central quantity, therefore, is the electrical current defined by

$$\mathbf{j} = \sum_a \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} \frac{\mathbf{p}_a}{m_a} f_a(p_a, t), \quad (2)$$

where $f_a(p_a, t)$ is the single-particle gauge invariant Wigner function of the species a . For the determination of f_a we start from the kinetic equation [2, 3]

$$\left\{ \frac{\partial}{\partial t} + e_a \mathbf{E}(t) \cdot \nabla_{\mathbf{k}_a} \right\} f_a(\mathbf{k}_a, t) = I_a(\mathbf{k}_a, t), \quad (3)$$

where the collision integral is given by (\mathcal{V} being the volume and V_{ab} the interaction potential)

$$I_a(\mathbf{p}, t) = \sum_b \frac{1}{\mathcal{V}} \int \frac{d\mathbf{q} d\mathbf{p}_1 d\mathbf{p}_2}{(2\pi\hbar)^6} [\delta(\mathbf{p} - \mathbf{p}_1 - \mathbf{q}) - \delta(\mathbf{p} - \mathbf{p}_1)] \times V_{ab}(\mathbf{q}) F_{ab}(\mathbf{p}_1, \mathbf{p}_1 + \mathbf{q}, \mathbf{p}_2, \mathbf{p}_2 - \mathbf{q}, t). \quad (4)$$

The two-particle density matrix is connected with a two-particle correlation function by

$$F_{ab}(t) = (i\hbar)^2 g_{ab}^<(12t, 1'2't) = (i\hbar)^2 g_a^<(1t, 1't) g_b^<(2t, 2't) \pm \text{exch.} + (i\hbar)^2 g_{ab}^{\text{corr}}(12t, 1'2't),$$

where we denoted $1 = \mathbf{r}_1, s_1$ etc. For special physical situations, it is easier to derive appropriate approximations for other functions. In place of the correlation part of $g_{ab}^<$, we can put any single-time two-particle correlation function, especially also $L_{ab}^< - L_{ab}^{0<}$ which is the correlation function of density fluctuations [4, 5],

$$\begin{aligned} (i\hbar) L_{ab}^<(11't, 22't) &= \langle \delta \hat{\rho}_b(22't) \delta \hat{\rho}_a(11't) \rangle \\ \text{with } \delta \hat{\rho}_a(11't) &= \Psi_a^+(1', t) \Psi_a(1, t) - \langle \Psi_a^+(1', t) \Psi_a(1, t) \rangle \end{aligned} \quad (5)$$

In the momentum balance, there are contributions from collisions between different species only. For the case $a \neq b$, we have

$$F_{ab}(t) = (i\hbar)^2 g_a^<(1t, 1't) g_b^<(2t, 2't) + (i\hbar) L_{ab}^<(11't, 2'2't). \quad (6)$$

Explicit expressions for the collision integral where given in the classical case first by Silin [6] and later by Klimontovich [7] and in the quantum mechanical case in the papers [2, 8]. In generalization to the well-known Boltzmann-like collision integrals, they describe short-time processes, i.e., they are non-Markovian [9, 10, 11, 12].

In a partially ionized plasma, we have to account for bound and scattering states and transitions between them caused by the external field. Therefore, it is useful to expand F_{ab} with respect to the eigenstates of the unperturbed two-particle Hamiltonian

$$F_{ab}(\mathbf{p}_1, \mathbf{p}_1 + \mathbf{q}, \mathbf{p}_2, \mathbf{p}_2 - \mathbf{q}, t) = \sum_{\mathbf{P}\alpha, \mathbf{P}'\alpha'} \langle \mathbf{p}_1 \mathbf{p}_2 | \alpha \mathbf{P} \rangle \langle \mathbf{P} \alpha | F_{ab}(t) | \alpha' \mathbf{P}' \rangle \langle \mathbf{P}' \alpha' | \mathbf{p}_1 - \mathbf{q}, \mathbf{p}_2 + \mathbf{q} \rangle \quad (7)$$

with $\alpha = \{n, l, m\}$ for bound states and $\alpha = \mathbf{p}$ for scattering states. As we will show later on, due to the external field, there are nonvanishing nondiagonal elements $F_{ab}^{\alpha\alpha'}$. They are connected with transitions between different states α and α' .

The balance equation for the electrical current density follows in well-known manner from Eq. (3). Using, for instance, the function L_{ab} it reads [4]

$$\frac{d}{dt} \mathbf{j}_a(t) - n_a \frac{e_a^2}{m_a} \mathbf{E}(t) = \sum_{b \neq a} \int \frac{d\mathbf{q}}{(2\pi\hbar)^3} \frac{e_a \mathbf{q}}{m_a} V_{ab}(q) L_{ba}^<(\mathbf{q}; t, t). \quad (8)$$

Further progress is connected with the determination of the functions $L_{ba}^<$ and $g_{ab}^<$, respectively.

In the following, we will consider the free–free transition using approximations for $L_{ba}^<$ in the next section. Then the current is calculated from Eq. (8), and the collisional absorption (inverse bremsstrahlung) in the linear response regime as well as in the high–field case is determined and discussed.

In the third section, bound states are included. The equation of motion for $g_{ab}^<$ is the Bethe–Salpeter equation (BSE) for plasmas which has been discussed in detail in [13]. We will derive plasma Bloch equations which form a system of equations for the determination of the matrix elements of (7). Here, the equations for the diagonal elements describe the time evolution of the occupation of bound and scattering states, respectively, driven by the non-diagonal elements. The coupled system, therefore, describes the ionization kinetics of a partially ionized plasma.

2. Collisional absorption in a fully ionized plasma

Let us first consider the energy exchange between field and plasma due to transitions between scattering states. This problem has been dealt with in several papers, first for classical plasmas, e.g. [14, 6, 7, 15, 16]. Quantum mechanical treatments were given in more recent papers, for the strong field case, e.g., [17, 18, 19], and in linear response theory [20].

Starting point for our investigation is Eq. (8). The r.h.s. of this equation is determined by the function $L^<$. This quantity follows from the Bethe–Salpeter equation on the Keldysh contour

$$L_{ab}(12, 1' 2') = \Pi_{ab}(12, 1' 2') + \sum_{cd} \int_C d3 d4 \Pi_{ac}(13, 1' 3^+) V_{cd}(34) L_{db}(42, 4^+ 2'). \quad (9)$$

In a plasma in a strong laser field, an approximation in lowest order of V_{ie} is appropriate because the coupling between species with different charges is weaker than their coupling to the field. The coupling between particles with equal charges in the subsystem, however, is not affected by the field. Then the so-called polarization functions Π_{ab} can be adopted to be diagonal, $\Pi_{ab} = \delta_{ab} \Pi_a$. The solution has now the following structure (for brevity all arguments are suppressed)

$$L_{ei} = \mathcal{L}_{ee} V_{ei} \mathcal{L}_{ii}; \quad \mathcal{L}_{aa} = \Pi_a + \Pi_a V_{aa} \mathcal{L}_{aa}; \quad (a = e, i). \quad (10)$$

Going back from the contour to the physical time axis using the Langreth–Wilkins rules we find for the correlation functions of the density fluctuation [5]

$$L_{ei}^{\gtrless}(\mathbf{q}; t, t') = \int d\bar{t} \left[\mathcal{L}_{ee}^{\gtrless}(\mathbf{q}; t, \bar{t}) V_{ei}(q) \mathcal{L}_{ii}^A(\mathbf{q}; \bar{t}, t') + \mathcal{L}_{ee}^R(\mathbf{q}; t, \bar{t}) V_{ei}(q) \mathcal{L}_{ii}^{\gtrless}(\mathbf{q}; \bar{t}, t') \right]. \quad (11)$$

For the electron current (8) there follows that

$$\begin{aligned} \frac{d}{dt} \mathbf{j}_e(t) - n_e \frac{e_e^2}{m_e} \mathbf{E}(t) &= \text{Re} \int \frac{d^3 q}{(2\pi\hbar)^3} \frac{e_e \mathbf{q}}{m_e \hbar} V_{ei}(q) 2\pi i \int_{t_0}^t d\bar{t} \left[\mathcal{S}_{ee}(\mathbf{q}; t, \bar{t}) V_{ei}(q) \mathcal{L}_{ii}^A(\mathbf{q}; \bar{t}, t) \right. \\ &\quad \left. + \mathcal{L}_{ee}^R(\mathbf{q}; t, \bar{t}) V_{ei}(q) \mathcal{S}_{ii}(\mathbf{q}; \bar{t}, t) \right], \end{aligned} \quad (12)$$

where we introduced the dynamical structure factor

$$2\pi \mathcal{S}_{aa}(\mathbf{q}; t, \bar{t}) = \frac{i\hbar}{2} \left[\mathcal{L}_{aa}^>(\mathbf{q}; t, \bar{t}) + \mathcal{L}_{aa}^<(\mathbf{q}; t, \bar{t}) \right]. \quad (13)$$

The functions in the collision term depend on the current and the electrical field, respectively. This dependence can be made explicit if one assumes that each subsystem (electrons and ions) is in local equilibrium with a temperature T_a with respect to a coordinate frame moving with the mean velocity $\mathbf{u}_a(t)$ [1, 21]. The transformation between such a coordinate system and a

system at rest is given by $\tilde{\mathbf{r}} = \mathbf{r} - \int_{t_0}^t d\bar{t} \mathbf{u}_a(\bar{t})$. The Fourier transforms in the two coordinate systems are connected by

$$\mathcal{L}_{aa}(\mathbf{q}, t_1 t_2) = e^{-\frac{i}{\hbar} \mathbf{q} \cdot \int_{t_2}^{t_1} d\bar{t} \mathbf{u}_a(\bar{t})} \tilde{\mathcal{L}}_{aa}(\mathbf{q}, t_1 - t_2), \quad (14)$$

where $\tilde{\mathcal{L}}_{aa}$ denotes the local equilibrium function depending on the time difference only. One gets (omitting the tilde from now on)

$$\begin{aligned} \frac{d}{dt} \mathbf{j}(t) = & \varepsilon_0 \omega_p^2 \mathbf{E}(t) - \text{Re} \int \frac{d^3 q}{(2\pi\hbar)^3} \mathbf{q} \frac{e_e}{m_e} V_{ie}(q) \frac{2\pi}{i\hbar} \int_{t_0}^t d\bar{t} \left[\mathcal{S}_{ee}(\mathbf{q}; t - \bar{t}) V_{ei}(q) \right. \\ & \times \mathcal{L}_{ii}^A(\mathbf{q}; \bar{t} - t) + \mathcal{L}_{ee}^R(\mathbf{q}; t - \bar{t}) V_{ei}(q) \mathcal{S}_{ii}(\mathbf{q}; \bar{t} - t) \left. \right] \exp \left\{ -\frac{i}{\hbar} \frac{1}{n_e e_e} \mathbf{q} \cdot \int_{\bar{t}}^t d\bar{t}_1 \mathbf{j}(\bar{t}_1) \right\}. \end{aligned} \quad (15)$$

Thus the source term in the current balance equation depends on the current itself in a non-linear way.

We will consider this non-linear equation in the following in two limiting cases.

2.1. Linear response case

Let us first consider the energy exchange in a weak laser field. To describe the situation we introduce the quiver velocity

$$\mathbf{v}_0 = \frac{e_e \mathbf{E}_0}{m_e \omega}. \quad (16)$$

Then the weak field case is characterized by $v_0/v_{\text{th}} \ll 1$ where v_{th} is the electron thermal velocity. Therefore, the situation corresponds to the linear response regime. In this case we can expand the exponential function in Eq. (ball),

$$\exp \left\{ -\frac{i}{\hbar} \frac{1}{n_e e_e} \mathbf{q} \cdot \int_{\bar{t}}^t d\bar{t}_1 \mathbf{j}(\bar{t}_1) \right\} \approx 1 - \frac{i}{\hbar} \frac{1}{n_e e_e} \mathbf{q} \cdot \int_{\bar{t}}^t d\bar{t}_1 \mathbf{j}(\bar{t}_1). \quad (17)$$

Adopting a harmonic time dependence of the electric field, we have for the current $\mathbf{j}(t) = \mathbf{j}(\omega)e^{-i\omega t} + \mathbf{j}^*(\omega)e^{i\omega t}$ and get from the above equation

$$\mathbf{j}(\omega) = \frac{\varepsilon_0 \omega_p^2}{-i\omega + \nu_{ei}(\omega)} \mathbf{E}(\omega), \quad (18)$$

which is a generalized Drude equation with a complex electron-ion collision frequency ν_{ei} given by

$$\nu_{ei}(\omega) = i \frac{n_i e_i^2}{n_e m_e \hbar^2} \int \frac{d^3 q}{(2\pi\hbar)^3} \frac{1}{\omega} (\mathbf{q} \cdot \mathbf{n})^2 \mathcal{S}_{ii}(q) V(q) \left[\varepsilon_{ee}^{-1}(\mathbf{q}; \omega) - \varepsilon_{ee}^{-1}(\mathbf{q}; 0) \right] \quad (19)$$

with the dielectric function of the electron subsystem

$$\varepsilon_{ee}^{-1}(\mathbf{q}; \omega) = 1 + e^2 V(q) \mathcal{L}_{ee}(\mathbf{q}; \omega), \quad (20)$$

and the static ion-ion structure factor $\mathcal{S}_{ii}(q)$. Using for ε_{ee}^{-1} the RPA, this is the well-known expression first given by Bekefi [22] and later for instance by Reinholz et al. [20]. One gets immediately the dynamical conductivity

$$\sigma(\omega) = \frac{\varepsilon_0 \omega_p^2}{-i\omega + \nu_{ei}(\omega)}. \quad (21)$$

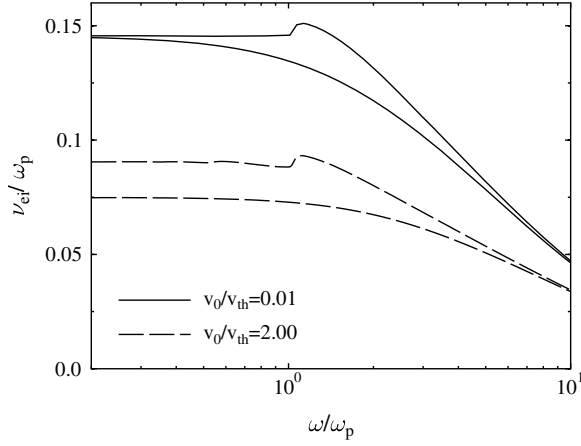


Figure 3. Electron-ion collision frequency as a function of the laser frequency for a hydrogen plasma ($n_e = 10^{21} \text{cm}^{-3}$; $T = 10^5 \text{ K}$) for two different field strengths. The upper curve of each pair corresponds to the full dynamical screening, the lower one to the static screening approximation.

In the limit of high frequencies, one gets $\sigma(\omega) \approx (\varepsilon_0 \omega_p^2 / \omega^2) [i\omega + \nu_{ei}(\omega)]$ [14], and the averaged absorbed energy is given by

$$\langle \mathbf{j} \cdot \mathbf{E} \rangle \equiv \frac{1}{T} \int_{t-T}^T dt' \mathbf{j}(t') \cdot \mathbf{E}(t') = \frac{\omega_p^2}{\omega^2} \text{Re } \nu_{ei}(\omega) \frac{\varepsilon_0 \mathbf{E}_0^2}{2} = \frac{\omega_p^2}{\omega^2} \text{Re } \nu_{ei}(\omega) \langle \varepsilon_0 \mathbf{E}^2 \rangle. \quad (22)$$

The dependence of the quantity $\text{Re } \nu_{ei}$ on the laser frequency is shown in Fig. 3 for two different relations v_0/v_{th} . Only the upper set of curves correspond to the linear response formula (19). Obviously, static and dynamical screening exhibit a qualitatively different behavior. The dynamical screening leads to a maximum in the vicinity of the plasma frequency, $\omega \approx \omega_p$. The lower set of curves is beyond the linear response regime. This strong field case will be considered in the next paragraph.

2.2. Strong high-frequency fields

The strong field case is characterized by $v_0/v_{th} \gtrsim 1$. For the first time such a situation was discussed by Silin [6] in the framework of classical kinetic theory.

For a strong high-frequency laser field a linearization is not possible. Instead, we follow the idea of Silin and assume that the influence of the collisions may be considered as a small perturbation compared to the external field. We decompose the current according to $\mathbf{j} = \mathbf{j}_0 + \mathbf{j}_1$ and assume

$$\mathbf{j} \approx \mathbf{j}^0 = \sum_a \frac{e_a^2 n_a}{m_a} \int_{t_0}^t dt' \mathbf{E}(t'). \quad (23)$$

For a harmonic electric field, $\mathbf{E} = \mathbf{E}_0 \cos \omega t$, the exponential factor in Eq. (15) can be expanded into a Fourier series. The current balance is given then by

$$\begin{aligned} \frac{d}{dt} \mathbf{j}_e(t) - n_e \frac{e_e^2}{m_e} \mathbf{E}(t) &= \text{Re} \int \frac{d^3 q}{(2\pi \hbar)^3} \frac{2\pi e_e \mathbf{q}}{m_e \hbar} V_{ei}^2(q) \sum_m \sum_n (-i)^{m+1} J_n \left(\frac{\mathbf{q} \cdot \mathbf{v}_0}{\hbar \omega} \right) J_{n-m} \left(\frac{\mathbf{q} \cdot \mathbf{v}_0}{\hbar \omega} \right) \\ &\times e^{im\omega t} \int_{-\infty}^{\infty} \frac{d\bar{\omega}}{2\pi} [\mathcal{S}_{ee}(\mathbf{q}; \bar{\omega} - n\omega) \mathcal{L}_{ii}^A(\mathbf{q}; \bar{\omega}) + \mathcal{L}_{ee}^R(\mathbf{q}; \bar{\omega} - n\omega) \mathcal{S}_{ii}(\mathbf{q}; \bar{\omega})] \end{aligned} \quad (24)$$

with the one-component structure factors and response functions \mathcal{S}_{aa} and \mathcal{L}_{aa} , respectively [4, 5]. We will assume in the following of this section that the subsystems are in local

thermodynamic equilibrium with temperatures T_e and T_i , respectively (the influence of non-Maxwellian distribution functions was considered in [23], and a numerical solution of a kinetic equation for a strong laser field was performed in [24]). J_l is the Bessel function of l th order. In the above equation, electron and ion functions contribute formally equal to the screening. The ion functions, however, are localized in the low-frequency region, i.e., for a high-frequency electric field, $\bar{\omega}$ can be neglected in comparison with $n\omega$. In this case, the first term in the brackets in Eq. (24) vanishes because $\int d\bar{\omega} \mathcal{L}_{ii}^A(\mathbf{q}; \bar{\omega}) = 0$, and for the current it follows that

$$\mathbf{j}_e(t) - \int_{-\infty}^t d\bar{t} \frac{n_e e_e^2}{m_e} \mathbf{E}(\bar{t}) = \text{Re} \int \frac{d^3 q}{(2\pi\hbar)^3} \sum_m \sum_n \frac{e_e}{m_e} \frac{\mathbf{q}}{m\hbar\omega} V_{ei}^2(q) (-i)^{m+2} e^{im\omega t} \times J_n \left(\frac{\mathbf{q} \cdot \mathbf{v}_0}{\hbar\omega} \right) J_{n-m} \left(\frac{\mathbf{q} \cdot \mathbf{v}_0}{\hbar\omega} \right) \mathcal{L}_{ee}^R(\mathbf{q}; -n\omega) n_i \mathcal{S}_{ii}(\mathbf{q}) \quad (25)$$

where screening by the ions is accounted for by the static structure factor $\mathcal{S}_{ii}(\mathbf{q})$ defined by

$$\mathcal{S}_{ii}(\mathbf{q}) \equiv \frac{1}{n_i} \int d\bar{\omega} \mathcal{S}_{ii}(\mathbf{q}, \bar{\omega}) = 1 + n_i \int d^3 r [g_{ii}(\mathbf{r}) - 1] e^{-\frac{i}{\hbar} \mathbf{q} \cdot \mathbf{r}}, \quad (26)$$

where g_{ii} is the pair distribution function. \mathcal{L}_{ee}^R is the exact density response function of the electron subsystem. Appropriate approximations can be expressed via local field corrections [5].

Equation (25) is clearly an expansion of the current in terms of higher harmonics of the laser frequency

$$\mathbf{j}_e(t) - \int_{-\infty}^t d\bar{t} \frac{n_e e_e^2}{m_e} \mathbf{E}(\bar{t}) = \sum_{m=-\infty}^{\infty} \mathbf{j}_m(t) e^{im\omega t}. \quad (27)$$

The Fourier coefficients \mathbf{j}_m of the current (with $\mathbf{j}_m = \mathbf{j}_{-m}^*$) can be easily identified from (25). One can show that only the odd harmonics are allowed due to the symmetry of the interaction, cf. [18]. The appearance of higher harmonics in strong laser fields was first observed by Silin [6] and is a very interesting effect.

Let us now consider again the cycle averaged energy (22). This quantity is now given by

$$\langle \mathbf{j} \cdot \mathbf{E} \rangle \equiv \frac{1}{T} \int_{t-T}^t dt' \mathbf{j}(t') \cdot \mathbf{E}(t') = \mathbf{E}_0 \cdot \text{Re} \mathbf{j}_1 = \frac{\omega_p^2}{\omega^2} \text{Re} \nu_{ei}(\omega) \frac{\varepsilon_0 \mathbf{E}_0^2}{2}. \quad (28)$$

Interesting is here the restriction of the sum over m to its first term, i.e., higher harmonics do not contribute to the average energy exchange.

Using the explicit expression for $\text{Re} \mathbf{j}_1$ we obtain

$$\langle \mathbf{j} \cdot \mathbf{E} \rangle = n_i \int \frac{d^3 q}{(2\pi\hbar)^3} V_{ii}(q) \mathcal{S}_{ii}(\mathbf{q}, T_i) \sum_{n=-\infty}^{\infty} n\omega J_n^2 \left(\frac{\mathbf{q} \cdot \mathbf{v}_0}{\hbar\omega} \right) \text{Im} \varepsilon_{ee}^{-1}(\mathbf{q}, -n\omega, T_e). \quad (29)$$

Finally, the effective electron-ion collision frequency can be obtained using Eq. (28).

In Eq. (29), ε_{ee}^{-1} is the full quantum mechanical dielectric function of the electrons, e.g. the RPA one. The physical meaning of the sum over n follows from the frequency argument $n\omega$ of the dielectric function. It indicates that every term of the sum describes the absorption of n photons (multi-photon absorption).

In order to compare with the classical theory, it is advantageous to consider the non-degenerate case in which some integrations can be done analytically. For the case of a Maxwellian distribution function, we get

$$\langle \mathbf{j} \cdot \mathbf{E} \rangle = \frac{8\sqrt{2\pi} Z^2 e^4 n_e n_i \sqrt{m_e}}{(4\pi\varepsilon_0)^2 (k_B T)^{3/2}} \omega^2 \sum_{n=1}^{\infty} n^2 \int_0^{\infty} \frac{dk}{k^3} \frac{1}{|\varepsilon(k, n\omega)|^2}$$

$$\times e^{-\frac{n^2 m_e \omega^2}{2k_B T k^2}} e^{-\frac{\hbar^2 k^2}{8m_e k_B T}} \frac{\sinh \frac{n\hbar\omega}{2k_B T}}{\frac{n\hbar\omega}{2k_B T}} \int_0^1 dz J_n^2 \left(\frac{eE_0 k}{m_e \omega^2} z \right). \quad (30)$$

Quantum effects, marked by \hbar , occur here at several places. The first place is one of the exponential functions describing the quantum diffraction effects at large momenta k . This exponential function ensures the convergence of the integral. The second place is the sinh term which is connected with the Bose statistics of multiple photon emission and absorption. Finally, quantum effects enter via the calculation of dielectric function itself.

In the limit $\hbar \rightarrow 0$ we get from Eq. (29) well-known classical results derived for the first time by Klimontovich [7] and later by Decker et al. [15]. The classical formulas exhibit the well-known problem of divergencies at large k which has to be overcome by some cut-off procedure. In contrast, in our quantum approach, no divergencies exist.

Unfortunately, a further analytical simplification is only possible in limiting cases. Therefore, we present in the following some results of the numerical evaluation of Eq. (29) [25, 26]. In Fig. 4, the collision frequency is shown as a function of the quiver velocity. For comparison, results from Silin [6] and Decker [15] are given. Our quantum expression (29) was evaluated with the full dynamical and the static RPA dielectric function. The frequency dependence has been shown

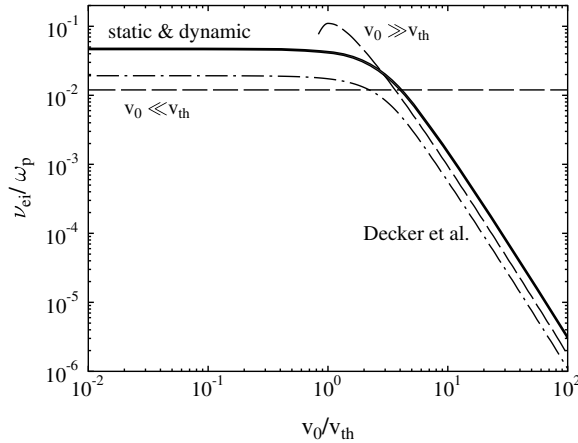


Figure 4. Real part of the electron-ion collision frequency as a function of the quiver velocity $v_0 = eE/\omega m_e$ for a hydrogen plasma in a laser field ($Z = 1; n_e = 10^{22} \text{ cm}^{-3}; T = 3 \cdot 10^5 \text{ K}; \omega/\omega_p = 5$). For comparison, results of Decker et al. (dash-dotted line) and the asymptotic formulae of Silin [6] (dashed line) are given.

already in Fig. 3. Again, we observe the resonance maximum at the plasma frequency. The collision frequency as a function of the coupling parameter $\Gamma = \left(\frac{4\pi n}{3}\right)^{1/3} \frac{e^2}{k_B T}$ is drawn in Fig. 5. According to this figure our quantum mechanical results are in good agreement with numerical simulations in this special situation. In contrast, the classical theories break down for $\Gamma \gtrsim 0.2$. This behavior results from the cut-off procedure at large momentum k inherent in classical approaches.

The influence of the Bose character of the photons [hyperbolic sine term in (30)] can be seen in Fig. 6. An interesting consequence is the plateau-like behavior of the collision frequency with a subsequent sharp drop down in dependence of the photon number. Neglecting the sinh term destroys this behavior.

3. Bound-free transitions. Ionization kinetics

3.1. Basic equations

So far we have considered transitions between scattering states. Now we take into account bound states and consider bound-free transitions in a partially ionized plasma, i.e., radiation

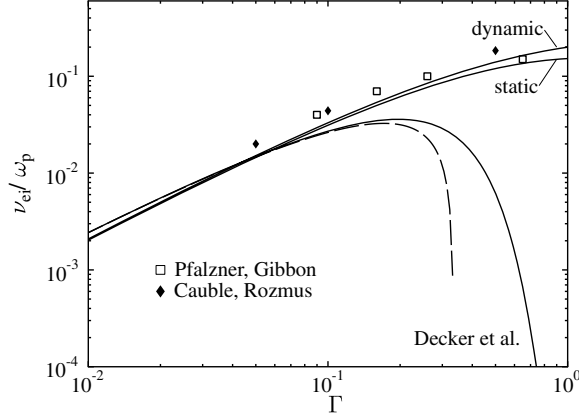


Figure 5. Electron–ion collision frequency as a function of the coupling parameter Γ for a hydrogen plasma in a laser field ($Z = 1$; $v_0/v_{th} = 0.2$; $n_e = 10^{22} \text{ cm}^{-3}$; $\omega/\omega_p = 3$). A comparison is given with the theory of Decker et al. and with the asymptotic formula of Silin [6] (dashed line) is given. Furthermore, results of Cauble and Rozmus [27] and simulation data of Pfalzner and Gibbon [28] are shown.

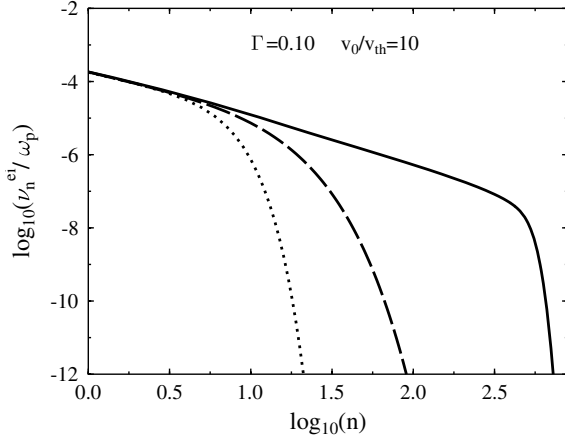


Figure 6. Contributions ν_n vs. photon number n in a hydrogen plasma ($n_e = 10^{22} \text{ cm}^{-3}$; $\omega/\omega_p = 5$; $\Gamma = 0.1$) for $v_0/v_{th} = 10$. Present approach (solid line), sinh term neglected (dashed line); classical dielectric theory (dotted line)

ionization and recombination processes. Therefore, our task is to develop a kinetic description of these processes, especially we have to determine rate and absorption coefficients of radiation ionization. To investigate the problem of build-up and decay of bound states, an appropriate starting point of our considerations is the equation of motion of the two-particle correlation function $g_{ab}^<$, well-known as Bethe–Salpeter equation. In order to take into account bound states, we have to consider the two-particle correlation function $g_{ab}^<$ at least in binary collision approximation [10, 29]. Then the BSE can be written in the form [13, 1]

$$\left[i\hbar \frac{\partial}{\partial t} - H_{ab}^0 - U_{ab} - N_{ab} V_{ab} - (\Sigma_a^{\text{HF}}(t) + \Sigma_b^{\text{HF}}(t)) \right] g_{ab}^<(t, t') = 0. \quad (31)$$

Here, $\Sigma_a^{\text{HF}}(t)$ is the Hartree–Fock self-energy, and $N_{ab} = 1 - F_a - F_b$ the Pauli blocking factor. It is convenient to introduce relative and center-of-mass coordinates \mathbf{p} and \mathbf{P} , respectively, which yields

$$H_{ab}^0 = \frac{\mathbf{P}^2}{2M} + \frac{\mathbf{p}^2}{2\mu}; \quad U_{ab} = -\mathbf{A} \cdot \left(\frac{e_t \mathbf{P}}{M} + \frac{e_r \mathbf{p}}{\mu} \right) + \frac{A^2}{2} \left(\frac{e_t^2}{M} + \frac{e_r^2}{\mu} \right), \quad (32)$$

where M and μ are the total and reduced masses, and $e_t = e_a + e_b$ the total and e_r the reduced charge defined by $Me_r = m_b e_a - m_a e_b$, respectively.

Now we take the time-diagonal limit of the two-particle correlation function, $F_{ab}(t) = (i\hbar)^2 g_{ab}^<(t, t)$. From Eq. (31) there follows that [3]

$$\begin{aligned} i\hbar \frac{\partial}{\partial t} F_{ab}(t) - [H_{ab}^0 + V_{ab}, F_{ab}(t)] - [U_{ab}(t), F_{ab}(t)] \\ - [(F_a(t) + F_b(t)) V_{ab}, F_{ab}(t)] - [(\Sigma_a^{\text{HF}}(t) + \Sigma_b^{\text{HF}}(t)), F_{ab}(t)] = 0, \end{aligned} \quad (33)$$

which describes the kinetics of the two-particle properties like bound and scattering states in the plasma. In Eq. (33), the first two terms represent the usual binary collision approximation, the third term contains the direct external field contribution, whereas the further terms arise from many-particle effects: Pauli blocking and Hartree–Fock (mean field + exchange) self-energy.

For the further investigation, we choose the presentation (7) the two-particle density matrix basis, where the bound and scattering eigenstates of the field-free two-particle Hamiltonian are given by

$$(H_{ab}^0 + V_{ab}) |\mathbf{P}\alpha\rangle = E_{\alpha\mathbf{P}} |\mathbf{P}\alpha\rangle, \quad (34)$$

$$|\mathbf{P}\alpha t\rangle = e^{-\frac{i}{\hbar} E_{\alpha\mathbf{P}} t} |\mathbf{P}\alpha\rangle. \quad (35)$$

Here, α denotes a complete set of quantum numbers n, l, m in the case of bound states and the relative momentum \mathbf{p} in the case of scattering states, respectively,

$$\alpha = \begin{cases} \{n, l, m\} & \text{for bound states} \\ \mathbf{p} & \text{for scattering states.} \end{cases}$$

In the representation (34), we get for the equation (33) [30]

$$\left[i\hbar \frac{\partial}{\partial t} - (E_{\alpha\mathbf{P}} - E_{\alpha'\mathbf{P}}) \right] F_{ab}^{\alpha\alpha'}(\mathbf{P}; t) = \sum_{\bar{\alpha}} \left[\hbar\Omega_{\mathbf{R}}^{\alpha\bar{\alpha}}(\mathbf{P}; t) F_{ab}^{\bar{\alpha}\alpha'}(\mathbf{P}; t) - F_{ab}^{\alpha\bar{\alpha}}(\mathbf{P}; t) \hbar\tilde{\Omega}_{\mathbf{R}}^{\bar{\alpha}\alpha'}(\mathbf{P}; t) \right], \quad (36)$$

with the matrix elements given by

$$\langle \alpha\mathbf{P} | F_{ab}(t) | \mathbf{P}'\alpha' \rangle = F_{ab}^{\alpha\alpha'}(\mathbf{P}; t) (2\pi\hbar)^3 \delta(\mathbf{P} - \mathbf{P}'), \quad (37)$$

$$\begin{aligned} \langle \alpha\mathbf{P} | U_{ab}(t) + \Sigma_a^{\text{HF}}(t) + \Sigma_b^{\text{HF}}(t) \pm [F_a(t) + F_b(t)] V_{ab} | \mathbf{P}'\alpha' \rangle \\ = \hbar\Omega_{\mathbf{R}}^{\alpha\alpha'}(\mathbf{P}; t) (2\pi\hbar)^3 \delta(\mathbf{P} - \mathbf{P}'), \end{aligned} \quad (38)$$

$$\begin{aligned} \langle \alpha\mathbf{P} | U_{ab}(t) + \Sigma_a^{\text{HF}}(t) + \Sigma_b^{\text{HF}}(t) \pm V_{ab} [F_a(t) + F_b(t)] | \mathbf{P}'\alpha' \rangle \\ = \hbar\tilde{\Omega}_{\mathbf{R}}^{\alpha\alpha'}(\mathbf{P}; t) (2\pi\hbar)^3 \delta(\mathbf{P} - \mathbf{P}'). \end{aligned} \quad (39)$$

Here, $\hbar\Omega_{\mathbf{R}}$ is an abbreviation which can be interpreted as a generalized Rabi energy, i.e., the field contribution renormalized by Hartree–Fock and Pauli blocking terms. Notice that U_{ab} in Eqs. (38,39) does not contain the term quadratic in A (cf. Eq. (32)), because it cancels out in the homogeneous case in the commutator. The quantities defined in (38,39) are related to each other by $[\hbar\Omega_{\mathbf{R}}^{\alpha\alpha'}(\mathbf{P}; t)]^* = \hbar\tilde{\Omega}_{\mathbf{R}}^{\alpha'\alpha}(\mathbf{P}; t)$.

Thus, the representation chosen above leads to a transformation of (33) into a matrix equation. Due to the inclusion of the scattering continuum, F_{ab} is, in principle, a matrix of infinite rank. If there are $N - 1$ bound states, we obtain N^2 coupled equations. Even if one makes use of the symmetry relation $[F_{ab}^{\alpha\alpha'}(\mathbf{P}; t)]^* = F_{ab}^{\alpha'\alpha}(\mathbf{P}; t)$, there remain $\frac{N}{2}(N + 1)$ equations.

Obviously, the diagonal elements $F_{ab}^{\alpha\alpha}$ are related to the occupation numbers of the respective state. Due to the external field we have nonvanishing nondiagonal elements $F_{ab}^{\alpha\alpha'}$. They are connected with transitions between two states α and α' . Here, we have to consider several situations:

(i) If both α and α' denote bound states, the equations describe excitation and deexcitation processes of atoms in the plasma. Then we recover the familiar atomic Bloch equations [31, 32, 33].

(ii) If both α and α' denote continuum (scattering) states, processes like (inverse) bremsstrahlung are described.

(iii) In the case most interesting for our investigations, however, α denotes a bound state and α' a scattering state (or vice versa). The transitions between them, i.e. ionization and recombination processes, are thus included in Eq. (36).

3.2. Plasma Bloch equations

3.2.1. Derivation of the plasma Bloch equations Since we are especially interested in the ionization kinetics, in the following, we do not include (de)excitation and (de)acceleration processes, i.e., transitions between discrete states and transitions within the scattering continuum. Therefore, we consider a system with one single bound state and denote

$$\alpha = \begin{cases} b & \text{"bound"} & \text{bound state} \\ f & \text{"free"} & \text{scattering state.} \end{cases}$$

Obviously, for that system only the physics beyond the familiar atomic Bloch equations remains from Eq. (36). We get a system of three coupled equations [30]:

$$i\hbar \frac{\partial}{\partial t} F_{bb}(\mathbf{P}; t) = 2i \int \frac{d\bar{\mathbf{p}}}{(2\pi\hbar)^3} \text{Im} \left\{ \hbar\Omega_R^{bf}(\mathbf{P}, \bar{\mathbf{p}}; t) F_{fb}(\mathbf{P}, \bar{\mathbf{p}}; t) \right\}, \quad (40)$$

$$\begin{aligned} & \left[i\hbar \frac{\partial}{\partial t} - (E(\mathbf{p}) - E(\mathbf{p}')) \right] F_{ff}(\mathbf{P}, \mathbf{p}, \mathbf{p}'; t) \\ & - \int \frac{d\bar{\mathbf{p}}}{(2\pi\hbar)^3} \left[\hbar\Omega_R^{ff}(\mathbf{P}, \mathbf{p}, \bar{\mathbf{p}}; t) F_{ff}(\mathbf{P}, \bar{\mathbf{p}}, \mathbf{p}'; t) - F_{ff}(\mathbf{P}, \mathbf{p}, \bar{\mathbf{p}}; t) \hbar\tilde{\Omega}_R^{ff}(\mathbf{P}, \bar{\mathbf{p}}, \mathbf{p}'; t) \right] \\ & = \hbar\Omega_R^{fb}(\mathbf{P}, \mathbf{p}; t) F_{bf}(\mathbf{P}, \mathbf{p}'; t) - F_{fb}(\mathbf{P}, \mathbf{p}; t) \hbar\tilde{\Omega}_R^{bf}(\mathbf{P}, \mathbf{p}'; t), \end{aligned} \quad (41)$$

$$\begin{aligned} & \left[i\hbar \frac{\partial}{\partial t} - (E(\mathbf{p}) - \tilde{E}_b) \right] F_{fb}(\mathbf{P}, \mathbf{p}; t) - \int \frac{d\bar{\mathbf{p}}}{(2\pi\hbar)^3} \hbar\Omega_R^{ff}(\mathbf{P}, \mathbf{p}, \bar{\mathbf{p}}; t) F_{fb}(\mathbf{P}, \bar{\mathbf{p}}; t) \\ & = \hbar\Omega_R^{fb}(\mathbf{P}, \mathbf{p}; t) F_{bb}(\mathbf{P}; t) - \int \frac{d\bar{\mathbf{p}}}{(2\pi\hbar)^3} F_{ff}(\mathbf{P}, \mathbf{p}, \bar{\mathbf{p}}; t) \hbar\tilde{\Omega}_R^{fb}(\mathbf{P}, \bar{\mathbf{p}}; t), \end{aligned} \quad (42)$$

$$F_{bf}(\mathbf{P}, \mathbf{p}; t) = [F_{fb}(\mathbf{P}, \mathbf{p}; t)]^*, \quad (43)$$

where \tilde{E}_b denotes the binding energy renormalized by Hartree–Fock and Pauli blocking contributions,

$$\begin{aligned} \tilde{E}_b = & E_b + \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} |\varphi_b(\mathbf{p})|^2 \left\{ \Sigma_1^{\text{HF}}(\mathbf{P}, \mathbf{p}; t) + \Sigma_2^{\text{HF}}(\mathbf{P}, \mathbf{p}; t) \right. \\ & \left. \mp [E(\mathbf{p}) - E_b] [F_1(\mathbf{P}, \mathbf{p}; t) + F_2(\mathbf{P}, \mathbf{p}; t)] \right\} \end{aligned} \quad (44)$$

with $F_{1/2}(\mathbf{P}, \mathbf{p}; t) = F_{1/2}\left(\frac{m_{1/2}}{M}\mathbf{P} \pm \mathbf{p}; t\right)$ and $\Sigma_{1/2}^{\text{HF}}(\mathbf{P}, \mathbf{p}; t) = \Sigma_{1/2}^{\text{HF}}\left(\frac{m_{1/2}}{M}\mathbf{P} \pm \mathbf{p}; t\right)$. Here, $\varphi_b(\mathbf{p}) = \langle \mathbf{p} | n \rangle$ being the Fourier transform of the bound state wave function.

Keeping in mind the relation (43), we have obtained a system of three equations for three unknown functions the physical meaning of which is quite obvious. F_{bb} is, in principle, equivalent to the distribution function of the atoms in the plasma. F_{ff} is the binary distribution of the unbound electron–ion pairs. From the system (40–43) we see that the dynamics of both distributions is driven by the function F_{fb} describing the transition between bound and continuum states. The latter quantity is closely connected with the polarization function. Its time evolution, in turn, is determined by the distributions. Notice that, even if we neglect three-particle collisions, ionization and recombination takes place due to the presence of the external field contained in the Rabi energies $\hbar\Omega_R$.

3.2.2. Field terms. Renormalized Rabi energies As we have seen before, the matrix elements of the field contribution to the Hamiltonian U_{ab} can be represented in the form of Rabi energies $\hbar\Omega_R$. Since we consider electron–ion pairs, U_{ab} from Eq. (32) simplifies to

$$U_{ab} = -\frac{e}{\mu} \mathbf{A} \cdot \mathbf{p}.$$

Thus, the Rabi energies are given by

$$\begin{aligned} \hbar\Omega_R^{ff}(\mathbf{P}, \mathbf{p}, \mathbf{p}'; t) &= -\frac{e}{\mu} \mathbf{A}(t) \cdot \langle +\mathbf{p} | \mathbf{p} | \mathbf{p}' + \rangle + \int \frac{d\bar{\mathbf{p}}}{(2\pi\hbar)^3} \varphi_{\mathbf{p}+}^*(\bar{\mathbf{p}}) \varphi_{\mathbf{p}'+}(\bar{\mathbf{p}}) \\ &\times \{ \Sigma_1^{\text{HF}}(\mathbf{P}, \bar{\mathbf{p}}; t) + \Sigma_2^{\text{HF}}(\mathbf{P}, \bar{\mathbf{p}}; t) \mp [E(\bar{\mathbf{p}}) - E(\mathbf{p})] [F_1(\mathbf{P}, \bar{\mathbf{p}}; t) + F_2(\mathbf{P}, \bar{\mathbf{p}}; t)] \}, \end{aligned} \quad (45)$$

$$\begin{aligned} \hbar\Omega_R^{fb}(\mathbf{P}, \mathbf{p}; t) &= -\frac{e}{\mu} \mathbf{A}(t) \cdot \langle +\mathbf{p} | \mathbf{p} | n \rangle + \int \frac{d\bar{\mathbf{p}}}{(2\pi\hbar)^3} \varphi_{\mathbf{p}+}^*(\bar{\mathbf{p}}) \varphi_b(\bar{\mathbf{p}}) \\ &\times \{ \Sigma_1^{\text{HF}}(\mathbf{P}, \bar{\mathbf{p}}; t) + \Sigma_2^{\text{HF}}(\mathbf{P}, \bar{\mathbf{p}}; t) \mp [E(\bar{\mathbf{p}}) - E(\mathbf{p})] [F_1(\mathbf{P}, \bar{\mathbf{p}}; t) + F_2(\mathbf{P}, \bar{\mathbf{p}}; t)] \} \end{aligned} \quad (46)$$

with $\varphi_{\mathbf{p}+}(\mathbf{p})$ being the Fourier transform of the scattering wave function. These relations show that the bare field terms are renormalized by many-particle effects in form of Hartree–Fock energies and Pauli blocking contributions.

3.3. Analysis of the transition function

In order to investigate the full ionization kinetics, the system (40–43) has to be solved selfconsistently. This is, of course, a very complicated task especially due to the momentum dependences of the quantities which is, up to now, numerically not feasible. Therefore, we approach the problem by analyzing limiting cases which simplify the system (40–43) and allow for an insight into the underlying physics.

First let us assume an external field given by

$$\mathbf{A}(t) = \mathbf{A}_0 \sin(\omega t), \quad \mathbf{E}(t) = -\frac{\partial}{\partial t} \mathbf{A}(t) = \mathbf{E}_0 \cos(\omega t), \quad \mathbf{E}_0 = -\omega \mathbf{A}_0. \quad (47)$$

Further, we look at Eq. (42). Introducing the approximations (i) neglect of Hartree–Fock and Pauli blocking renormalizations, (ii) replacement of scattering states $|\mathbf{p}+\rangle$ by free momentum states $|\mathbf{p}\rangle$, and (iii) neglect of binary correlations in the scattering state, i.e.,

$$F_{ff}(\mathbf{P}, \mathbf{p}, \mathbf{p}'; t) \approx F_1(\mathbf{P}, \mathbf{p}; t) F_2(\mathbf{P}, \mathbf{p}; t) (2\pi\hbar)^3 \delta(\mathbf{p} - \mathbf{p}'),$$

this equation can be written as

$$\begin{aligned} &\left[i\hbar \frac{\partial}{\partial t} - (E(\mathbf{p}) - E_b) + \frac{e}{\mu} \mathbf{A}(t) \cdot \mathbf{p} \right] F_{fb}(\mathbf{P}, \mathbf{p}; t) \\ &= -\frac{e}{\mu} \mathbf{A}(t) \cdot \mathbf{p} \varphi_b(\mathbf{p}) \{ F_{bb}(\mathbf{P}; t) - F_1(\mathbf{P}, \mathbf{p}; t) F_2(\mathbf{P}, \mathbf{p}; t) \}. \end{aligned} \quad (48)$$

The formal solution of Eq. (48) is given by

$$F_{fb}(\mathbf{P}, \mathbf{p}; t) = -\frac{1}{i\hbar\mu} \frac{e}{\mu} \int_{t_0}^t d\bar{t} e^{-\frac{i}{\hbar}(E(\mathbf{p})-E_b)(t-\bar{t})} e^{\frac{i}{\hbar}\frac{\mathbf{e}}{\mu}\mathbf{p}\cdot\bar{\mathbf{t}}} \bar{\mathbf{A}}(\bar{\mathbf{t}}) \mathbf{A}(\bar{\mathbf{t}}) \cdot \mathbf{p} \varphi_b(\mathbf{p}) \\ \times \{F_{bb}(\mathbf{P}; \bar{t}) - F_1(\mathbf{P}, \mathbf{p}; \bar{t})F_2(\mathbf{P}, \mathbf{p}; \bar{t})\}. \quad (49)$$

For monochromatic fields (Eqs. (47)), the integral in the second exponent on the r.h.s. of (49) can be solved:

$$\frac{e}{\mu} \mathbf{p} \cdot \int_{\bar{t}}^t d\bar{\mathbf{t}} \bar{\mathbf{A}}(\bar{\mathbf{t}}) = \frac{\mathbf{v}_0 \cdot \mathbf{P}}{\omega} (\cos \omega t - \cos \omega \bar{t}), \quad (50)$$

where \mathbf{v}_0 is the so-called quiver velocity, $\mathbf{v}_0 = \frac{e\mathbf{E}_0}{\mu\omega}$.

Keeping in mind that the field is switched on adiabatically, i.e. at $t_0 \rightarrow -\infty$, we have to write for the vector potential

$$\mathbf{A}(t) = \lim_{\epsilon \rightarrow 0} \mathbf{A}_0 \frac{1}{2i} \left[e^{i(\omega - i\epsilon)t} - e^{-i(\omega + i\epsilon)t} \right].$$

Assuming that the time dependence of the field amplitude and the distribution functions is weak compared to the very fast field oscillations, they can be taken out of the integral in (49). Then, using the relation

$$e^{\pm iz \cos \omega t} = \sum_{n=-\infty}^{\infty} (\pm i)^n J_n(z) e^{\pm in\omega t},$$

we obtain

$$F_{fb}(\mathbf{P}, \mathbf{p}; t) = \frac{1}{2\hbar\mu} \frac{e}{\mu} e^{-\frac{i}{\hbar}(E(\mathbf{p})-E_b)t} \sum_{k=-\infty}^{\infty} i^k J_k\left(\frac{\mathbf{v}_0 \cdot \mathbf{P}}{\hbar\omega}\right) e^{ik\omega t} \\ \times \sum_{n=-\infty}^{\infty} (-i)^n J_n\left(\frac{\mathbf{v}_0 \cdot \mathbf{P}}{\hbar\omega}\right) \int_{t_0}^t d\bar{t} e^{\frac{i}{\hbar}(E(\mathbf{p})-E_b)\bar{t}} e^{-in\omega\bar{t}} \left[e^{i(\omega-i\epsilon)\bar{t}} - e^{-i(\omega+i\epsilon)\bar{t}} \right] \\ \times \mathbf{A}_0 \cdot \mathbf{p} \varphi_b(\mathbf{p}) [F_{bb}(\mathbf{P}; t) - F_1(\mathbf{P}, \mathbf{p}; t)F_2(\mathbf{P}, \mathbf{p}; t)]. \quad (51)$$

The time integration can be carried out, and after an index shift $k = n + l$ we obtain the result

$$F_{fb}(\mathbf{P}, \mathbf{p}; t) = -\frac{e}{2\mu} \sum_{n=-\infty}^{\infty} \sum_{l=-\infty}^{\infty} i^{l-1} J_n\left(\frac{\mathbf{v}_0 \cdot \mathbf{P}}{\hbar\omega}\right) J_{n+l}\left(\frac{\mathbf{v}_0 \cdot \mathbf{P}}{\hbar\omega}\right) \\ \times \left[\frac{e^{i((l+1)\omega-i\epsilon)t}}{E_b - E(\mathbf{p}) + (n-1)\hbar\omega + i\hbar\epsilon} - \frac{e^{i((l-1)\omega-i\epsilon)t}}{E_b - E(\mathbf{p}) + (n+1)\hbar\omega + i\hbar\epsilon} \right] \\ \times \mathbf{A}_0 \cdot \mathbf{p} \varphi_b(\mathbf{p}) [F_{bb}(\mathbf{P}; t) - F_1(\mathbf{P}, \mathbf{p}; t)F_2(\mathbf{P}, \mathbf{p}; t)]. \quad (52)$$

This formula can be rewritten once more by appropriate index shifts and using the Bessel function relations

$$J_{-n}(z) = (-1)^n J_n(z); \quad J_{n-1}(z) + J_{n+1}(z) = \frac{2n}{z} J_n(z). \quad (53)$$

We arrive at

$$F_{fb}(\mathbf{P}, \mathbf{p}; t) = \frac{e}{\mu} \sum_{n=-\infty}^{\infty} \sum_{l=-\infty}^{\infty} i^l \frac{n\hbar\omega}{\mathbf{v}_0 \cdot \mathbf{p}} J_n\left(\frac{\mathbf{v}_0 \cdot \mathbf{P}}{\hbar\omega}\right) J_{n+l}\left(\frac{\mathbf{v}_0 \cdot \mathbf{p}}{\hbar\omega}\right) \\ \times \frac{e^{i(l\omega - i\epsilon)t}}{E_b - E(\mathbf{p}) + n\hbar\omega + i\hbar\epsilon} \mathbf{A}_0 \cdot \mathbf{p} \varphi_b(\mathbf{p}) [F_{bb}(\mathbf{P}; t) - F_1(\mathbf{P}, \mathbf{p}; t) F_2(\mathbf{P}, \mathbf{p}; t)]. \quad (54)$$

As can be seen from this result, the field causes interesting physical effects. First, the sum over l indicates the generation of higher harmonics of the field frequency, cf. the term $e^{i(l\omega - i\epsilon)t}$. On the other hand, the sum over n reflects the absorption or emission of multiple photons, i.e. multiphoton ionization, which finds its expression in the denominator $E_b - E(\mathbf{p}) + n\hbar\omega + i\hbar\epsilon$. Similar effects have been found, e.g., in [34, 35, 36].

3.4. Absorption and ionization

3.4.1. Electrical current and absorption coefficient Now we are able to consider the current density and other interesting physical quantities connected to \mathbf{j} like polarization, absorption and ionization coefficients. For this purpose, let us come back to the balance equation (12).

With the completeness relation for the eigenstates $|\mathbf{P}\alpha\rangle$ and the Schrödinger equation

$$1 = \sum_{\mathbf{P}\alpha} |\mathbf{P}\alpha\rangle \langle \mathbf{P}\alpha|, \quad (H_{ab}^0 - E_{\alpha\mathbf{P}}) |\mathbf{P}\alpha\rangle = -V_{ab} |\mathbf{P}\alpha\rangle,$$

we express the right hand side of the balance equation (12) for the current density in terms of the $F_{ab}^{\alpha\alpha'}(\mathbf{P}; t)$. It follows

$$\frac{d\mathbf{j}(t)}{dt} - \omega_{pl}^2 \mathbf{E} = -\frac{\mathcal{V}}{2i\hbar} \sum_{ab} \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} \int \frac{d\mathbf{P}}{(2\pi\hbar)^3} \sum_{\alpha\bar{\alpha}} \left(\frac{e_t \mathbf{P}}{M} + \frac{e_r \mathbf{p}}{\mu} \right) (E_{\mathbf{p}\mathbf{P}} - E_{\alpha\mathbf{P}}) \\ [\varphi_{\alpha}(\mathbf{p}) \varphi_{\bar{\alpha}}^*(\mathbf{p}) F_{ab}^{\alpha\bar{\alpha}}(\mathbf{P}; t) - \varphi_{\alpha}^*(\mathbf{p}) \varphi_{\bar{\alpha}}(\mathbf{p}) F_{ab}^{\bar{\alpha}\alpha}(\mathbf{P}; t)] = \sum_{ab} I_{ab}(t) \quad (55)$$

This equation shows that the different transition processes discussed in Sec. 3.1 contribute to the current in a partially ionized plasma. In the following, we consider a hydrogen plasma as a simple model case. We should notice, however, that the theory can be applied to more complicated systems, too. We consider here the electron-proton part and especially the bound-free contribution ($\alpha = n = 1, \bar{\alpha} = \mathbf{p}$ and vice versa)

$$I_{ep}(t) = \frac{\mathcal{V}}{\hbar} \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} \int \frac{d\mathbf{P}}{(2\pi\hbar)^3} \frac{e\mathbf{p}}{\mu} (E(\mathbf{p}) - E_b) \varphi_b(\mathbf{p}) \text{Im} F_{fb}(\mathbf{P}; t). \quad (56)$$

Then we introduce the expression (54) into (56) and ignore the weak time dependence of the distribution functions and the field amplitude. Finally we integrate Eq. (55) over the time. Then follows

$$\mathbf{j}(t) = \mathbf{j}_0(t) + \sum_{l=-\infty}^{\infty} \mathbf{j}_l(\omega) e^{-il\omega t} \quad (57)$$

which is clearly the Fourier expansion of the current in terms of all harmonics of the field frequency ω . The Fourier coefficients $\mathbf{j}_l(\omega)$ are given by

$$\mathbf{j}_l(\omega) = -\frac{\mathcal{V}}{2} \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} \int \frac{d\mathbf{P}}{(2\pi\hbar)^3} \frac{e}{\mu} \frac{E(\mathbf{p}) - E_b}{l\hbar\omega} \mathbf{p} \varphi_b(\mathbf{p}) \frac{e}{\mu} \mathbf{A}_0 \cdot \mathbf{p} \varphi_b(\mathbf{p})$$

$$\begin{aligned}
& \times [F_{bb}(\mathbf{P}; t) - F_e(\mathbf{P}, \mathbf{p}; t)F_p(\mathbf{P}, \mathbf{p}; t)] \sum_{n=-\infty}^{\infty} i^l \frac{n\hbar\omega}{\mathbf{v}_0 \cdot \mathbf{p}} J_n\left(\frac{\mathbf{v}_0 \cdot \mathbf{p}}{\hbar\omega}\right) J_{n-l}\left(\frac{\mathbf{v}_0 \cdot \mathbf{p}}{\hbar\omega}\right) \\
& \times \left[\frac{(-1)^l}{E_b - E(\mathbf{p}) + n\hbar\omega + i\hbar\epsilon} + \frac{1}{E_b - E(\mathbf{p}) - n\hbar\omega - i\hbar\epsilon} \right]
\end{aligned} \tag{58}$$

Furthermore, $\mathbf{j}_0(t) = \omega_{pl}^2 \int_{-\infty}^t dt' \mathbf{E}(t')$ is the current of the collisionless plasma.

As already mentioned above, the current determines several physical quantities, e.g. the polarization. Here, we will consider the energy transfer $\mathbf{j} \cdot \mathbf{E}$ between the field and the plasma. For the field \mathbf{E} we assume the time dependence (47). We consider the dissipation of the energy averaged over one cycle of oscillation. Then we obtain after a simple calculation

$$\langle \mathbf{j} \cdot \mathbf{E} \rangle = \frac{1}{T} \int_{t-T}^t dt' \mathbf{j}(t) \cdot \mathbf{E}(t) = \mathbf{E}_0 \cdot \text{Re } \mathbf{j}_1(\omega). \tag{59}$$

Now we introduce the absorption coefficient for the bound-free transition by

$$\alpha_{bf}(\omega) = \frac{1}{c \varepsilon_0} \frac{\langle \mathbf{j} \cdot \mathbf{E} \rangle}{\langle \mathbf{E}^2 \rangle} \tag{60}$$

With the help of the Dirac identity, the substitution $n \rightarrow -n$ in the second term and the Bessel function relations (53), after some algebra follows

$$\begin{aligned}
\alpha_{bf}(\omega) &= \frac{\mathcal{V}}{c \varepsilon_0 \hbar} \frac{2\pi}{E_0^2} \sum_{n=-\infty}^{\infty} (n\hbar\omega)^3 \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} \int \frac{d\mathbf{P}}{(2\pi\hbar)^3} \frac{E(\mathbf{p}) - E_b}{n\hbar\omega} |\varphi_b(\mathbf{p})|^2 \\
&\times J_n^2\left(\frac{\mathbf{v}_0 \cdot \mathbf{p}}{\hbar\omega}\right) \delta(E(\mathbf{p}) - E_b - n\hbar\omega) [F_{bb}(\mathbf{P}; t) - F_e(\mathbf{P}, \mathbf{p}; t)F_p(\mathbf{P}, \mathbf{p}; t)]
\end{aligned} \tag{61}$$

In order to determine the absorption coefficient we need the diagonal elements of the density matrix which have to be calculated from the corresponding equations (40) and (41). In order to obtain first results, we replace these quantities by thermodynamic equilibrium functions. This assumption represents a significant simplification, in particular, the influence of the field on the bound states is neglected. An improved determination of F_{bb} from Eq. (40) leads to a Stark shift of levels quadratic in the $\mathbf{A} \cdot \mathbf{p}$ interaction as was discussed also by Milonni and Ackerhalt [40].

The quantity $\alpha_{bf}(\omega)$ describes the process of the absorption of n photons in the process of ionization of atoms. It is shown in Fig. 7 in dependence on the laser frequency. An interesting feature is the occurrence of additional maxima for photon energies $\hbar\omega$ smaller than the ionization energy. They are caused by two- and more-photon processes.

If we restrict the sum in formula (61) to $n = 1$, our approach reproduces the relations of linear response theory. A further significant simplification follows in the case of thermodynamic equilibrium, i.e., for Maxwellian distributions for F_{bb} and $F_{e/p}$.

3.4.2. Rate coefficients The absorption of photons is, under the condition $E(\mathbf{p}) - E_b = n\hbar\omega$, connected with a bound-free transition, i.e. the ionization of the atoms. The process of ionization, of course, changes the plasma composition. From the macroscopic point of view, the time dependence of the densities of the plasma particles is given by rate equations. For the change of the atomic density $n_A(t)$ we have

$$\frac{\partial}{\partial t} n_A(t) = -\alpha(t) n_A(t) + \beta(t) n_e(t) n_p(t). \tag{62}$$

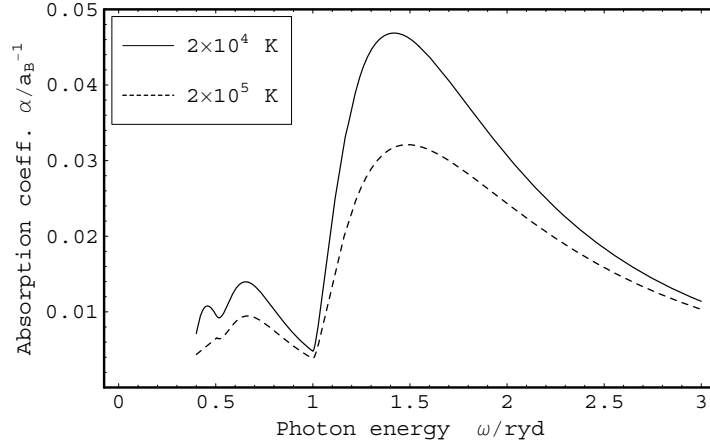


Figure 7. Absorption coefficient α_{bf} vs. laser frequency for two different temperatures.

Here, $n_e(t)$ and $n_p(t)$ are the densities of free particles (electrons and protons), and the coefficients $\alpha(t)$ and $\beta(t)$ are the rate coefficients.

From the microscopic point of view, the rate equation (62) follows from the equation for the occupation of the atoms F_{bb} (40),

$$\frac{\partial}{\partial t} F_{bb}(\mathbf{P}; t) = \frac{2}{\hbar} \sin \omega t \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} \mathbf{v}_0 \cdot \mathbf{p} \varphi_b(\mathbf{p}) \text{Im} F_{fb}(\mathbf{P}, \mathbf{p}; t). \quad (63)$$

Inserting the solution (54) and using the Dirac identity, we arrive at

$$\begin{aligned} \frac{\partial}{\partial t} F_{bb}(\mathbf{P}; t) = & -\frac{2}{\hbar} \sin \omega t \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} \mathbf{v}_0 \cdot \mathbf{p} |\varphi_b(\mathbf{p})|^2 \sum_{n=-\infty}^{\infty} \sum_{l=-\infty}^{\infty} n\hbar\omega J_n\left(\frac{\mathbf{v}_0 \cdot \mathbf{P}}{\hbar\omega}\right) J_{n+l}\left(\frac{\mathbf{v}_0 \cdot \mathbf{p}}{\hbar\omega}\right) \\ & \times \left\{ \sin\left[l\left(\omega t + \frac{\pi}{2}\right)\right] \frac{P}{E_b - E(\mathbf{p}) + n\hbar\omega} - \pi \cos\left[l\left(\omega t + \frac{\pi}{2}\right)\right] \delta(E(\mathbf{p}) - E_b - n\hbar\omega) \right\} \\ & \times [F_{bb}(\mathbf{P}; t) - F_e(\mathbf{P}, \mathbf{p}; t) F_p(\mathbf{P}, \mathbf{p}; t)]. \end{aligned} \quad (64)$$

In the following, we assume Maxwellian distributions for the free particles, i.e.

$$F_{e/p}(\mathbf{P}, \mathbf{p}; t) = \frac{n_{e/p}(t) \Lambda_{e/p}^3}{2} e^{-\frac{1}{2m_{e/p} k_B T} \frac{m_{e/p}}{M} \mathbf{P} \pm \mathbf{p}}^2,$$

with the thermal wavelength $\Lambda_{e/p} = \sqrt{\frac{2\pi\hbar^2}{m_{e/p} k_B T}}$. Integrating Eq. (64) over \mathbf{P} , we obtain the rate equation (62) for the density of the atoms $n_A(t)$. Obviously, the ionization coefficient $\alpha(t)$ is then given by the microscopic expression

$$\begin{aligned} \alpha(t) = & \frac{2}{\hbar} \sin \omega t \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} \mathbf{v}_0 \cdot \mathbf{p} |\varphi_b(\mathbf{p})|^2 \sum_{n=-\infty}^{\infty} \sum_{l=-\infty}^{\infty} n\hbar\omega J_n\left(\frac{\mathbf{v}_0 \cdot \mathbf{P}}{\hbar\omega}\right) J_{n+l}\left(\frac{\mathbf{v}_0 \cdot \mathbf{p}}{\hbar\omega}\right) \\ & \times \left\{ \sin\left[l\left(\omega t + \frac{\pi}{2}\right)\right] \frac{P}{E_b - E(\mathbf{p}) + n\hbar\omega} - \pi \cos\left[l\left(\omega t + \frac{\pi}{2}\right)\right] \delta(E(\mathbf{p}) - E_b - n\hbar\omega) \right\}, \end{aligned} \quad (65)$$

and the recombination coefficient follows from

$$\begin{aligned} \beta(t) = & \frac{\Lambda_r^3}{2\hbar} \sin \omega t \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} \mathbf{v}_0 \cdot \mathbf{p} |\varphi_b(\mathbf{p})|^2 \sum_{n=-\infty}^{\infty} \sum_{l=-\infty}^{\infty} n\hbar\omega J_n\left(\frac{\mathbf{v}_0 \cdot \mathbf{p}}{\hbar\omega}\right) J_{n+l}\left(\frac{\mathbf{v}_0 \cdot \mathbf{p}}{\hbar\omega}\right) \\ & \times \left\{ \sin\left[l\left(\omega t + \frac{\pi}{2}\right)\right] \frac{P}{E_b - E(\mathbf{p}) + n\hbar\omega} - \pi \cos\left[l\left(\omega t + \frac{\pi}{2}\right)\right] \delta(E(\mathbf{p}) - E_b - n\hbar\omega) \right\} \\ & \times e^{-\frac{p^2}{2\mu k_B T}} \end{aligned} \quad (66)$$

with $\Lambda_r = \sqrt{\frac{2\pi\hbar^2}{\mu k_B T}}$. It is useful to average again over one period of the field oscillation. Using the relations $\frac{1}{T} \int_0^T dt \sin \omega t \cos l\omega t = 0$, $\frac{1}{T} \int_0^T dt \sin \omega t \sin l\omega t = \frac{l}{2} \delta_{|l|1}$, we get the averaged coefficients

$$\bar{\alpha} = \frac{2\pi}{\hbar} \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} |\varphi_b(\mathbf{p})|^2 \sum_{n=-\infty}^{\infty} (n\hbar\omega)^2 J_n^2\left(\frac{\mathbf{v}_0 \cdot \mathbf{p}}{\hbar\omega}\right) \delta(E(\mathbf{p}) - E_b - n\hbar\omega), \quad (67)$$

$$\bar{\beta} = \frac{\pi\Lambda_r^3}{2\hbar} \int \frac{d\mathbf{p}}{(2\pi\hbar)^3} |\varphi_b(\mathbf{p})|^2 \sum_{n=-\infty}^{\infty} (n\hbar\omega)^2 J_n^2\left(\frac{\mathbf{v}_0 \cdot \mathbf{p}}{\hbar\omega}\right) \delta(E(\mathbf{p}) - E_b - n\hbar\omega) e^{-\frac{p^2}{2\mu k_B T}}. \quad (68)$$

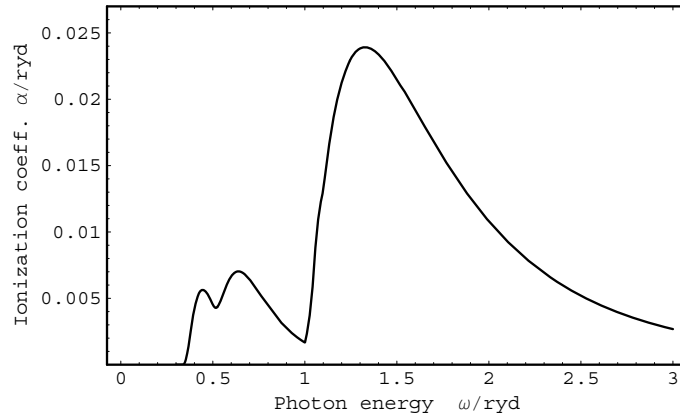


Figure 8. Ionization coefficient α vs. laser frequency.

This strongly simplified expression for the ionization coefficient is essentially equivalent to results from atomic physics [37, 38, 39, 40]. However, in contrast to the results of Faisal and Reiss, no effect of the A^2 term occurs. Effects of higher order in A following from the $\mathbf{p} \cdot \mathbf{A}$ coupling are contained, in principle, in the basic equations (40) to (43) as mentioned already above in the discussion of the absorption coefficient.

In Fig. 8 we show the ionization coefficient α in dependence on the laser energy. Similar to the absorption coefficient we again observe a finite probability for the ionization for photon energies smaller than the binding energy due to multi-photon processes. However, we want to remark that this multi-photon interpretation of the ionization process is limited by the Keldysh frequency [37]

$$\omega_K = \frac{eE}{\sqrt{2mI}}, \quad (69)$$

where I is the ionization energy. Only for frequencies larger than ω_K the multi-photon interpretation is meaningful. For frequencies smaller than ω_K the ionization has to be interpreted by the usual tunnel effect.

Acknowledgments

This work was supported by the Deutsche Forschungsgemeinschaft (SFB 198).

References

- [1] D. Kremp, M. Schlanges, and W.-D. Kraeft, *Quantum Statistics of Nonideal Plasmas* (Springer, Berlin 2005)
- [2] D. Kremp, Th. Bornath, M. Bonitz, and M. Schlanges, Phys. Rev. E **60**, 4725 (1999)
- [3] M. Bonitz, *Quantum Kinetic Theory* (Teubner, Stuttgart/Leipzig 1998)
- [4] Th. Bornath, M. Schlanges, P. Hilse, and D. Kremp, J. Phys. A: Math. Gen. **36** 5941-5948 (2003).
- [5] Th. Bornath, M. Schlanges, P. Hilse, and D. Kremp, in: K. Morawetz (Ed.) Nonequilibrium physics at short time scales: formation of correlations (Springer-Verlag, Berlin 2004), pp. 153-172.
- [6] V.P. Silin, Zh. Eksp. Teor. Fiz. **47**, 2254 (1964), [JETP **20**, 1510 (1965)]
- [7] Yu.L. Klimontovich, *Kinetic Theory of Nonideal Gases and Nonideal Plasmas* (russ.), (Nauka, Moscow 1975) [Engl. transl.: Pergamon Press, Oxford 1982]
- [8] M. Bonitz, Th. Bornath, D. Kremp, M. Schlanges, and W.D. Kraeft, Contrib. Plasma Phys. **39**, 329 (1999)
- [9] D. Semkat, D. Kremp, and M. Bonitz, Phys. Rev. E **59**, 1557 (1999)
- [10] D. Semkat, D. Kremp, and M. Bonitz, Contrib. Plasma Phys. **42**, 31 (2002)
- [11] D. Semkat, M. Bonitz, and D. Kremp, Contrib. Plasma Phys. **43**, 321 (2003)
- [12] D. Kremp, D. Semkat, and M. Bonitz, J. Phys.: Conference Series **11**, 1 (2005)
- [13] Th. Bornath, D. Kremp, and M. Schlanges, Phys. Rev. E **60**, 6382 (1999)
- [14] C. Oberman, A. Ron, and J. Dawson, Phys. Fluids **5**, 1514 (1962); J.M. Dawson and C. Oberman: Phys. Fluids **6**, 394 (1963)
- [15] C.D. Decker, W.B. Mori, J.M. Dawson, and T. Katsouleas, Phys. Plasmas **1**, 4043 (1994)
- [16] P. Mulser, F. Cornolti, E. Bèsuelle, and R. Schneider, Phys. Rev. E **63**, 63 (2000)
- [17] H.-J. Kull and L. Plagne, Phys. of Plasmas **8**, 5244 (2001)
- [18] Th. Bornath, M. Schlanges, P. Hilse, and D. Kremp, Phys. Rev. E **64**, 26414 (2001)
- [19] G. Hazak, N. Metzler, M. Klapisch, and J. Gardner, Phys. of Plasmas **9**, 345 (2002)
- [20] H. Reinholz, R. Redmer, G. Röpke, and A. Wierling, Phys. Rev. E **62**, 5648 (2000)
- [21] Th. Bornath, D. Kremp, and M. Schlanges, unpublished (2004)
- [22] G. Bekefi, *Radiation processes in plasmas* (Wiley, New York 1966)
- [23] M. Schlanges, P. Hilse, Th. Bornath, and D. Kremp, in: M. Bonitz and D. Semkat (Eds.) Progress in Nonequilibrium Green's Functions (World Scientific, Singapore 2003), pp. 50-65.
- [24] H. Haberland, M. Bonitz, and D. Kremp, Phys. Rev. E **64**, 026405 (2001)
- [25] M. Schlanges, Th. Bornath, D. Kremp, and P. Hilse, Contrib. Plasma Phys. **43**, 360 (2003)
- [26] Th. Bornath, D. Kremp, P. Hilse, and M. Schlanges, J. Phys.: Conference Series **11**, 180 (2005)
- [27] R. Cauble and W. Rozmus, Phys. Fluids **28**, 3387 (1985)
- [28] S. Pfalzner and P. Gibbon, Phys. Rev. E **57**, 4698 (1998)
- [29] M. Bonitz, D. Semkat, D. Kremp, and V. S. Filinov, in *Progress in Nonequilibrium Green's functions II*, M. Bonitz and D. Semkat (Eds.), World Scientific Publ., Singapore 2003, pp. 445-463
- [30] D. Semkat, D. Kremp, and M. Bonitz, J. Phys.: Conference Series **11**, 25 (2005)
- [31] Yu. L. Klimontovich, *Kinetic theory of electromagnetic processes*, Springer, Berlin, Heidelberg, New York (1983)
- [32] R. Schepe, D. Tamme, and K. Henneberger, Contrib. Plasma Phys. **39**, 29 (1999)
- [33] Schepe R, Tamme D and Henneberger K 2000 in *Progress in Nonequilibrium Green's Functions* ed M Bonitz (Singapore: World Scientific Publ.) p 170
- [34] Kremp D, Bornath Th, Bonitz M and Schlanges M 1999 *Phys. Rev. E* **60** 4725
- [35] Bonitz M, Bornath Th, Kremp D, Schlanges M and Kraeft W D 1999 *Contrib. Plasma Phys.* **39** 329
- [36] Bornath Th, Schlanges M, Hilse P and Kremp D 2001 *Phys. Rev. E* **64** 26414
- [37] L. V. Keldysh, Zh. Eksp. Teor. Fiz. **47**, 1945 (1964) [Sov. Phys.-JETP **20**, 1307 (1965)]
- [38] F. Faisal, J. Phys. B: Atom. Molec. Phys. **6**, L 89 (1973)
- [39] H. R. Reiss, Phys. Rev. A **22**, 1786 (1980); Phys. Rev. A **42**, 1476 (1989)
- [40] P. W. Milonni and J. R. Ackerhalt, Phys. Rev. A **39**, 1139 (1989)