## Momentum distribution function and short-range correlations in dense quantum plasmas – ab initio quantum Monte Carlo results

Michael Bonitz, Alexey Filinov, and Tobias Dornheim<sup>†</sup>

Institute of Theoretical Physics and Astrophysics, Kiel University <sup>†</sup>Center for Advanced Systems Understanding

Plasma Theory and Simulation, PTS-2022

pdf at www.theo-physik.uni-kiel.de/bonitz/research.html





The momentum distribution function (thermodynamic equilibrium)

#### Classical plasma

- ideal plasma: Maxwell distribution
- interacting plasma: Maxwell distribution
  - $\Rightarrow$  exponential decay for large momenta

#### Quantum plasma

ideal plasma: Fermi/Bose function

#### $\Rightarrow$ exponential decay for large momenta

### The momentum distribution function (thermodynamic equilibrium)

#### Classical plasma

- ideal plasma: Maxwell distribution
- interacting plasma: Maxwell distribution
  - $\Rightarrow$  exponential decay for large momenta

#### Quantum plasma

- ideal plasma: Fermi/Bose function
  - $\Rightarrow$  exponential decay for large momenta

#### What about nonideal Quantum plasmas?

- ► slower non-exponential decay, ~ p<sup>-8</sup>, predicted<sup>1</sup>
- relevant for threshold processes: dramatic increase of reaction rates, fusion rates could be possible
- important for electrons under warm dense matter (WDM) conditions or ions in dense stars
- First *ab initio* Quantum Monte Carlo results for WDM available:
   K. Hunger *et al.*, Phys. Rev. E **103**, 053204 (2021), arXiv:2101.00842
   T. Dornheim *et al.*, Phys. Rev. B **103**, 205142 (2021)

<sup>&</sup>lt;sup>1</sup>Daniel, Vosko (1960); Galitskii, Migdal (1967)

### Warm Dense Matter: Occurences and Applications [Andrew NG (2000): "missing link between CM, plasmas"]

#### Astrophysics:

- Giant planet interiors (e.g. Jupiter)
- Brown dwarfs
- Earth interior, Meteor Impacts
- Recently discovered planets



Source: Sci-News.com [Img4]

### Warm Dense Matter: Occurences and Applications [Andrew NG (2000): "missing link between CM, plasmas"]

#### Astrophysics:

- Giant planet interiors (e.g. Jupiter)
- Brown dwarfs
- Earth interior, Meteor Impacts
- Recently discovered planets

#### Laboratory Experiments, shock compression:

- Lasers, FELs, Z-pinch, ion beams
- Properties of matter under extreme conditions, e.g. Kritcher et al., Nature 2020
- major driver: Inertial confinement fusion

#### Potential abundance of clean energy!



Source: en.wikipedia.org [Img5] and arstechnica.com [Img6]

### Warm Dense Matter: Occurences and Applications [Andrew NG (2000): "missing link between CM, plasmas"]

#### Astrophysics:

- Giant planet interiors (e.g. Jupiter)
- Brown dwarfs
- Earth interior, Meteor Impacts
- Recently discovered planets

#### Laboratory Experiments, shock compression:

- Lasers, FELs, Z-pinch, ion beams
- Properties of matter under extreme conditions, e.g. Kritcher et al., Nature 2020
- major driver: Inertial confinement fusion

#### Potential abundance of clean energy!

US: NIF, Omega (Rochester), LCLS (Stanford): Fundamental research into WDM properties:  $\rightarrow$  Equation of state,  $S(q, \omega)$ , conductivity etc.

#### National Ignition Facility (Livermore, California)



area: 70000*m*<sup>2</sup> cost: ~1 billion Dollar <u>Source:</u> C. Stolz, *Phil. Trans. R. Soc. A* **370**, 4115 (2012) [Img7]

### **Progress in Inertial Confinement Fusion**

Lasor Energy Info Hohliraum

continuous optimization of target design, pulse shape etc.

Record shot on August 8 2021: 1.92 MJ UV laser energy source: https://lasers.llnl.gov/news/hybridexperiments-drive-nif-toward-ignition



The "Hybrid-E experiment" on Aug. 8 achieved a hot-spot absorbed energy of about 65 kJ—about 20 kJ from the implosion, and the rest from "self-heating" from the fusion reactions (self-sustained burn). 1.35 MJ fusion energy yield, corresponds to 70% of ignition threshold (NAS criterion). Zylstra *et al.*, Nature (2022)

### Facilities for WDM experiments in Europe and Asia:

### Free electron lasers:

- FLASH (DESY, Hamburg)
- European X-ray Free-Electron Laser, Hamburg – Schenefeld
- HIBEF Beamline and consortium. 2021 first successful experiments
- Fermi (Triest, Italy)
- SACLA (Riken, Japan)



source: photon-science.desy.de

### Facilities for WDM experiments in Europe and Asia:

### Free electron lasers:

- FLASH (DESY, Hamburg)
- European X-ray Free-Electron Laser, Hamburg – Schenefeld
- HIBEF Beamline and consortium. 2021 first successful experiments
- Fermi (Triest, Italy)
- SACLA (Riken, Japan)



source: photon-science.desy.de

### Heavy ion beams:

- Facility for Antiproton and Ion Research, Darmstadt
- Construction started in 2017
- Heavy ion beams: Isochoric heating up to ~ 10<sup>6</sup>K





source: dw.com

source: inspirehep.net

### Facilities for WDM experiments in Europe and Asia:

### Free electron lasers:

- FLASH (DESY, Hamburg)
- European X-ray Free-Electron Laser, Hamburg – Schenefeld
- HIBEF Beamline and consortium. 2021 first successful experiments



Fermi (Triest\_Italy)

► SACL

## Warm dense matter: indeed a HOT topic

### Heavy ion beams:

- Facility for Antiproton and Ion Research, Darmstadt
- Construction started in 2017
- Heavy ion beams: Isochoric heating up to  $\sim 10^6 K$





source: dw.com

source: inspirehep.net

#### Extreme and exotic state of matter:

- $\rightarrow$  High temperature:  $T\sim 10^3-10^8~K$
- ightarrow Extreme density:  $n \sim 10^{21} 10^{27} \ cm^{-3}$



#### Extreme and exotic state of matter:

- $\rightarrow$  High temperature:  $T\sim 10^3-10^8~K$
- ightarrow Extreme density:  $n \sim 10^{21} 10^{27} \ cm^{-3}$

Characteristic parameters:



#### Extreme and exotic state of matter:

- $\rightarrow$  High temperature:  $T\sim 10^3-10^8~K$
- ightarrow Extreme density:  $n \sim 10^{21} 10^{27} \ cm^{-3}$

#### Characteristic parameters:

• Density (coupling) parameter  $r_s = \overline{r}/a_B \sim 1$ 



#### Extreme and exotic state of matter:

- $\rightarrow$  High temperature:  $T \sim 10^3 10^8 K$
- $\rightarrow$  Extreme density:  $n \sim 10^{21} 10^{27} cm^{-3}$

#### Characteristic parameters:

- Density (coupling) parameter  $r_s = \overline{r}/a_B \sim 1$ Degeneracy temperature  $\theta = k_B T/E_F \sim 1$
- $\Theta < 1$ : quantum plasma.

  - $\Theta > 1$ : classical plasma



0910 T / K

#### Extreme and exotic state of matter:

- $\rightarrow$  High temperature:  $T\sim 10^3-10^8~{\it K}$
- ightarrow Extreme density:  $n \sim 10^{21} 10^{27} \ cm^{-3}$

#### Characteristic parameters:

- Density (coupling) parameter r<sub>s</sub> =  $\bar{r}/a_B \sim 1$  Degeneracy temperature  $\theta = k_B T/E_F \sim 1$
- Degeneracy temperature  $\theta = k_{\rm B}T/E_{\rm F} \sim \Theta < 1$ : quantum plasma,
  - $\Theta > 1$ : classical plasma

Classical coupling parameter 
$$\Gamma = e^2 / r_s k_B T \sim 1$$

#### Source: T. Dornheim, S. Groth, and M. Bonitz, Phys. Reports 744, 1-86 (2018) 10 r<sub>s</sub>=10 Classical 9 8 7 6 5 4 Ideal 3 0=0 Fermi gas Metals 2 18 20 22 24 26 28 30 log<sub>10</sub> n / cm<sup>-3</sup>

#### Extreme and exotic state of matter:

- $\rightarrow$  High temperature:  $T\sim 10^3-10^8~K$
- ightarrow Extreme density:  $n \sim 10^{21} 10^{27} \ cm^{-3}$

#### Characteristic parameters:

- Density (coupling) parameter  $r_s = \overline{r}/a_B \sim 1$
- Degeneracy temperature  $\theta = k_{\rm B}T/E_{\rm F} \sim 1$  $\Theta < 1$ : quantum plasma,
  - $\Theta > 1$ : classical plasma
- Classical coupling parameter  $\Gamma = e^2/r_s k_B T \sim 1$

#### Nontrivial interplay of many effects:

Coulomb coupling (non-ideality)



#### Extreme and exotic state of matter:

- $\rightarrow$  High temperature:  $T\sim 10^3-10^8~{\it K}$
- ightarrow Extreme density:  $n\sim 10^{21}-10^{27}~cm^{-3}$

#### Characteristic parameters:

- Density (coupling) parameter  $r_s = \overline{r}/a_B \sim 1$
- Degeneracy temperature  $\theta = k_{\rm B}T/E_{\rm F} \sim 1$  $\Theta < 1$ : quantum plasma,  $\Theta > 1$ : classical plasma
- Classical coupling parameter  $\Gamma = e^2/r_s k_B T \sim 1$

#### Nontrivial interplay of many effects:

- Coulomb coupling (non-ideality)
- Fermionic exchange (anti-symmetry)



Source: cidehom.com [Img2]

#### Extreme and exotic state of matter:

- $\rightarrow$  High temperature:  $T\sim 10^3-10^8~K$
- ightarrow Extreme density:  $n \sim 10^{21} 10^{27} \ cm^{-3}$

#### Characteristic parameters:

- Density (coupling) parameter  $r_s = \overline{r}/a_B \sim 1$
- Degeneracy temperature  $\theta = k_{\rm B} T/E_{\rm F} \sim 1$  $\Theta < 1$ : quantum plasma,  $\Theta > 1$ : classical plasma
- Classical coupling parameter  $\Gamma = e^2/r_s k_B T \sim 1$

#### Nontrivial interplay of many effects:

- Coulomb coupling (non-ideality)
- Fermionic exchange (anti-symmetry)
- Thermal excitations (statistical description)



Source: en.wikipedia.org [Img3]

### Importance of the uniform electron gas (UEG)

Model system of Coulomb interacting quantum electrons in a uniform positive background

### Ground state (T = 0):

- Simple model for conduction electrons in metals
- Exchange-correlation (XC) energy:

 $e_{xc}(r_s) = e_{tot}(r_s) - e_0(r_s)$ 

- $\rightarrow$  Input for density functional theory (DFT) simulations (in LDA and GGA)
- $\rightarrow$  Parametrization<sup>1</sup> of  $e_{xc}(r_s)$  from ground state quantum Monte Carlo data<sup>2</sup>
- ightarrow this made DFT-MD the basis of modern atomic, molecular physics, chemistry, material science

<sup>&</sup>lt;sup>1</sup> J.P. Perdew and A. Zunger, PRB **23**, 5048 (1981) <sup>2</sup> D.M. Ceperley and B. Alder, PRL **45**, 566 (1980) <sup>3</sup> N.D. Mermin, Phys. Rev **137**, A1441 (1965) <sup>4</sup> A.Y. Potekhin and G. Chabrier, *A&A 550*, *A43* (2013)

### Importance of the uniform electron gas (UEG)

Model system of Coulomb interacting quantum electrons in a uniform positive background

### Ground state (T = 0):

- Simple model for conduction electrons in metals
- Exchange-correlation (XC) energy:

 $e_{\rm xc}(r_s) = e_{\rm tot}(r_s) - e_0(r_s)$ 

- $\rightarrow$  Input for density functional theory (DFT) simulations (in LDA and GGA)
- $\rightarrow$  Parametrization<sup>1</sup> of  $e_{xc}(r_s)$  from ground state quantum Monte Carlo data<sup>2</sup>
- ightarrow this made DFT-MD the basis of modern atomic, molecular physics, chemistry, material science

### Warm dense matter ( $T \sim T_F$ ):

Thermal DFT<sup>3</sup>: minimize free energy F = E − TS → Requires parametrization of XC free energy of UEG:

$$f_{xc}(r_s, \theta) = f_{tot}(r_s, \theta) - f_0(r_s, \theta)$$

- $f_{xc}(r_s, \theta)$  direct input for Equation of state (EOS) models of astrophysical objects<sup>4</sup>
- f<sub>xc</sub>(r<sub>s</sub>, θ) contains complete thermodynamic information of UEG

<sup>&</sup>lt;sup>1</sup> J.P. Perdew and A. Zunger, PRB **23**, 5048 (1981) <sup>2</sup> D.M. Ceperley and B. Alder, PRL **45**, 566 (1980) <sup>3</sup> N.D. Mermin, Phys. Rev **137**, A1441 (1965) <sup>4</sup> A.Y. Potekhin and G. Chabrier, *A&A 550*, *A43* (2013)

### Path Integral Monte Carlo (PIMC): Fermions

#### Fermionic antisymmetry:

$$Z = rac{1}{N!} \sum_{\sigma \in S_N} \mathrm{sgn}(\sigma) \int \mathrm{d}\mathbf{R} \, \left\langle \mathbf{R} 
ight| e^{-eta \hat{H}} \left| \hat{\pi}_\sigma \mathbf{R} 
ight
angle$$

 $\Rightarrow$  We must include **permutation-cycles**!



PIMC configuration of N = 3 particles,  $W(\mathbf{X}) < 0$ 

Taken from: T. Dornheim, S. Groth, A. Filinov, and M. Bonitz, *J. Chem. Phys.* **151**, 014108 (2019)

### Path Integral Monte Carlo (PIMC): Fermions

#### Fermionic antisymmetry:

$$Z = rac{1}{N!} \sum_{\sigma \in S_N} \operatorname{sgn}(\sigma) \int \mathrm{d}\mathbf{R} \, \left\langle \mathbf{R} \right| e^{-eta \hat{H}} \left| \hat{\pi}_\sigma \mathbf{R} 
ight
angle$$

- $\Rightarrow$  We must include **permutation-cycles**!
- Randomly generate all possible paths X using the Metropolis algorithm



Snapshot of PIMC simulation of UEG with N = 19,  $r_s = 2$ ,  $\theta = 0.5$  (fluctuating probability density)

Taken from: T. Dornheim, S. Groth, A. Filinov, and M. Bonitz, *J. Chem. Phys.* **151**, 014108 (2019)

### Path Integral Monte Carlo (PIMC): Fermions

#### Fermionic antisymmetry:

$$Z = rac{1}{N!} \sum_{\sigma \in S_N} \mathrm{sgn}(\sigma) \int \mathrm{d}\mathbf{R} \, \left\langle \mathbf{R} \right| e^{-eta \hat{H}} \left| \hat{\pi}_\sigma \mathbf{R} 
ight
angle$$

- $\Rightarrow$  We must include **permutation-cycles**!
- Randomly generate all possible paths X using the Metropolis algorithm
- Sign changes due to particle exchange lead to vanishing signal-to-noise ratio
  - ⇒ Fermion Sign Problem (unsolved!)



Exponential decrease of the average sign S with system size N and quantum degeneracy  $\theta^{-1}$ 

Taken from: T. Dornheim, Phys. Rev. E 100, 023307 (2019)

Standard PIMC in warm dense regime severely hampered by *fermion sign problem*:



- <sup>2</sup> V. Filinov et al., Phys. Rev. E 91, 033108 (2015)
- <sup>3</sup> T. Schoof *et al.*, Phys. Rev. Lett **115**, 130402 (2015)
  - <sup>5</sup> T. Dornheim *et al.*, New J. Phys. **17**, 073017 (2015)
- <sup>4</sup> T. Schoof *et al.*, Contrib. Plasma Phys. **51**, 687 (2011)
  - <sup>6</sup> T. Dornheim *et al.*, J. Chem. Phys. **143**, 204101 (2015)

- Standard PIMC in warm dense regime severely hampered by *fermion sign problem*:
  - First results<sup>1</sup> by E. Brown, D. Ceperley et al. (2013) based on fixed node approximation (RPIMC)
  - Induces systematic errors of unknown magnitude
  - RPIMC limited to r<sub>s</sub> ≥ 1
  - Fermionic **PIMC**: Filinov et al.<sup>2</sup> limited to  $r_s \gtrsim 1$



- <sup>2</sup> V. Filinov *et al.*, Phys. Rev. E **91**, 033108 (2015)
- <sup>3</sup> T. Schoof *et al.*, Phys. Rev. Lett **115**, 130402 (2015)
  - <sup>5</sup> T. Dornheim et al., New J. Phys. 17, 073017 (2015)
- <sup>4</sup> T. Schoof *et al.*, Contrib. Plasma Phys. **51**, 687 (2011)
  - <sup>6</sup> T. Dornheim *et al.*, J. Chem. Phys. **143**, 204101 (2015)

- Standard PIMC in warm dense regime severely hampered by *fermion sign problem*:
  - First results<sup>1</sup> by E. Brown, D. Ceperley et al. (2013) based on fixed node approximation (RPIMC)
  - Induces systematic errors of unknown magnitude
  - RPIMC limited to r<sub>s</sub> ≥ 1
  - Fermionic **PIMC**: Filinov et al.<sup>2</sup> limited to  $r_s \gtrsim 1$

### Our approach:

Avoid fermion sign problem by combining two exact and complementary QMC methods:



- <sup>2</sup> V. Filinov et al., Phys. Rev. E 91, 033108 (2015)
- <sup>3</sup> T. Schoof *et al.*, Phys. Rev. Lett **115**, 130402 (2015)
  - <sup>5</sup> T. Dornheim *et al.*, New J. Phys. **17**, 073017 (2015)
- <sup>4</sup> T. Schoof *et al.*, Contrib. Plasma Phys. **51**, 687 (2011)
  - <sup>6</sup> T. Dornheim et al., J. Chem. Phys. 143, 204101 (2015)

- Standard PIMC in warm dense regime severely hampered by *fermion sign problem*:
  - First results<sup>1</sup> by E. Brown, D. Ceperley et al. (2013) based on fixed node approximation (RPIMC)
  - Induces systematic errors of unknown magnitude
  - RPIMC limited to r<sub>s</sub> ≥ 1
  - Fermionic **PIMC**: Filinov et al.<sup>2</sup> limited to  $r_s \gtrsim 1$

### Our approach:

Avoid fermion sign problem by combining two exact and complementary QMC methods:

- 1. Configuration PIMC (CPIMC)<sup>3,4</sup>
  - $\rightarrow$  Excels at high density  $\mathit{r_s} \lesssim$  1 and strong degeneracy



- <sup>2</sup> V. Filinov et al., Phys. Rev. E 91, 033108 (2015)
- <sup>3</sup> T. Schoof *et al.*, Phys. Rev. Lett **115**, 130402 (2015)
  - <sup>5</sup> T. Dornheim et al., New J. Phys. 17, 073017 (2015)
- <sup>4</sup> T. Schoof *et al.*, Contrib. Plasma Phys. **51**, 687 (2011)
  - <sup>6</sup> T. Dornheim et al., J. Chem. Phys. 143, 204101 (2015)

- Standard PIMC in warm dense regime severely hampered by *fermion sign problem*:
  - First results<sup>1</sup> by E. Brown, D. Ceperley et al. (2013) based on fixed node approximation (RPIMC)
  - Induces systematic errors of unknown magnitude
  - **RPIMC** limited to  $r_s \gtrsim 1$
  - Fermionic **PIMC**: Filinov et al.<sup>2</sup> limited to  $r_s \gtrsim 1$

### Our approach:

Avoid fermion sign problem by combining two exact and complementary QMC methods:

- 1. Configuration PIMC (CPIMC)<sup>3,4</sup>
  - ightarrow Excels at high density  $r_s \lesssim$  1 and strong degeneracy
- 2. Permutation blocking PIMC (PB-PIMC)<sup>5,6</sup>
  - $\rightarrow$  Extends standard PIMC towards stronger degeneracy



- <sup>2</sup> V. Filinov et al., Phys. Rev. E 91, 033108 (2015)
- <sup>3</sup> T. Schoof *et al.*, Phys. Rev. Lett **115**, 130402 (2015)
  - <sup>5</sup> T. Dornheim *et al.*, New J. Phys. **17**, 073017 (2015)
- <sup>4</sup> T. Schoof *et al.*, Contrib. Plasma Phys. **51**, 687 (2011)
  - <sup>6</sup> T. Dornheim et al., J. Chem. Phys. 143, 204101 (2015)

- Standard PIMC in warm dense regime severely hampered by *fermion sign problem*:
  - First results<sup>1</sup> by E. Brown, D. Ceperley et al. (2013) based on fixed node approximation (RPIMC)
  - Induces systematic errors of unknown magnitude
  - **RPIMC** limited to  $r_s \gtrsim 1$
  - Fermionic **PIMC**: Filinov et al.<sup>2</sup> limited to  $r_s \gtrsim 1$

### Our approach:

Avoid fermion sign problem by combining two exact and complementary QMC methods:

- 1. Configuration PIMC (CPIMC)<sup>3,4</sup>
  - ightarrow Excels at high density  $r_s \lesssim$  1 and strong degeneracy
- 2. Permutation blocking PIMC (PB-PIMC)<sup>5,6</sup>
  - $\rightarrow$  Extends standard PIMC towards stronger degeneracy



### Ab initio simulations over broad range of parameters possible

- <sup>1</sup> E.W. Brown *et al.*, PRL **110**, 146405 (2013)
- <sup>2</sup> V. Filinov et al., Phys. Rev. E 91, 033108 (2015)
- <sup>3</sup> T. Schoof *et al.*, Phys. Rev. Lett **115**, 130402 (2015)
  - <sup>5</sup> T. Dornheim et al., New J. Phys. 17, 073017 (2015)
- <sup>4</sup> T. Schoof *et al.*, Contrib. Plasma Phys. **51**, 687 (2011)
- <sup>6</sup> T. Dornheim et al., J. Chem. Phys. 143, 204101 (2015)

1. Exact exchange-correlation energy  $E_{xc} = E - E_0$  ( $E_0$ : ideal energy) (N = 33 spin-polarized electrons,  $\theta \ge 0.5$ ,  $\forall r_s$ )



<sup>1</sup>S. Groth *et al.*, Phys. Rev. B **93**, 085102 (2016) <sup>3</sup>F.D. Malone *et al.*, Phys. Rev. Lett. **117**, 115701 (2016)

<sup>2</sup>T. Dornheim *et al.*, Phys. Rev. B **93**, 205134 (2016)

<sup>4</sup>T. Dornheim et al., Phys. Rev. Lett. (2016)

- 1. Exact exchange-correlation energy  $E_{xc} = E E_0$  ( $E_0$ : ideal energy) (N = 33 spin-polarized electrons,  $\theta \ge 0.5$ ,  $\forall r_s$ )
- **RPIMC** limited to  $r_s \ge 1$
- CPIMC excels at high density



<sup>1</sup>S. Groth *et al.*, Phys. Rev. B **93**, 085102 (2016) <sup>3</sup>F.D. Malone *et al.*, Phys. Rev. Lett. **117**, 115701 (2016)

<sup>2</sup>T. Dornheim *et al.*, Phys. Rev. B **93**, 205134 (2016)

<sup>4</sup>T. Dornheim et al., Phys. Rev. Lett. (2016)

- 1. Exact exchange-correlation energy  $E_{xc} = E E_0$  ( $E_0$ : ideal energy) (N = 33 spin-polarized electrons,  $\theta \ge 0.5$ ,  $\forall r_s$ )
- **RPIMC** limited to  $r_s \ge 1$
- CPIMC excels at high density
- **PB-PIMC** applicable at  $\theta \gtrsim 0.5$



<sup>1</sup>S. Groth *et al.*, Phys. Rev. B **93**, 085102 (2016) <sup>3</sup>F.D. Malone *et al.*, Phys. Rev. Lett. **117**, 115701 (2016)

<sup>2</sup>T. Dornheim *et al.*, Phys. Rev. B **93**, 205134 (2016)

<sup>4</sup>T. Dornheim et al., Phys. Rev. Lett. (2016)

- 1. Exact exchange-correlation energy  $E_{xc} = E E_0$  ( $E_0$ : ideal energy) (N = 33 spin-polarized electrons,  $\theta \ge 0.5$ ,  $\forall r_s$ )
- **RPIMC** limited to  $r_s \ge 1$
- CPIMC excels at high density
- **PB-PIMC** applicable at  $\theta \gtrsim 0.5$

Combination<sup>1</sup> yields exact results over entire density range down to  $\theta \sim 0.5$ 

- Also applies to the unpolarized UEG<sup>2</sup>
- confirmed by independent DMQMC simulations<sup>3</sup>
- Extended to TD Limit<sup>4</sup> and to the ground state<sup>5</sup>
- Analytical parametrization of f<sub>xc</sub>(r<sub>s</sub>, θ, ξ), "GDSMFB", with error below 0.3%, Integrated in standard DFT libraries<sup>5</sup>



<sup>1</sup>S. Groth et al., Phys. Rev. B 93, 085102 (2016)

<sup>3</sup>F.D. Malone et al., Phys. Rev. Lett. 117, 115701 (2016)

<sup>2</sup>T. Dornheim *et al.*, Phys. Rev. B **93**, 205134 (2016)

<sup>4</sup>T. Dornheim et al., Phys. Rev. Lett. (2016)

### Recognition for our work...

APS John Dawson Award 2021

Kiel group:

Tim Schoof (PhD 2016), Simon Groth (PhD 2018): CPIMC, finite size corrections etc.

Tobias Dornheim (PhD 2018): PB-PIMC now at CASUS Görlitz, Extension to static and dynamic response, transport, DFT, machine learning, nonlinear density response etc.

UK and US collaborators: F. Malone, M. Foulkes and T. Sjostrom



63rd Annual Meeting of the APS Division of Plasma Physics



### 2021 PRIZES & AWARDS

#### John Dawson Award for Excellence in Plasma Physics Research

For developing Monte Carlo methods that overcome the fermion sign problem, leading to the first ab initio data for an electron gas under warm dense matter conditions.



William Foulkes Imperial College ondon







Travis Siostrom Los Alamos National Laboratory Tobias Dornheim Center for Advanced Systems Understanding



Fionn Malone



Michael Bonitz Kiel University



Tim Schoof DESY

### Ab Initio PIMC approach to equilibrium response and transport properties

### Quantities accessible in PIMC:

all thermodynamic functions from  $F(r_s, \theta)$ ; structural properties: g(r), S(q)fluctuations in response to excitation:  $\delta \hat{H}(\mathbf{q}) \longrightarrow \delta \rho(\mathbf{q})$ correlation functions: e.g.  $\langle \delta \rho(\mathbf{q}_1, \tau_1) \rho(\mathbf{q}_2, \tau_2) \rangle$  yield transport properties

### Susceptibilites from linear response theory (LRT):

 $\delta \rho(\mathbf{q}) = \chi(\mathbf{q}) \delta H(\mathbf{q}), \quad \chi:$  static density response  $\longrightarrow$  comparison for PIMC to LRT/experiment

### **Correlation and exchange effects:** encoded in "local field correction" $G(\mathbf{q}, \omega)$ straightforward connection to transport<sup>2</sup>, optics etc.: $\chi(\mathbf{q}, \omega)$ , $S(\mathbf{q}, \omega)^3$ , $\epsilon(\mathbf{q}, \omega)$ , $\sigma(\mathbf{q}, \omega)$ , plasmon dispersion<sup>4</sup>

### PIMC: susceptibilities beyond validity limits of LRT<sup>5</sup>

### Ab initio spectral properties, momentum distribution n(p)

<sup>&</sup>lt;sup>2</sup>Hamann et al., Phys. Rev. B (2020)

<sup>&</sup>lt;sup>3</sup>Dornheim et al., Phys. Rev. Lett. (2018)

<sup>&</sup>lt;sup>4</sup>Hamann et al., Contrib. Plasma Phys. (2020)

<sup>&</sup>lt;sup>5</sup>Dornheim et al., Phys. Rev. Lett. (2020)

### Momentum distribution of correlated electrons in WDM<sup>6</sup>

#### Key questions

- 1. Is the large momentum asymptotic of n(p) indeed of order  $p^{-8}$ ?
- 2. How does the asymptotic depend on density and momentum?
- 3. How do correlations and quantum effects influence the low-momentum states?

#### Earlier works

- **•** non-exponential decay,  $\sim p^{-8}$ , predicted by Daniel, Vosko (1960); Galitskii, Migdal (1967) and others
- Many ground state results: analytical and QMC: Gori-Giorgi et al., Calmels, Overhauser, Spink et al.
- Observed also in cold atoms, but there asymptotic  $\sim p^{-4}$ , e.g. Doggen, Kinnunen, (2015)
- importance for dense plasmas discussed e.g. by Starostin, Fisch and others
- recent QMC simulations: Militzer, V. Filinov et al.

<sup>&</sup>lt;sup>6</sup>K. Hunger et al., Phys. Rev. E 103, 053204 (2021)

### Momentum distribution of correlated electrons in WDM<sup>7</sup>

#### Key questions

- 1. Is the large momentum asymptotic of n(p) indeed of order  $p^{-8}$ ?
- 2. How does the asymptotic depend on density and momentum?
- 3. How do correlations and quantum effects influence the low-momentum states?

#### Earlier works

- non-exponential decay,  $\sim p^{-8}$ , predicted by Daniel, Vosko (1960); Galitskii, Migdal (1967) and others
- Many ground state results: analytical and QMC: Gori-Giorgi et al., Calmels, Overhauser, Spink et al.
- Observed also in cold atoms, but there asymptotic  $\sim \rho^{-4}$ , e.g. Doggen, Kinnunen, (2015)
- importance for dense plasmas discussed e.g. by Starostin, Fisch and others
- recent QMC simulations: Militzer, V. Filinov et al.
- Asymptotic given by on-top pair distribution, for all temperatures, via

$$\lim_{p\to\infty} n(p) = \frac{4}{9} \left(\frac{4}{9\pi}\right)^{2/3} \left(\frac{r_s}{\pi}\right)^2 \frac{p_F^8}{p^8} g^{\uparrow\downarrow}(0) \,,$$

Kimball (1975); Yasuhara, Kawazoe (1976)

#### Tasks

- Develop CPIMC and fermionic PIMC simulations for n(p) and  $g^{\uparrow\downarrow}(0)$
- Compute n(p) and  $g^{\uparrow\downarrow}(0)$  for WDM parameters, explore density and temperature dependence
- Generate accurate benchmark data for n(p) for all momenta. Input for models, reaction rates etc.

<sup>&</sup>lt;sup>7</sup>K. Hunger et al., Phys. Rev. E 103, 053204 (2021)

### CPIMC approach to the momentum distribution and the on-top PDF<sup>9</sup>

CPIMC is QMC in Fock space (second quantization)<sup>8</sup> Exact description of quantum electrons at  $r_s \lesssim 1$ 

$$\hat{\mathcal{H}} = \sum_{ij} h_{ij} \hat{a}^{\dagger}_i \hat{a}_j + rac{1}{2} \sum_{ijkl} w_{ijkl} \, \hat{a}^{\dagger}_i a^{\dagger}_j \hat{a}_l \hat{a}_k \,, \qquad \hat{
ho} = rac{1}{Z} \, e^{-eta \hat{\mathcal{H}}} \,,$$

Uniform electron gas: Use plane wave basis. Generate paths C in Fock space with weight W(C)

Estimators for single-particle and two-particle density matrix:

$$n_{i} = \langle a_{i}^{\dagger} a_{i} \rangle = -\frac{1}{\beta} \frac{\partial}{\partial h_{ii}} \ln Z = \frac{1}{Z} \sum_{C} \left( \sum_{\nu=0}^{K} n_{i}^{(\nu)} \frac{\tau_{i+1} - \tau_{i}}{\beta} \right) W(C)$$
$$d_{ijkl} = \langle a_{i}^{\dagger} a_{j}^{\dagger} a_{k} a_{l} \rangle = -\frac{1}{\beta} \frac{\partial}{\partial w_{ijkl}} \ln Z,$$
$$g^{\uparrow\downarrow}(0, C) = \frac{1}{2N_{\sigma_{1}}(C)N_{\sigma_{2}}(C)} \sum_{ijkl} \delta_{s_{i},s_{l}} \delta_{s_{j},s_{k}} (1 - \delta_{s_{i},s_{j}}) d_{ijkl}(C)$$

## <sup>8</sup>Schoof, Bonitz *et al.*, Contrib. Plasma Phys. (2011) <sup>9</sup>K. Hunger *et al.*, Phys. Rev. E **103**, 053204 (2021)

#### Illustration for 3 particles



Figure: Continuous time representation of the path integral, for N = 3. Paths *C* are classified by the number *K* of kinks, their times and involved orbitals. **Ideal Fermi gas:** one Slater determinant, corresponds to straight lines (no kinks). **Correlations:** mix of Slater determinants, leads to increase or *K*.

### Results for the momentum distribution – Overview<sup>10</sup>



Figure: Left: Temperature dependence at  $r_s = \bar{r}/a_B = 0.5$ . Full lines: CPIMC, dashed: Fermi function  $n^{id}$ . Right: Density dependence at  $\Theta = k_B T/E_F = 2.0$  Full lines: CPIMC, dashed: ground state, black:  $n^{id}$ .

<sup>&</sup>lt;sup>10</sup>K. Hunger *et al.*, Phys. Rev. E **103**, 053204 (2021) ground state results: Gori-Giorgi *et al.* (2001)

### Results for the momentum distribution – large-k asymptotic<sup>11</sup>



Tail of n(k),

 $r_s = 0.5$  and  $\Theta = 2$ 

pink: asymptotic  $n^{\infty}(k)$ , using CPIMC result for  $g^{\uparrow\downarrow}(0)$ 

 $\delta^\infty$  : relative deviation from asymptotic

$$\delta^{\infty}(k) = \frac{n(k)}{n^{\infty}(k)} - 1.$$

Ordering of curves determined by  $g^{\uparrow\downarrow}(0; \Theta)$ 

<sup>&</sup>lt;sup>11</sup>K. Hunger et al., Phys. Rev. E 103, 053204 (2021), ground state results: Gori-Giorgi et al. (2001)

### Results for the momentum distribution – low-k states<sup>12</sup>



Figure: Top: Difference correlated (CPIMC) minus ideal distribution,<br/>Bottom: Difference of kinetic energy densities. Total kinetic energy: area under curve<br/>Left: Temperature dependence at  $r_s = 0.5$ .Right: Density dependence at  $\Theta = 2$ .

Explanation: negative energy shift of low-momentum states:  $E(k) = \frac{k^2}{2m} + \Sigma_F(k) + \dots$ 

<sup>&</sup>lt;sup>12</sup>K. Hunger et al., Phys. Rev. E 103, 053204 (2021)

### Interaction-induced lowering of the kinetic energy<sup>13</sup>



Exchange-correlation contribution to kinetic energy,  $K_{\rm xc}$ ,

Black line:  $K_{xc} = 0$ , symbols: CPIMC data points blue line: Militzer *et al.* (2002) heat map (and pluses):

$$K_{\rm xc} = -f_{\rm xc} - \theta \frac{\partial f_{\rm xc}}{\partial \theta}\Big|_{r_{\rm s}} - r_{\rm s} \frac{\partial f_{\rm xc}}{\partial r_{\rm s}}\Big|_{\theta}$$

with  $f_{xc}$ : GDSMFB parametrization red circles:  $n(0) > n^{id}(0)$ green crosses:  $n(0) < n^{id}(0)$ 

<sup>13</sup>predicted by Militzer et al. (2002), and Kraeft et al. (2002), present results from: K. Hunger et al., Phys. Rev. E **103**, 053204 (2021)

### Interaction-induced lowering of the kinetic energy (contd.)<sup>14</sup>



<sup>14</sup> predicted by Militzer *et al.* (2002), and Kraeft *et al.* (2002), from: T. Dornheim *et al.*, Phys. Rev. B **103**, 205142 (2021)

## Top: Occupation of lowest momentum orbital

red, green symbols: PIMC data blue symbols: CPIMC data points

# Bottom: Exchange-correlation contribution to kinetic energy, $K_{\rm xc}$ ,

CPIMC data: minimum at  $r_s \approx 0.4^a$ 

<sup>a</sup>K. Hunger et al., Phys. Rev. E 103, 053204 (2021)

### On-top pair distribution function $g^{\uparrow\downarrow}(0)$



Figure: Left: Temperature dependence for three densities (CPIMC data where available, curves shifted vertically). Right: Density dependence at Θ = 1. Red: CPIMC, blue circles: fermionic propagator PIMC (A. Filinov), green and black line: ground state data of Spink *et al.* (2013) and Calmels *et al.* (1998) ESA: "effective static approximation" by Dornheim *et al.* (2020)

Minimum due to competition between exchange and Coulomb correlations.

### Particle number in the tail: temperature and density dependence



large-momentum asymptotic:

$$\begin{split} n(k) &\to s(r_s, \Theta) \cdot \left(\frac{k_F}{k}\right)^{-8} \\ &\sim r_s^2 \cdot g^{\uparrow\downarrow}(0, r_s, \Theta) \left(\frac{k_F}{k}\right)^{-8} \end{split}$$

,

s depends non-monotonically on  $\Theta$  and  $r_{s}$ 

black line: T = 0green symbols:  $\Theta = 2$ orange symbols:  $\Theta = 4$ 

minimum around  $\Theta \sim 0.65$  maximum around  $4 \lesssim r_s \lesssim 5$ 

### Momentum distribution for jellium at strong coupling...electron liquid<sup>15</sup>

- Fermionic PIMC simulations in grand canonical ensemble
- Extension of n(k) results to  $2 \le r_s \le 50$  and  $\Theta \ge 0.75$
- Good agreement with RPIMC data of Militzer, Pollock, and Ceperley (HEDP 2019)
- Data limited to moderate k-values



linear vs. log scale

<sup>&</sup>lt;sup>15</sup>Dornheim et al., Phys. Rev. B 103, (2021)

### Extension to hydrogen plasma

Single-bound electron problem, T = 0

ground state wave function:  $\psi_{100}(r) = rac{1}{\sqrt{\pi a_B^3}} \, e^{-r/a_B} \,, \qquad {
m p}$ 

probability density: 
$$g_{100}(r) = r^2 \frac{e^{-2r/a_B}}{\pi a_B^3}$$

momentum probability density, pr: radial momentum:16

$$|\langle \mathbf{p}|100
angle|^2 = ilde{
ho}_{100}(
ho_r) = rac{8}{\pi^2} rac{
ho_{
m Ryd}^{-3}}{\left(1+rac{
ho_r^2}{
ho_{
m Ryd}^2}
ight)^4} \,, \quad \mathcal{E}_{
m Ryd} = rac{m_r}{2} 
ho_{
m Ryd}^2$$

~

normalization:  $1 = 4\pi \int_{0}^{\infty} dp_r p_r^2 \tilde{\rho}_{100}(p_r)$ , maximum of  $p_r^2 \tilde{\rho}(p_r)$ :  $p_r^0 = p_{\text{Ryd}} / \sqrt{3}$ 

asymptotic: 
$$\lim_{p_r \to \infty} \tilde{\rho}(p_r) \sim \frac{8}{\pi^2 \rho_{\rm Ryd}^3} \left( \frac{p_{\rm Ryd}}{p_r} \right)^8$$

<sup>&</sup>lt;sup>16</sup>Podolski, Pauling, Phys. Rev. (1929)

Momentum distribution: hydrogen plasma versus jellium,  $r_s = 6$ 



Figure: Grand canonical Fermionic propagator PIMC, A. Filinov

### Pair distribution: hydrogen plasma, $r_s = 6$



Figure: Grand canonical Fermionic propagator PIMC, A. Filinov

### Electron-ion pair distribution: hydrogen plasma, $r_s = 6$



Figure: Grand canonical Fermionic propagator PIMC, A. Filinov

### Summary and outlook

### momentum distribution of quantum electrons in WDM:

- tail crucial for rates of threshold processes (e.g. fusion)
- CPIMC: unprecedented accuracy. Benchmarks, input in analytical models
- First *ab initio* results<sup>*a*</sup> for n(k) and  $g^{\uparrow\downarrow}$ :
  - based on combination of CPIMC and direct fermionic PIMC, thereby avoiding the fermion sign problem, extending previous thermodynamic results<sup>b</sup>
  - $p^{-8}$ -asymptotic quantified via on-top PDF
  - accurate FPIMC results for hydrogen

### Outlook:

- *ab initio* spectral function, energy dispersion  $n(k) = \int d\omega a(k, \omega) f^{EQ}(\omega)$
- further improvement of CPIMC, extension to two-component plasmas
  - <sup>a</sup>K. Hunger *et al.*, Phys. Rev. E **103**, 053204 (2021)
  - <sup>b</sup>T. Dornheim et al., Phys. Reports (2018)

