

# Towards *ab initio* simulation of quantum plasmas and warm dense matter

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[www.theo-physik.uni-kiel.de/bonitz](http://www.theo-physik.uni-kiel.de/bonitz)

**DFG**

**DAAD**

**HLRN**



Kiel: sailing capital

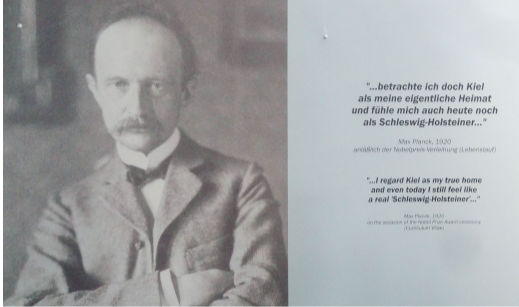


Kiel: capital of Schleswig-Holstein, sailing and science city



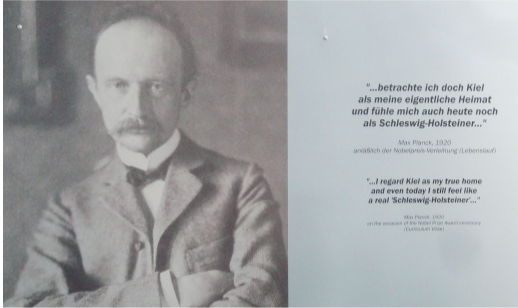
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## ► 114 years later: Chair Statistical Physics:

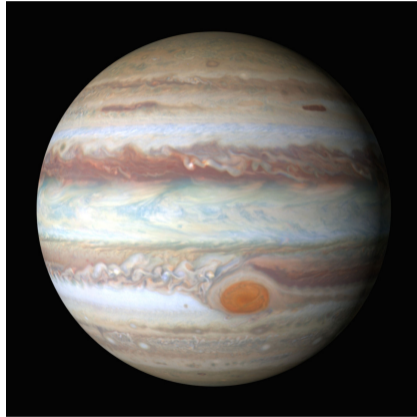
- Classical and quantum plasmas
- strongly correlated solids
- Quantum kinetic theory, nonequilibrium Green functions
- *Ab initio* simulations (MD, quantum Monte Carlo, DFT)
- Quantum hydrodynamics



# Warm Dense Matter: Occurrences and Applications

- ▶ **Astrophysics:**

- ▶ Giant planet interiors (e.g. Jupiter)
- ▶ Brown dwarfs
- ▶ Meteor Impacts



[Source: Sci-News.com \[Img4\]](#)

# Warm Dense Matter: Occurrences and Applications

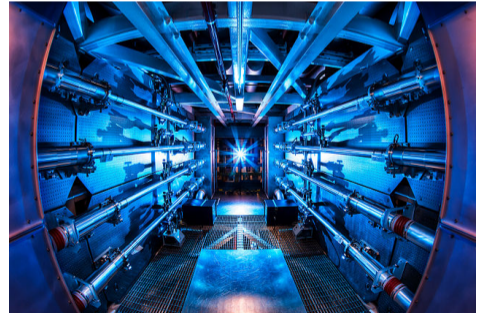
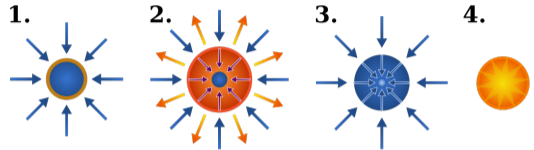
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- ▶ Meteor Impacts

## ▶ Experiments:

- ▶ Inertial confinement fusion

**Potential abundance of clean energy!**



Source: [en.wikipedia.org](https://en.wikipedia.org) [Img5] and [arstechnica.com](https://arstechnica.com) [Img6]

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**Potential abundance of clean energy!**

**WDM conditions routinely realized at  
large research facilities  
→ Equation of state,  $S(\mathbf{q}, \omega)$ , etc.**

## National Ignition Facility (Livermore, California)



area:  $70000m^2$

cost:  $\sim 1$  billion Dollar

Source: C. Stolz, *Phil. Trans. R. Soc. A* **370**, 4115 (2012) [Img7]



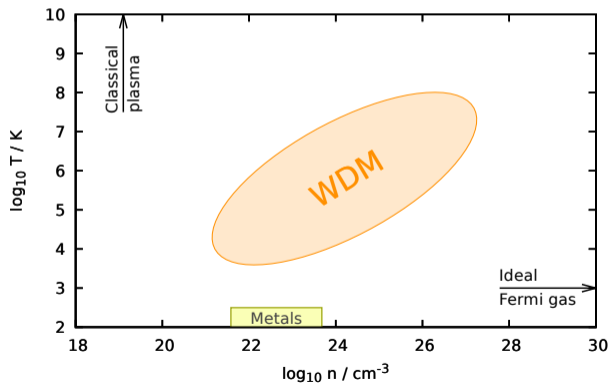
## Warm Dense Matter and quantum plasmas: relevant parameters

► **Extreme and exotic state of matter:**

→ High temperature:  $T \sim 10^4 - 10^8 \text{ K}$

→ Solid state density:  $n \sim 10^{21} - 10^{27} \text{ cm}^{-3}$

Source: [TD](#), S. Groth, and M. Bonitz,  
*Phys. Reports* **744**, 1-86 (2018)



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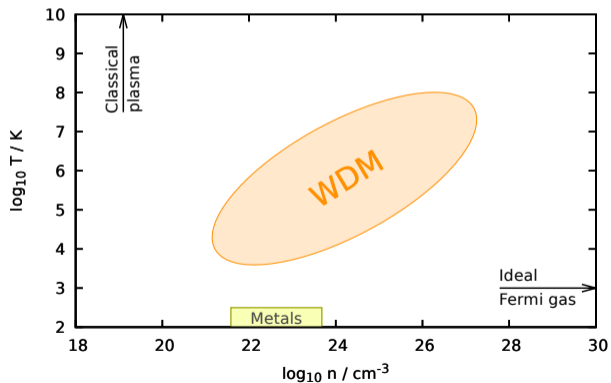
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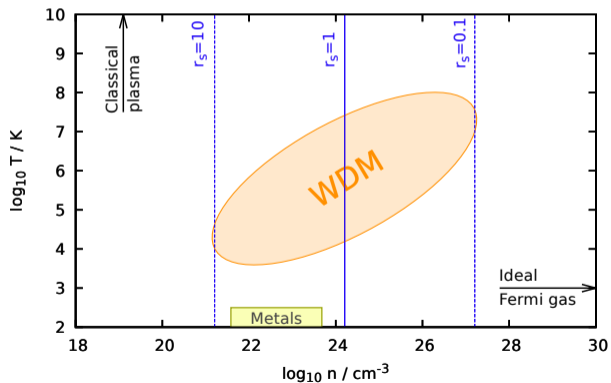
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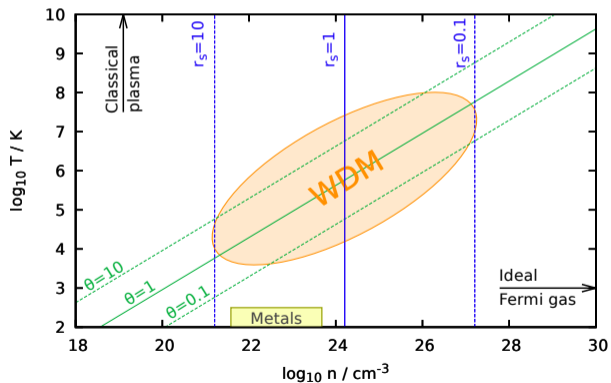
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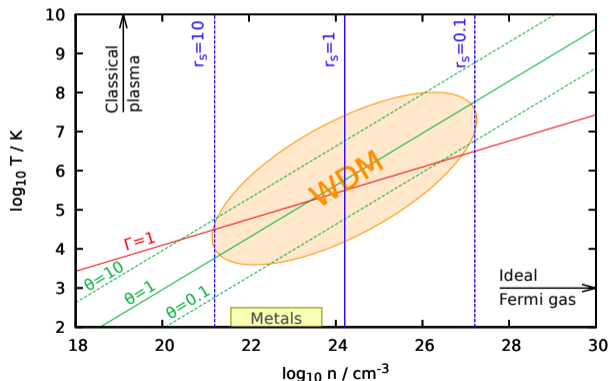
▶ Degeneracy temperature  $\theta = k_B T/E_F \sim 1$

▶  $\Theta > 1$ : quantum plasma,

$\Theta < 1$ : classical plasma

Classical coupling parameter  $\Gamma = e^2/r_s k_B T \sim 1$

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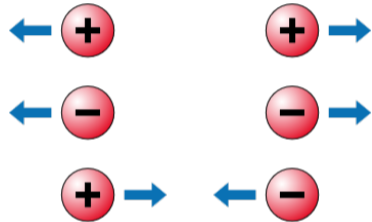
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## ▶ Nontrivial interplay of many effects:

▶ Coulomb coupling (non-ideality)



[Source: bin-br.at \[Img1\]](#)

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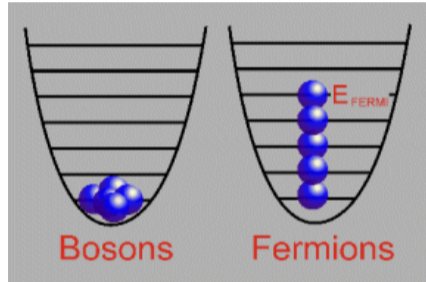
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▶ Fermionic exchange (anti-symmetry)



Source: [cidehom.com](http://cidehom.com) [Img2]

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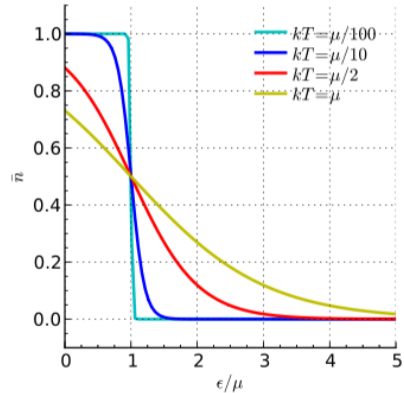
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## ▶ Nontrivial interplay of many effects:

▶ Coulomb coupling (non-ideality)

▶ Fermionic exchange (anti-symmetry)

▶ Thermal excitations (statistical description)





# Theoretical problems and simulation methods for WDM

## 1. WDM in equilibrium

- ▶ multiple components free electrons, ions, atoms, and (highly excited) solid
- ▶ ⇒ accessible only by DFT+MD

### Problems of DFT:

- ▶ assumes electrons in ground state
- ▶ poor electronic correlations, binding energies, band gaps etc.
- ▶ ⇒ accurate static input: QMC

## 2. WDM out of equilibrium

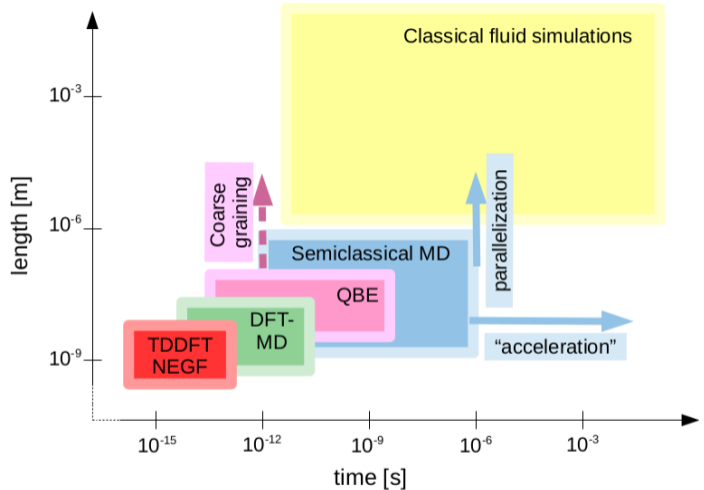


Figure: modified from M. Bonitz *et al.*, *Front. Chem. Science Engin.* (2019)  
QBE: quantum Boltzmann equation, NEGF: Nonequilibrium Green functions

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- ▶ time-dependent DFT (same problems as DFT), quantum kinetic theory, Nonequilibrium Green functions

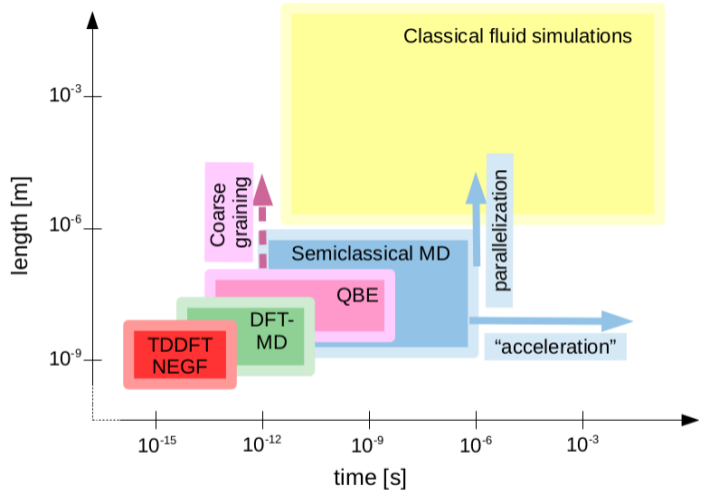


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- ▶ time-dependent DFT (same problems as DFT), quantum kinetic theory, Nonequilibrium Green functions
- ▶ complementary strengths and limitations need combination of methods

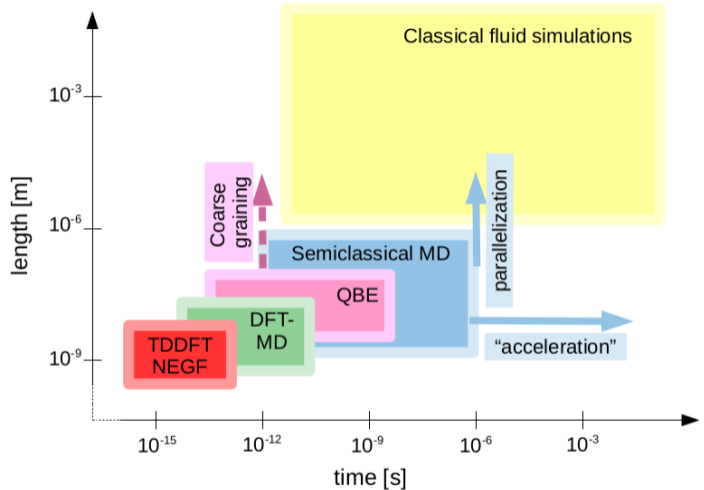


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# Importance of the uniform electron gas (UEG)

Model system of Coulomb interacting quantum electrons in a uniform positive background

## Ground state ( $T = 0$ ):

- ▶ Simple model for conduction electrons in metals
- ▶ **Exchange-correlation (XC) energy:**

$$e_{xc}(r_s) = e_{\text{tot}}(r_s) - e_0(r_s)$$

→ **Input for density functional theory (DFT) simulations (in LDA and GGA)**

→ Parametrization<sup>1</sup> of  $e_{xc}(r_s)$  from ground state quantum Monte Carlo data<sup>2</sup>

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## Warm dense matter ( $T \sim T_F$ ):

- ▶ **Thermal DFT<sup>3</sup>:** minimize free energy  $F = E - TS$
- **Requires parametrization of XC free energy of UEG:**

$$f_{xc}(r_s, \theta) = f_{\text{tot}}(r_s, \theta) - f_0(r_s, \theta)$$

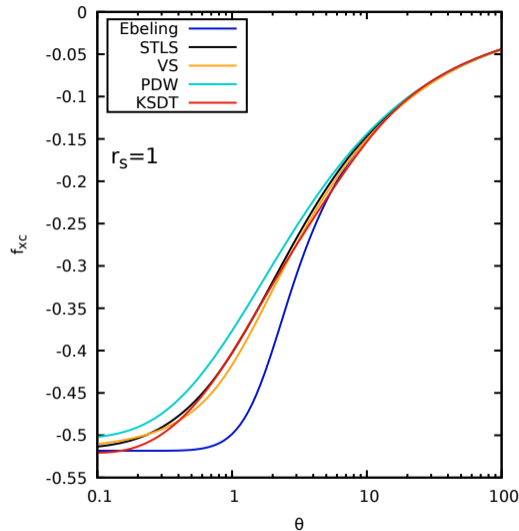
- ▶  $f_{xc}(r_s, \theta)$  direct input for **EOS models** of astrophysical objects<sup>4</sup>
- ▶  $f_{xc}(r_s, \theta)$  contains **complete thermodynamic information** of UEG

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## Many parametrizations for $f_{xc}$ based on different approximate approaches:

- ▶ Semi-analytical approaches by **Ebeling**<sup>1</sup>
- ▶ Dielectric methods, e.g. Singwi-Tosi-Land-Sjölander<sup>2</sup> (**STLS**) and Vashista-Singwi<sup>3</sup> (**VS**)
- ▶ Quantum-classical mappings, e.g. Perrot and Dharma-wardana<sup>4</sup> (**PDW**)
- ▶ **Most recent:** Fit by Karasiev<sup>5</sup> *et al.* (**KSDT**) to Restricted Path Integral Monte Carlo (**RPIMC**) data<sup>6</sup>



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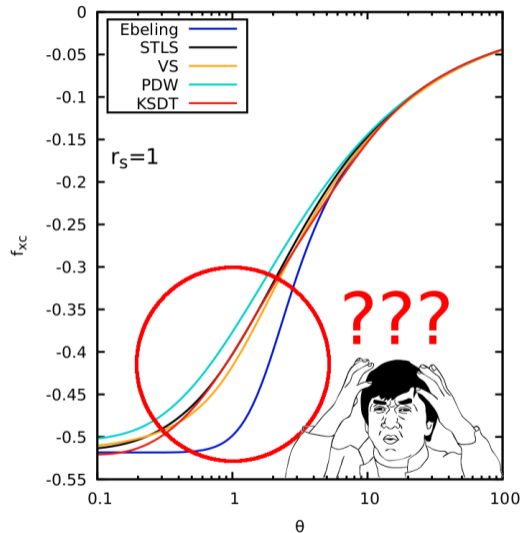
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**Accuracy of existing parametrizations for  $f_{xc}(r_s, \theta)$  unclear**



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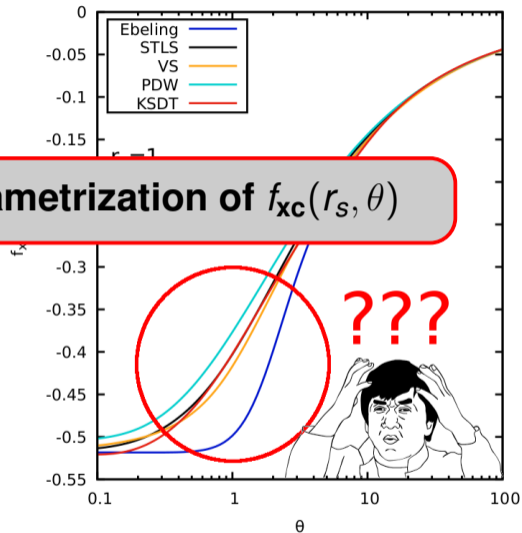
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**Goal: obtain *ab initio* parametrization of  $f_{xc}(r_s, \theta)$**

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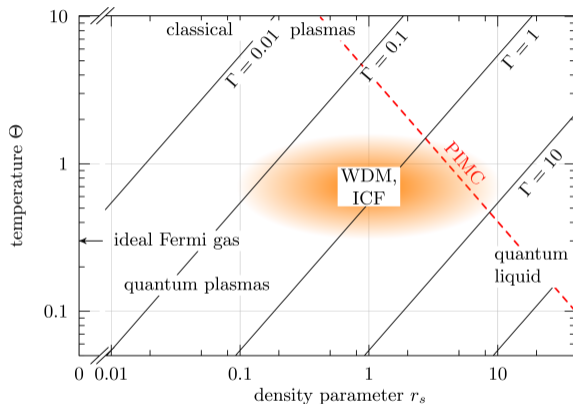
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# Path integral Monte Carlo (PIMC) simulation of the warm dense UEG

- ▶ Standard PIMC in warm dense regime severely hampered by **fermion sign problem**:



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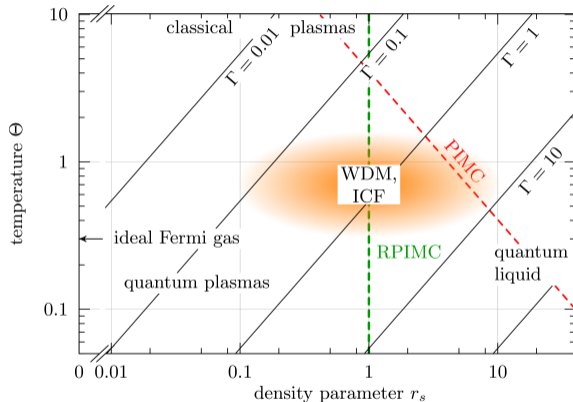
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  - ▶ Induces **systematic errors** of unknown magnitude
  - ▶ **RPIMC** limited to  $r_s \gtrsim 1$



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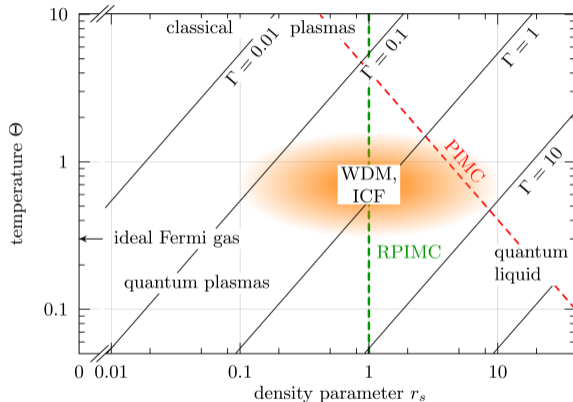
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## Our approach:

**Avoid fermion sign problem by combining two exact and complementary QMC methods:**



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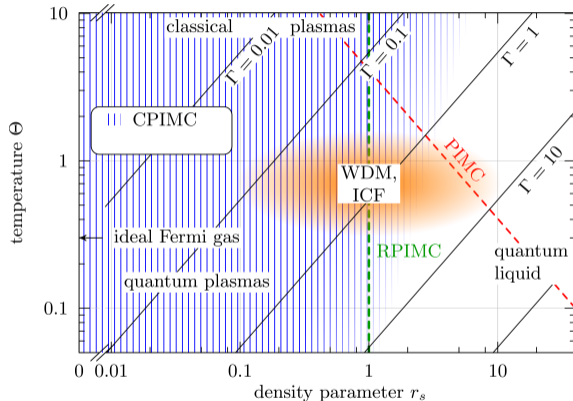
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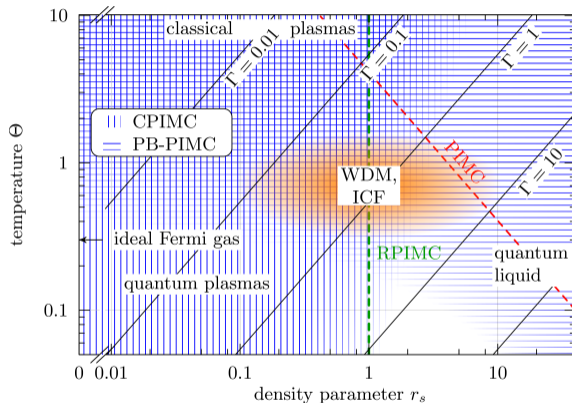
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2. **Permutation blocking PIMC (PB-PIMC)**<sup>4,5</sup>
  - Extends standard PIMC towards stronger degeneracy



<sup>1</sup> E.W. Brown *et al.*, PRL **110**, 146405 (2013)

<sup>2</sup> T. Schoof *et al.*, Phys. Rev. Lett **115**, 130402 (2015)

<sup>3</sup> T. Schoof *et al.*, Contrib. Plasma Phys. **51**, 687 (2011)

<sup>4</sup> T. Dornheim *et al.*, New J. Phys. **17**, 073017 (2015)

<sup>5</sup> T. Dornheim *et al.*, J. Chem. Phys. **143**, 204101 (2015)

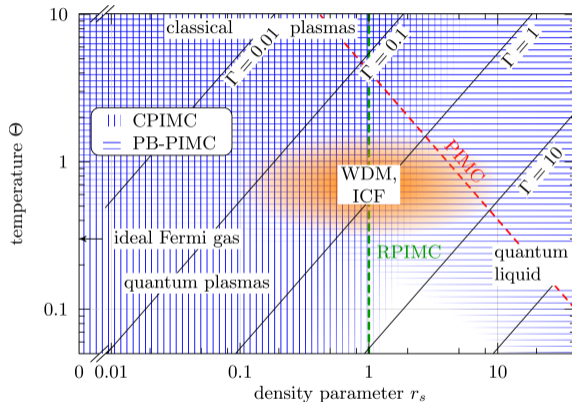
# Path integral Monte Carlo (PIMC) simulation of the warm dense UEG

- ▶ Standard PIMC in warm dense regime severely hampered by **fermion sign problem**:
  - ▶ First results<sup>1</sup> by E. Brown, D. Ceperley *et al.* (2013) based on **fixed node approximation (RPIMC)**
  - ▶ Induces **systematic errors** of unknown magnitude
  - ▶ **RPIMC** limited to  $r_s \gtrsim 1$

## Our approach:

**Avoid fermion sign problem by combining two exact and complementary QMC methods:**

1. **Configuration PIMC (CPIMC)**<sup>2,3</sup>
  - Excels at high density  $r_s \lesssim 1$  and strong degeneracy
2. **Permutation blocking PIMC (PB-PIMC)**<sup>4,5</sup>
  - Extends standard PIMC towards stronger degeneracy



***Ab initio* simulations over broad range of parameters**

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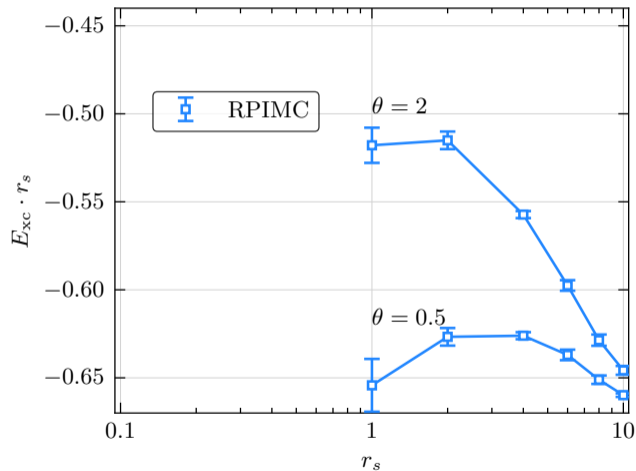
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Exact exchange-correlation energy  $E_{xc} = E - E_0$  ( $E_0$ : ideal energy)  
( $N = 33$  spin-polarized electrons,  $\theta \geq 0.5, \forall r_s$ )

► **RPIMC** limited to  $r_s \geq 1$



<sup>1</sup>S. Groth *et al.*, Phys. Rev. B **93**, 085102 (2016)

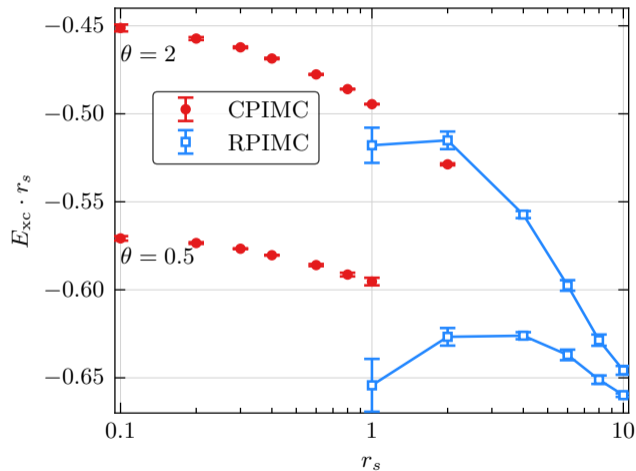
<sup>2</sup>T. Dornheim *et al.*, Phys. Rev. B **93**, 205134 (2016)

<sup>3</sup>F.D. Malone *et al.*, Phys. Rev. Lett. **117**, 115701 (2016)

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- ▶ **RPIMC** limited to  $r_s \geq 1$
- ▶ **CPIMC** excels at high density



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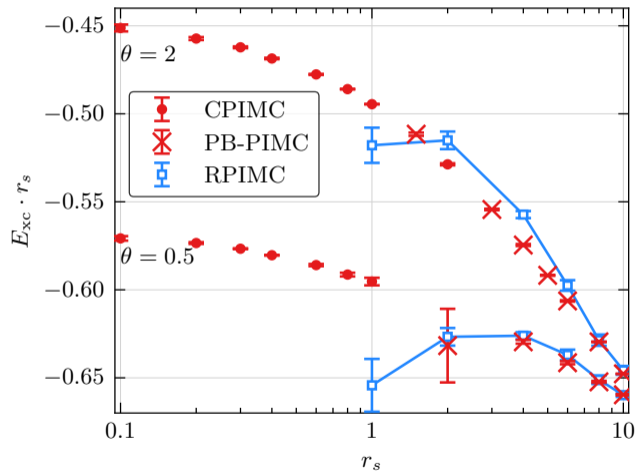
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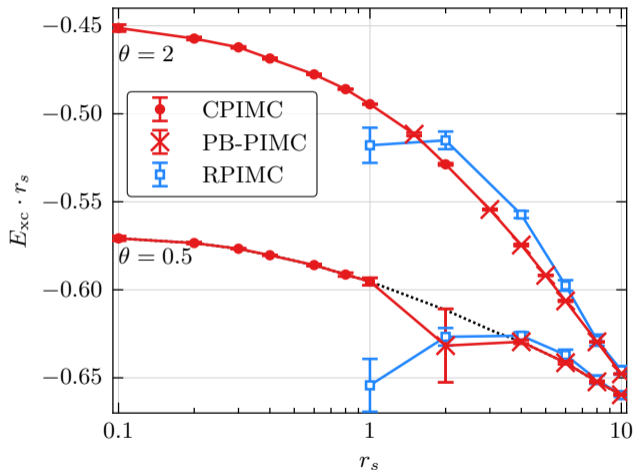
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- ▶ **RPIMC** limited to  $r_s \geq 1$
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- ▶ **PB-PIMC** applicable at  $\theta \gtrsim 0.5$

**Combination<sup>1</sup> yields exact results over entire density range down to  $\theta \sim 0.5$**

- ▶ Also applies to the **unpolarized UEG<sup>2</sup>**
- ▶ confirmed by independent **DMQMC** simulations<sup>3</sup>



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<sup>2</sup>T. Dornheim *et al.*, Phys. Rev. B **93**, 205134 (2016)

<sup>3</sup>F.D. Malone *et al.*, Phys. Rev. Lett. **117**, 115701 (2016)

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The first complete *ab initio*-based thermodynamic data for the UEG  
at WDM conditions:  $0 \leq r_s \leq 20$ ,  $0 \leq \Theta \leq 10$ ,  $0 \leq \xi \leq 1$ , error below 0.3%

1. Accurate results for  $f_{xc}$  for finite  $N$  for paramagnetic and ferromagnetic cases<sup>1,2</sup>

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<sup>3</sup>T. Dornheim *et al.*, PRL **117**, 156403 (2016)   <sup>4</sup>S. Groth *et al.*, PRL **119**, 135001 (2017)   <sup>6</sup>T. Dornheim, S. Groth, and M. Bonitz, *Physics Reports* **744**, 1-86 (2018)

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3. Connect our data to ground state data, cover entire temperature range<sup>4</sup>

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 $f_{xc}$ -functional implemented in [Libxc](#) (LDA\_XC\_GDSMFB)

## Comments and outlook:

- ▶ first unbiased tests of many earlier models and fits<sup>5,6</sup>: STLS, VS, Ichimaru, Dharma wardana, Ebeling, Green functions (Kraeft, Vorberger, Rehr...) etc.: benchmark data allow for model improvement
- ▶ earlier fit (KSDT) by Karasiev *et al.*: tested and subsequently corrected, *now good agreement* with GDSMFB

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- ▶ derivatives  $\partial f_{xc}/\partial n$  and  $\partial f_{xc}/\partial T$  potentially inaccurate;  $\Rightarrow$  *fits to separate ab initio data needed*

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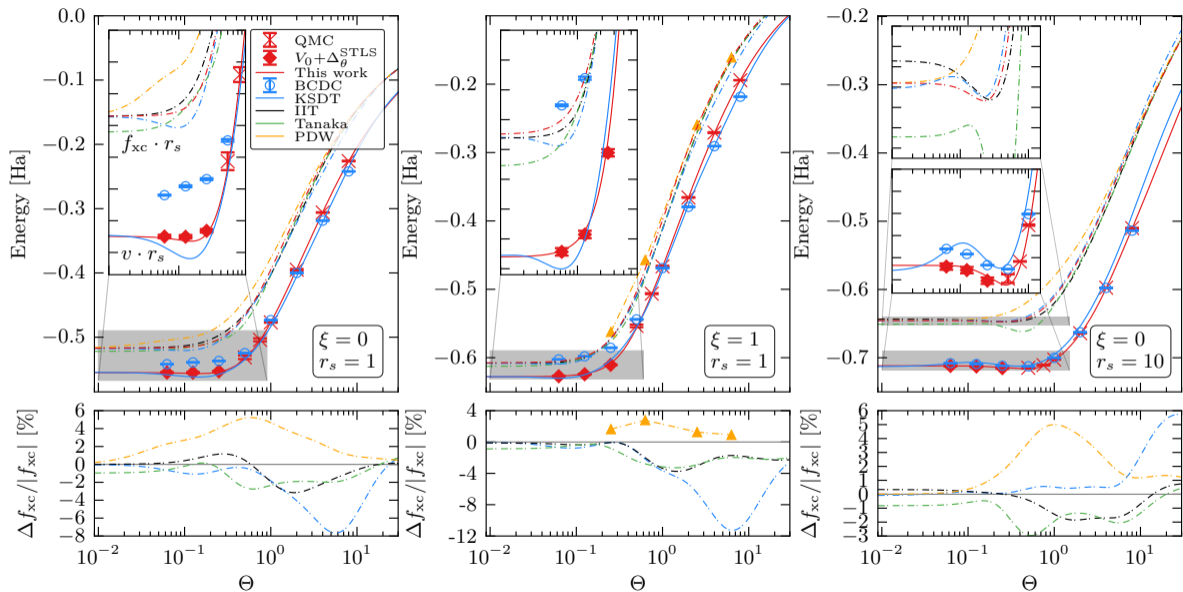
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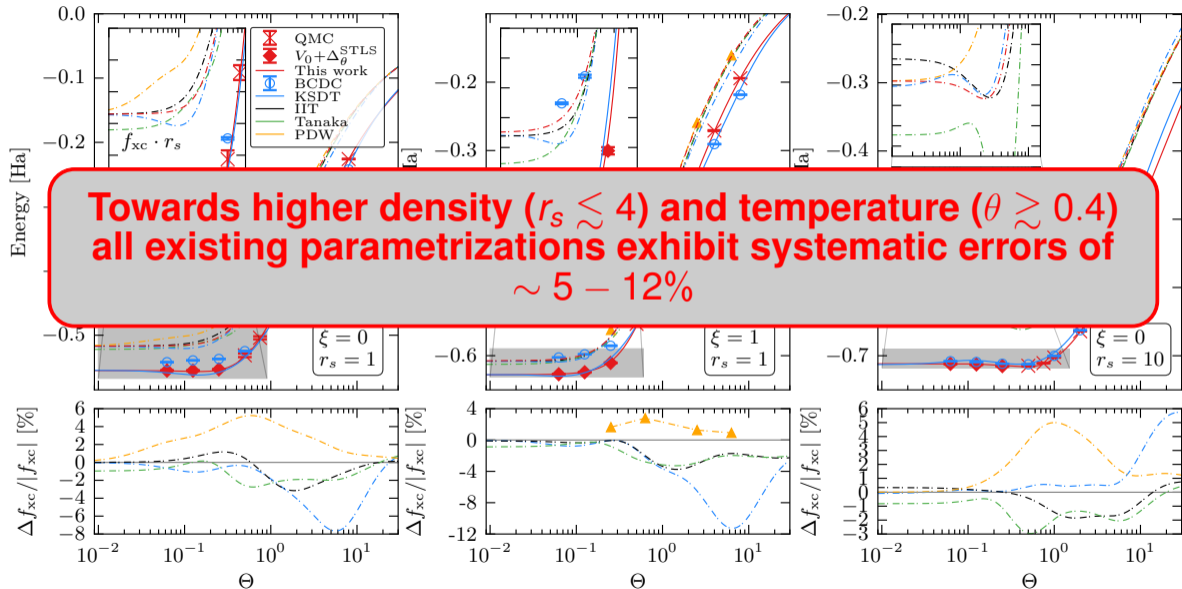
# Benchmarks for existing and future models and fits:

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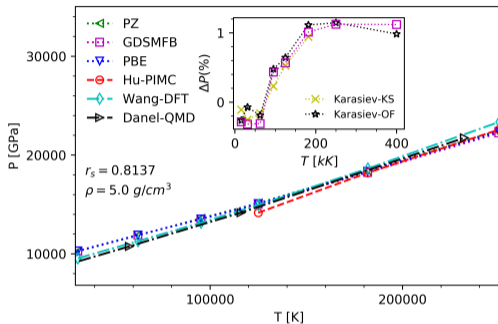
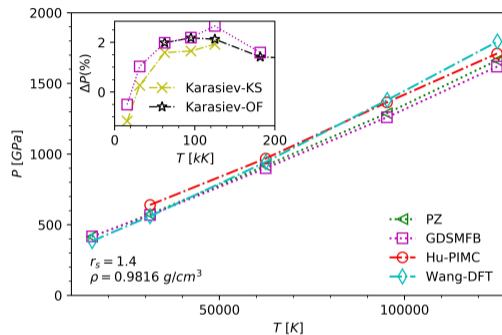
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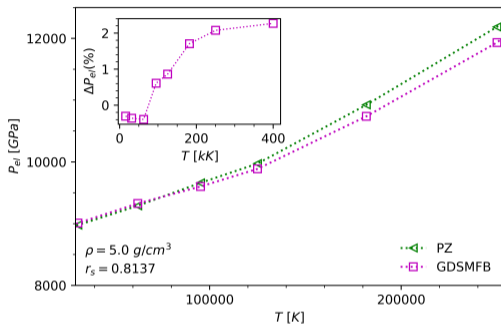
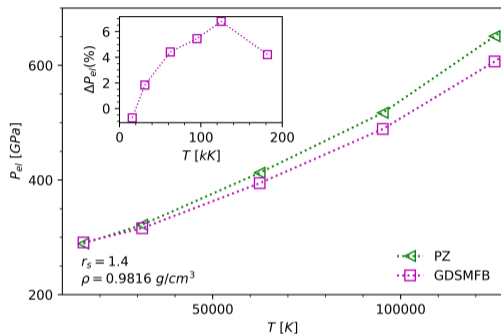


# Impact of finite-T functional on DFT-MD simulations: dense hydrogen<sup>1</sup>



**GDSMFB**: present finite-T functional, **PZ**: Perdew, Zunger (1981), **PBE**: Perdew, Burke, Enzerhof (1996)  
**Hu-PIMC**: Hu, Militzer, PRB 2011; **Wang-DFT**: Orbital-free MD, Phys. Plasmas (2013)  
**Karasiev et al.**, PRE (2016)

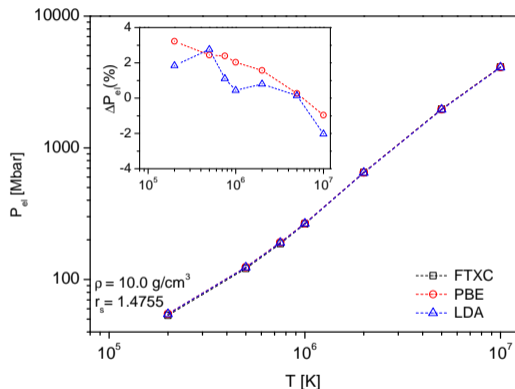
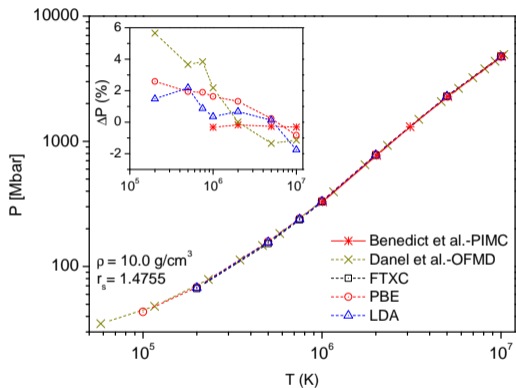
## Impact of finite-T functional: dense hydrogen, electronic pressure<sup>2</sup>



significantly larger differences in specific heat, susceptibilities

# Impact of finite-T functional on DFT-MD simulations: dense carbon<sup>3</sup>, $r_s \approx 1.5$

total pressure (left) vs. electronic pressure,  $E_F \approx 23\text{eV} \approx 267,000\text{K}$



**baseline=FTXC:** present finite-T functional (GDSMF), compared to  $T = 0$  PBE and LDA

**Danel:** Danel, Kazandjian, Piron, PRE 2018;

**Benedict:** Benedict, Driver, Hamel, Militzer, Qi, Correa, Saul, and Schwegler, Phys. Rev. B (2014)

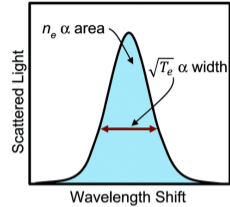
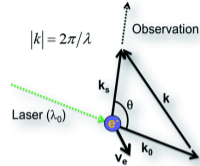
<sup>3</sup>Shen Zhang, unpublished (highly efficient high-T approach)

## Ab initio dynamic ( $\omega$ -dependent) results for the warm dense UEG

- **Key quantity:** dynamic structure factor

$$S(\mathbf{q}, \omega) := \frac{1}{2\pi} \int_{-\infty}^{\infty} dt \underbrace{\langle \hat{n}_{\mathbf{q}}(t) \hat{n}_{-\mathbf{q}}(0) \rangle}_{:=F(\mathbf{q}, t)} e^{i\omega t}$$

→ Directly measured in **scattering experiments**



**Thomson scattering:** signal depends on electron and ion temperature, density and correlations, degree of ionization, band structure of condensed phase etc.

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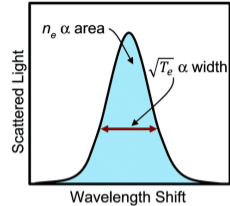
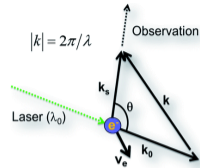
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- ▶ **Chihara decomposition** applies for non-collective scattering:

$$S(\mathbf{q}, \omega) = S_{b-b}(\mathbf{q}, \omega) + S_{b-f}(\mathbf{q}, \omega) + S_{f-f}(\mathbf{q}, \omega)$$

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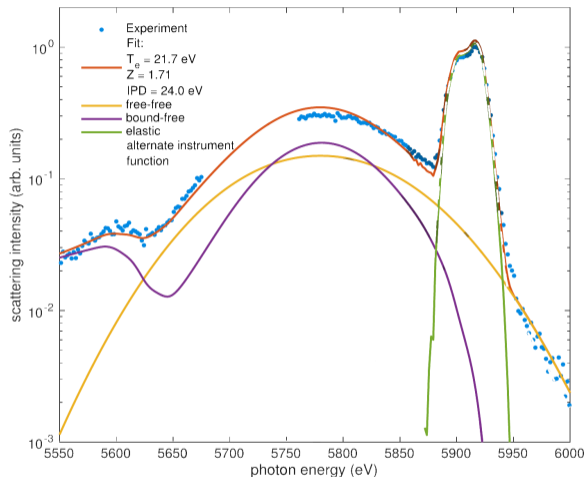
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- ▶ **Practical example:** Fit model for  $S(\mathbf{q}, \omega; T_e)$  to spectrum to determine electron temperature  $T_e$



**Figure:** Scattering spectrum of isochorically heated graphite at LCLS. Taken from D. Kraus *et al.*, *Plasma Phys. Control. Fusion* (2019)

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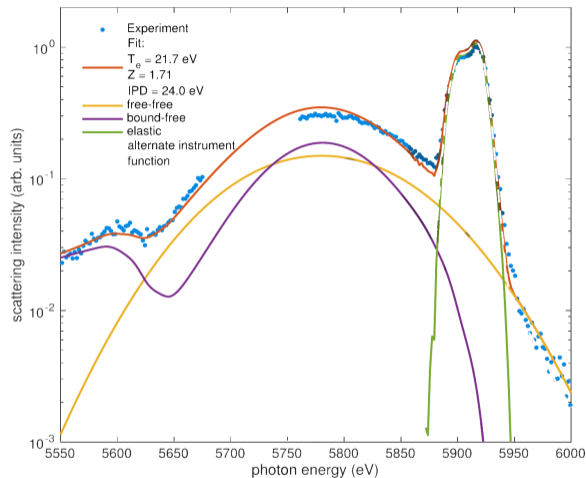
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## Density correlations: from real time to imaginary time

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→ Rigorous treatment of correlations **in general not feasible**

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**Alternative:** analytic continuation  $t \rightarrow -i\tau$

$$F(\mathbf{q}, \tau) = \int_{-\infty}^{\infty} d\omega S(\mathbf{q}, \omega) e^{-\tau\omega} \quad \forall \tau \in [0, \beta]$$

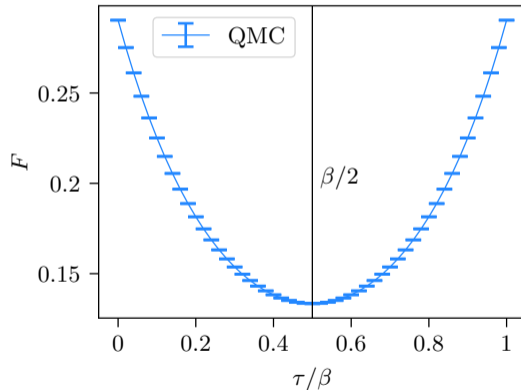
**Advantage:**

$F(\mathbf{q}, \tau)$  accessible within **PIMC** simulations

(**thermodynamic equilibrium** → *ab initio* feasible)

Imaginary-time density-density correlation function:

( $\theta = 1, r_s = 10, N = 34, q = 0.63q_F$ )



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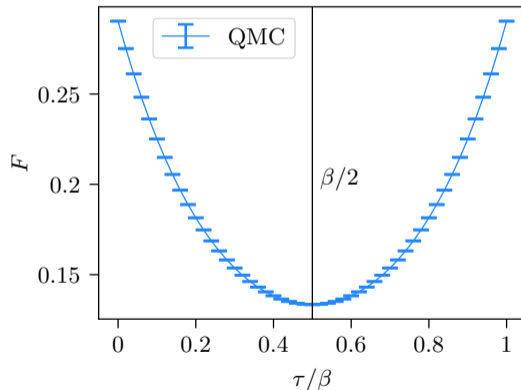
$F(\mathbf{q}, \tau)$  accessible within **PIMC** simulations  
(**thermodynamic equilibrium** → *ab initio* feasible)

**Disadvantage:**

$S(\mathbf{q}, \omega)$  requires **inverse Laplace transform**  
(**ill-posed problem** → sensitive to statistical errors in  $F$ )

Imaginary-time density-density correlation function:

( $\theta = 1, r_s = 10, N = 34, q = 0.63q_F$ )



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## Performing the inverse Laplace transformation for $S(\mathbf{q}, \omega)$

$$F(\mathbf{q}, \tau) = \int_{-\infty}^{\infty} d\omega S(\mathbf{q}, \omega) e^{-\tau\omega} \quad \forall \tau \in [0, \beta]$$

**Known frequency moments:**  $\langle \omega^{-1} \rangle, \langle \omega^0 \rangle, \langle \omega^1 \rangle, \langle \omega^3 \rangle$

$$\langle \omega^k \rangle = \int_{-\infty}^{\infty} d\omega \omega^k S(\mathbf{q}, \omega)$$

**General idea:**

- ▶ Find / optimize trial spectrum  $S^{\text{trial}}(\mathbf{q}, \omega)$  reproducing  $F(\mathbf{q}, \tau)$  and frequency moments

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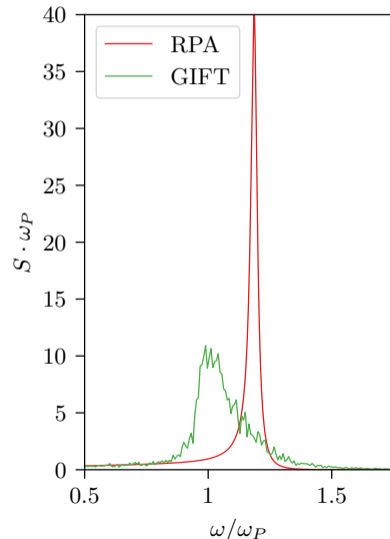
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**Method 1: Genetic Inversion via Falsification of Theories<sup>1</sup> (GIFT)**

- ▶ Stochastically evolve set of trial spectra  $\{S_i^{\text{trial}}(\mathbf{q}, \omega)\}$

Dynamic structure factor of the UEG:

( $\theta = 1, r_s = 10, N = 34, k = 0.63k_F$ )



<sup>1</sup>E. Vitali *et al.*, *Phys. Rev. B* **82**, 174510 (2010)

# Performing the inverse Laplace transformation for $S(\mathbf{q}, \omega)$

$$F(\mathbf{q}, \tau) = \int_{-\infty}^{\infty} d\omega S(\mathbf{q}, \omega) e^{-\tau\omega} \quad \forall \tau \in [0, \beta]$$

**Known frequency moments:**  $\langle \omega^{-1} \rangle, \langle \omega^0 \rangle, \langle \omega^1 \rangle, \langle \omega^3 \rangle$

$$\langle \omega^k \rangle = \int_{-\infty}^{\infty} d\omega \omega^k S(\mathbf{q}, \omega)$$

**General idea:**

- ▶ Find / optimize trial spectrum  $S^{\text{trial}}(\mathbf{q}, \omega)$  reproducing  $F(\mathbf{q}, \tau)$  and frequency moments

**Method 1: Genetic Inversion via Falsification of Theories<sup>1</sup> (GIFT)**

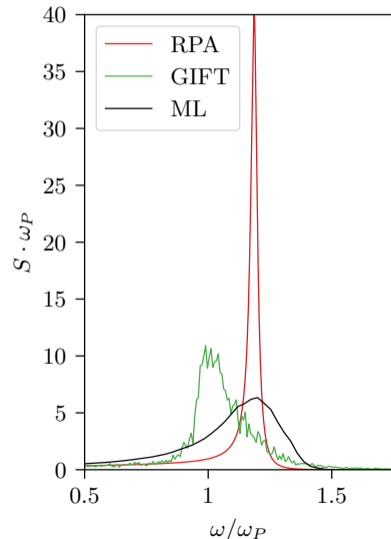
- ▶ Stochastically evolve set of trial spectra  $\{S_i^{\text{trial}}(\mathbf{q}, \omega)\}$

**Method 2: Machine learning (ML)**

- ▶ Compute versatile test set  $\{F_i, S_i\}, i = 1, \dots, 10^6$
- ▶ Train **neural net** to perform mapping  $\{F, \langle \omega^{-1} \rangle \dots \langle \omega^3 \rangle\} \rightarrow S$

Dynamic structure factor of the UEG:

( $\theta = 1, r_s = 10, N = 34, k = 0.63k_F$ )



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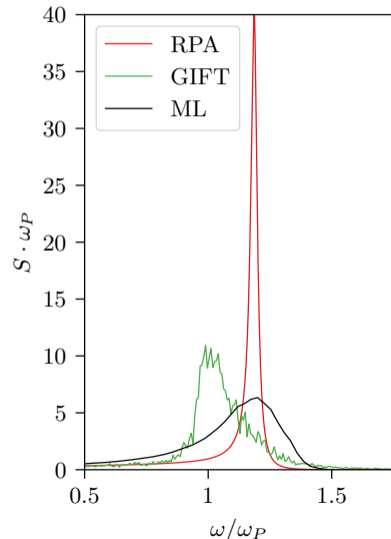
$$\langle \omega^k \rangle = \int_{-\infty}^{\infty} d\omega \omega^k S(\mathbf{q}, \omega)$$

GIFT and ML spectra reproduce correct  $F$  and  $\langle \omega^{-1} \rangle - \langle \omega^3 \rangle$

→ Knowledge of  $F$  and  $\langle \omega^{-1} \rangle - \langle \omega^3 \rangle$  not sufficient to determine  $S$

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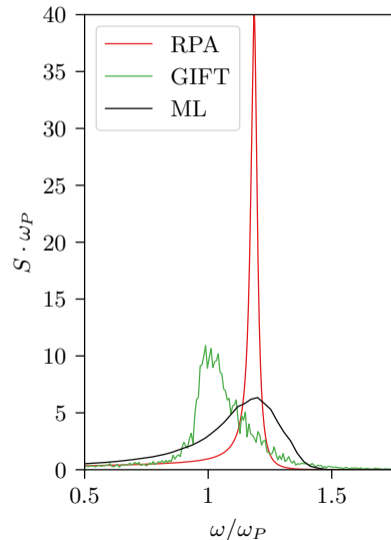


## Incorporating additional information on $S(\mathbf{q}, \omega)$ via dielectric formulation

► **Fluctuation-dissipation theorem:**

$$S(\mathbf{q}, \omega) = -\frac{\text{Im}\chi(\mathbf{q}, \omega)}{\pi n(1 - e^{-\beta\omega})}$$

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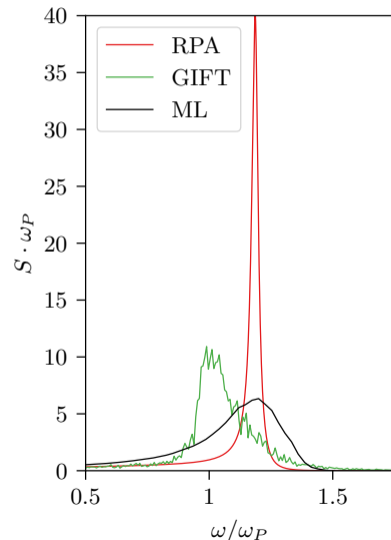
$$S(\mathbf{q}, \omega) = -\frac{\text{Im}\chi(\mathbf{q}, \omega)}{\pi n(1 - e^{-\beta\omega})}$$

- ▶ Express response function  $\chi$  via ideal response function  $\chi_0$  and **dynamic local field correction  $G$** :

$$\chi(\mathbf{q}, \omega) = \frac{\chi_0(\mathbf{q}, \omega)}{1 - v_q[1 - G(\mathbf{q}, \omega)]\chi_0(\mathbf{q}, \omega)}$$

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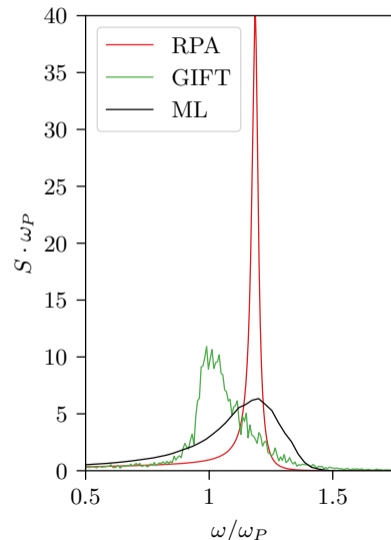
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**Make ansatz and optimize  $G(\mathbf{q}, \omega)$  instead of  $S(\mathbf{q}, \omega)$**

### Advantages:

- ▶ Limits  $G(\mathbf{q}, 0)$  and  $G(\mathbf{q}, \infty)$  known from PIMC simulation
- ▶ Other exact properties of  $G$  can be incorporated

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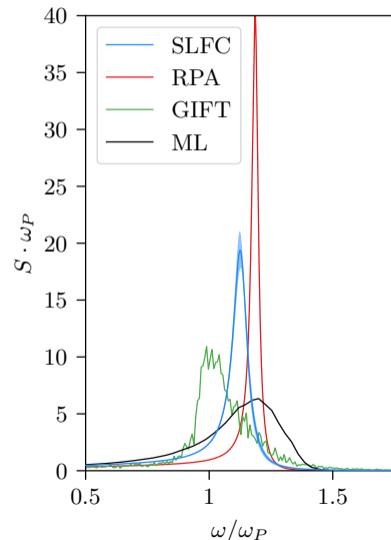
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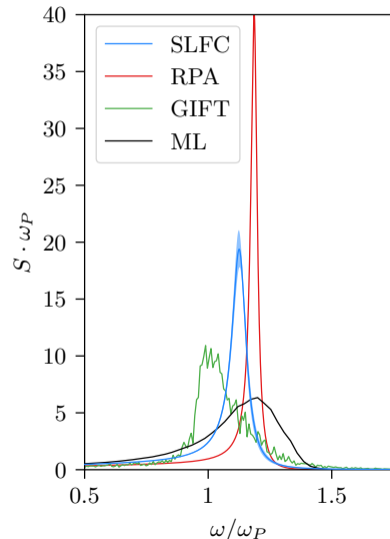
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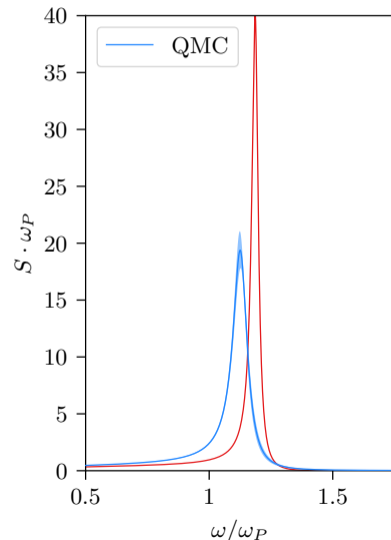
- ▶ **Random phase approximation (RPA):  $G \equiv 0$**

**Stochastic sampling of  $G(\mathbf{q}, \omega)$  accurately determines  $S(\mathbf{q}, \omega)$**

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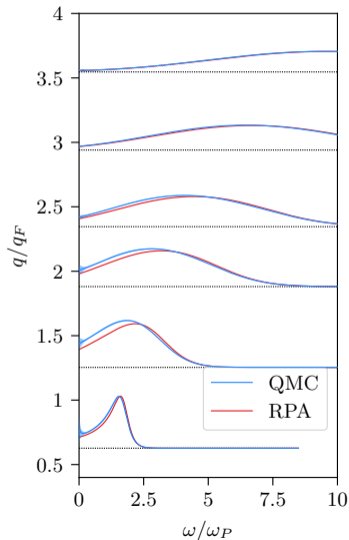


# Correlation effects in the dispersion relation: $\theta = 1$ , $r_s = 2$

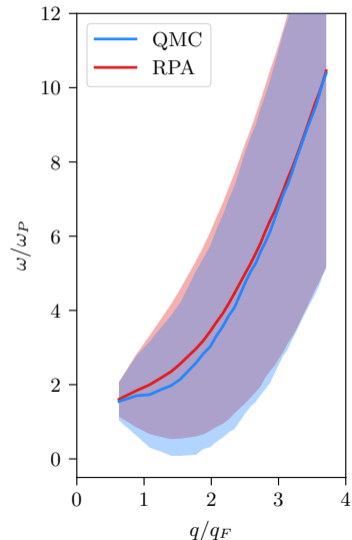
T. Dornheim, S. Groth, J. Vorberger, and M. Bonitz, *Phys. Rev. Lett.* **121**, 255001 (2018)

- ▶ Slight **correlation induced redshift** for intermediate  $q$  (at small  $r_s$ )

Dynamic structure factor of the UEG:



Peak position and FWHM:

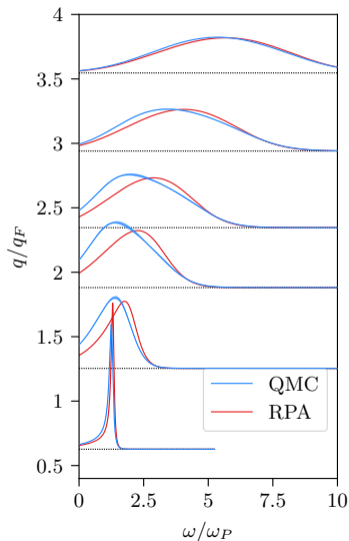


# Correlation effects in the dispersion relation: $\theta = 1$ , $r_s = 6$

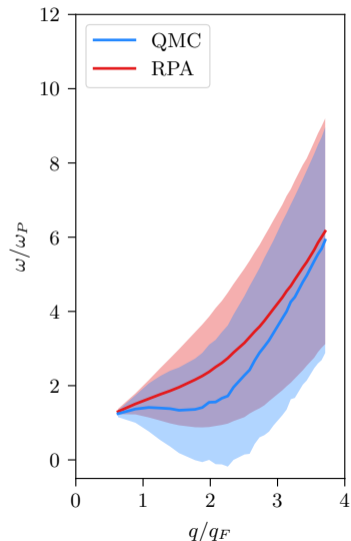
T. Dornheim, S. Groth, J. Vorberger, and M. Bonitz, *Phys. Rev. Lett.* **121**, 255001 (2018)

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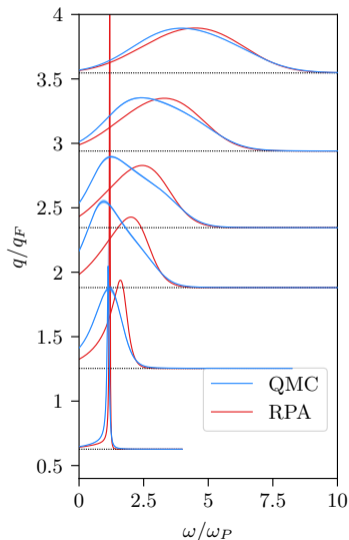


# Correlation effects in the dispersion relation: $\theta = 1$ , $r_s = 10$

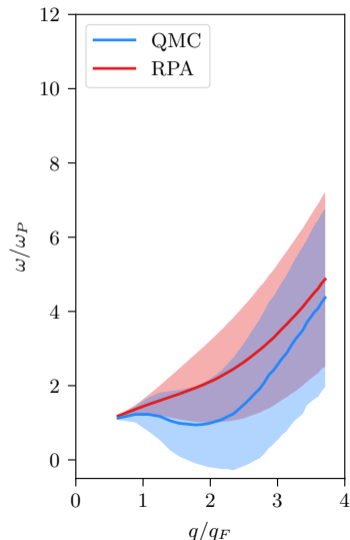
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- ▶ Slight **correlation induced redshift** for intermediate  $q$  (at small  $r_s$ )
- ▶ **Pronounced redshift and broadening** with increasing  $r_s$
- ▶ **Negative plasmon dispersion** for large  $r_s$  around  $q = 2q_F$  **predicted for dense hydrogen**
- ▶ **dispersion** and  $S(q, \omega)$  serve as rigorous benchmark for models

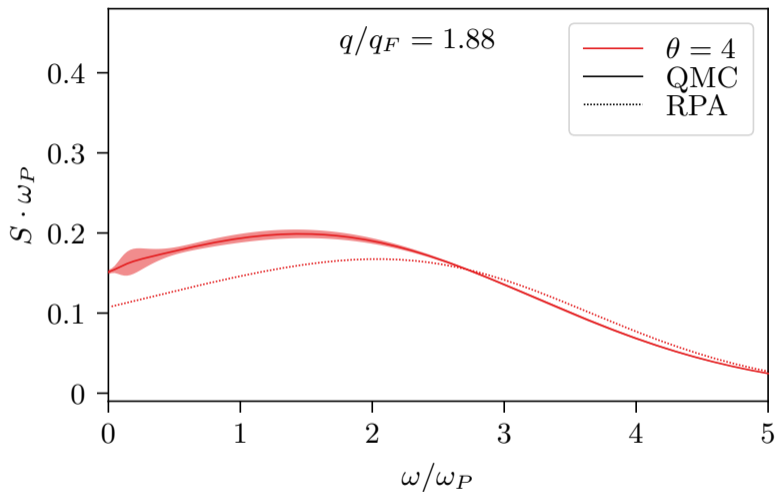
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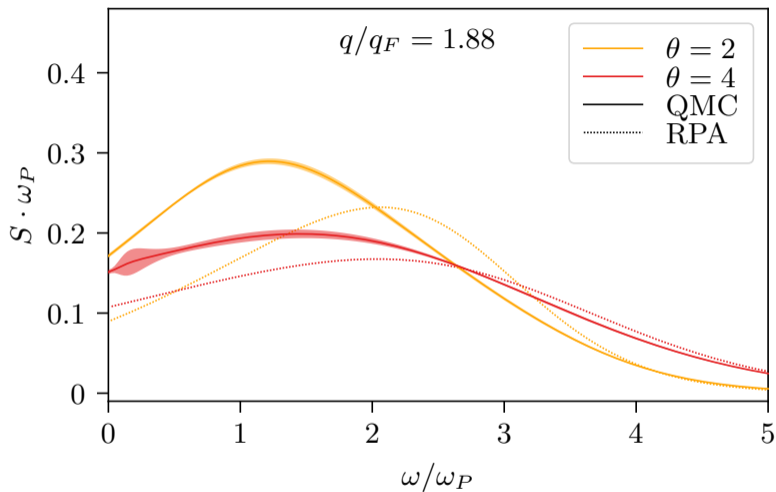


# Temperature effects in the dynamic structure factor at $r_s = 10$

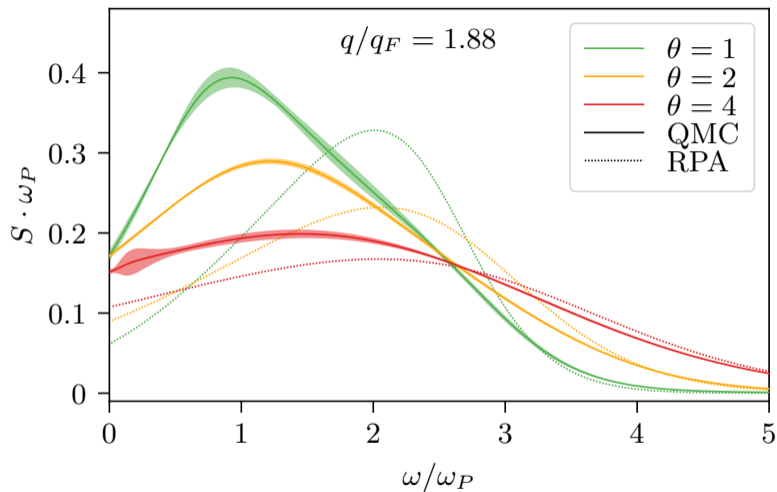




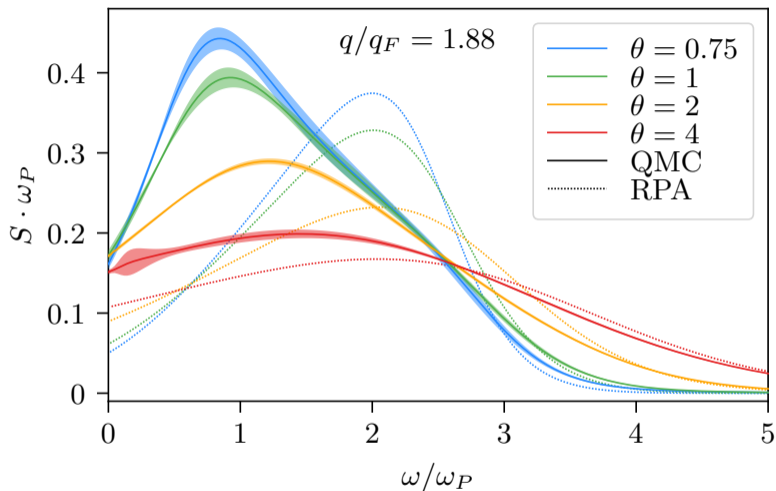
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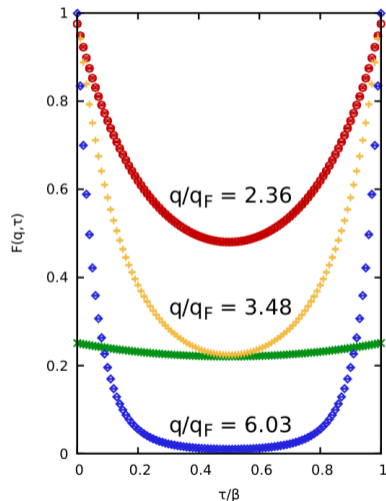
**Manifestation of non-trivial shapes in  $S(\mathbf{q}, \omega)$  towards low T**

# The Static Local Field Correction: *Ab initio* PIMC Simulations

- ▶ PIMC gives direct access to imaginary-time density–density correlation function:

$$F(\mathbf{q}, \tau) = \frac{1}{N} \langle \rho(\mathbf{q}, \tau) \rho(-\mathbf{q}, 0) \rangle$$

Source: S. Groth, TD, and J. Vorberger,  
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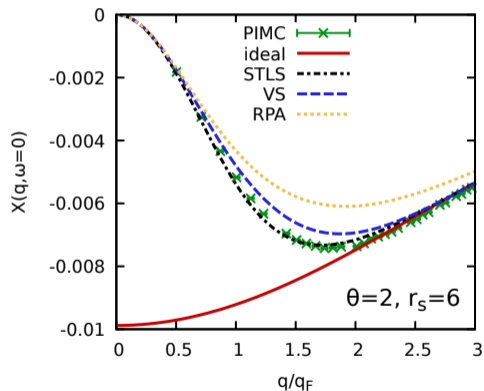
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- ▶  $F(\mathbf{q}, \tau)$  is directly connected to **static** density response  $\chi(\mathbf{q}) = \chi(\mathbf{q}, \omega = 0)$ :

$$\chi(\mathbf{q}) = -n \int_0^\beta d\tau F(\mathbf{q}, \tau)$$

→ Full  $\mathbf{q}$ -dependence from a single simulation of the **unperturbed UEG**

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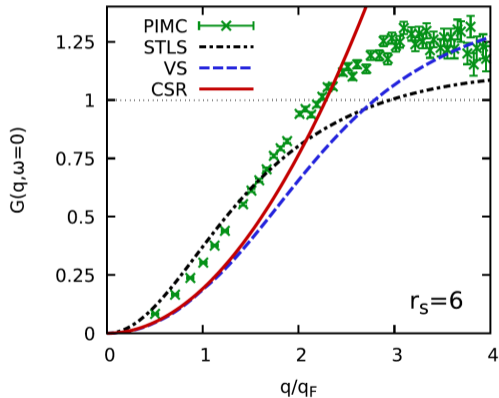
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- ▶  $G(q)$  can be obtained as the deviation from  $\chi_0(q)$ :

$$G(\mathbf{q}) = 1 - \frac{1}{v_q} \left( \frac{1}{\chi_0(\mathbf{q}, 0)} - \frac{1}{\chi(\mathbf{q})} \right) .$$

Source: S. Groth, **TD**, and J. Vorberger, *Phys. Rev. B* **99**, 235122 (2019)

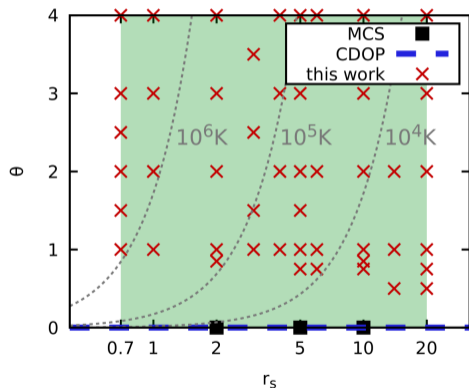


# The Static Local Field Correction: Neural-net representation

## Extensive set of new PIMC data

- ▶ QMC data available at discrete grid ( $q; \theta, r_s$ )

Source: T. Dornheim *et al.*,  
J. Chem. Phys. (2019)

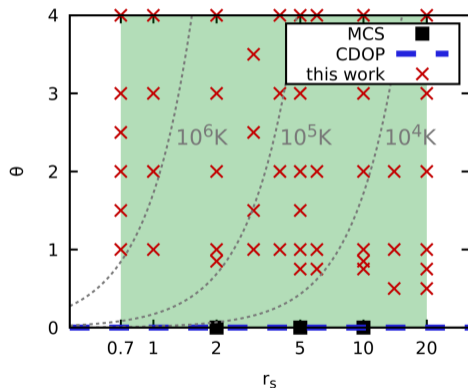


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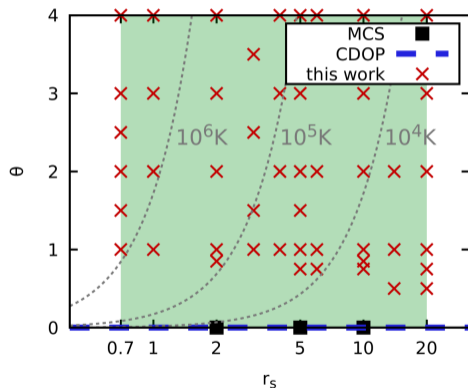


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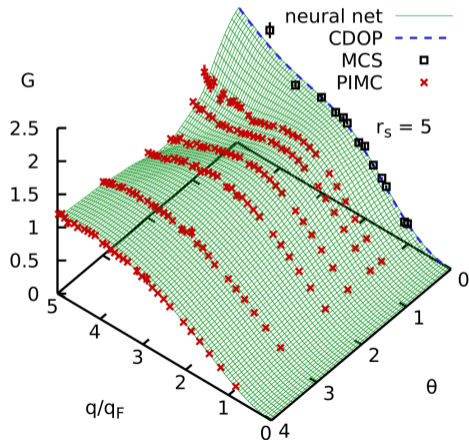


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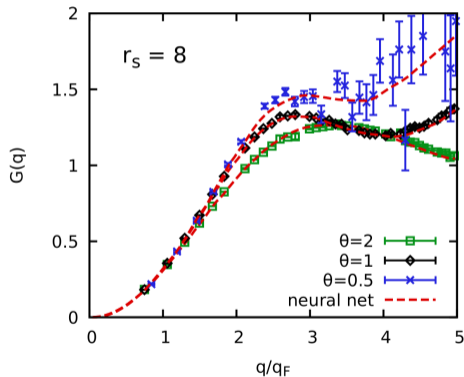


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- ▶ **Successful validation against independent data!**
- ▶ **Basis for transport quantities, screened ion potential**  
**Benchmarks for models and simulations**

Source: T. Dornheim *et al.*,  
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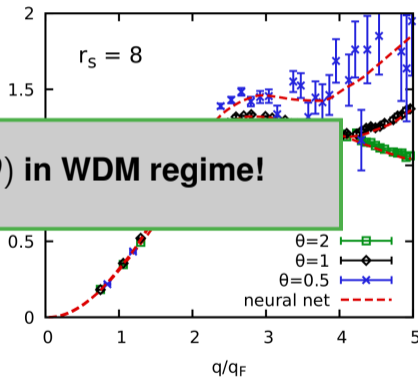


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**Complete *ab initio* data for  $G(q; r_s, \theta)$  in WDM regime!**

## Part II: Warm dense matter out of equilibrium

- ▶ **increasing experimental relevance:**

femtosecond pump-probe experiments with lasers and free electron lasers:  
XFEL (Hamburg), LCLS (Stanford), China  
e.g. A. Ng, S. Glenzer

- ▶ **Theoretical approaches**

1. Quantum kinetic theory<sup>4</sup>, nonequilibrium Green functions<sup>5</sup>: Gericke *et al.*, Vorberger *et al.*, Bornath *et al.*  
new simulations currently being developed in Kiel (WDM conference 2019)
2. Time-dependent DFT (TD-DFT), e.g. A. Baczewski *et al.*
3. (Quantum) Hydrodynamics

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<sup>4</sup>M. Bonitz, "Quantum Kinetic Theory", 2nd ed., Springer 2016

<sup>5</sup>D. Kremp *et al.*, "Quantum Statistics of Nonideal Plasmas", Springer 2005

## Real-Time Kadanoff-Baym Approach to Plasma Oscillations in a Correlated Electron Gas

N.-H. Kwong\* and M. Bonitz

*Fachbereich Physik, Universität Rostock, Universitätsplatz 3, D-18051 Rostock, Germany*

(Received 30 August 1999)

A nonequilibrium Green's functions approach to the collective response of correlated Coulomb systems at finite temperatures is presented. It is shown that solving Kadanoff-Baym-type equations of motion for the two-time correlation functions including the external perturbing field allows one to compute the plasmon spectrum with collision effects in a systematic and consistent way. The scheme has a "built-in" sum-rule preservation and is simpler to implement numerically than the equivalent equilibrium approach based on the Bethe-Salpeter equation.

PACS numbers: 73.20.Mf, 05.30.-d

- ▶ Relation to equilibrium approaches: M. Bonitz, "Quantum Kinetic Theory", 2nd ed., Springer (2016)
- ▶ Similar approach in TD-DFT (Bertsch, Yabana 1995)

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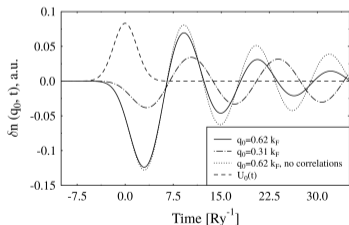


FIG. 1. Density fluctuation of a strongly correlated electron gas for two wave numbers. For comparison, the uncorrelated response for one wave number (dotted line) and the exciting field (dashes) are shown, too.  $k_F$  denotes the Fermi momentum, Ry = 13.6 eV.

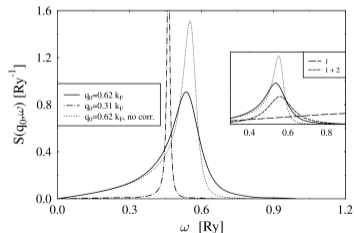


FIG. 2. Dynamic structure factor (12) for the correlated electron gas of Fig. 1 (same line styles). Inset shows  $S$  for  $q_0 = 0.62k_F$  and contains two other approximations to the correlations corresponding to retaining the first diagram in Eq. (17) and first plus second diagrams, respectively.

# Quantum hydrodynamics (QHD) for plasmas: problems and open questions

## ▶ Examples of quantum fluid theories

- ▶ Fermi liquid theory (Landau, Abrikosov...): metals, quasiparticles non-Fermi liquids (Luttinger)
- ▶ quantum spin liquids (magnetic materials); superfluid theory (Bogolyubov)

## ▶ kinetic theory approach: moments of the momentum distribution function:

$n \sim \langle v^0 \rangle$ ,  $\mathbf{u} \sim \langle \mathbf{v}^1 \rangle$ ,  $P \sim \langle \mathbf{v}\mathbf{v} \rangle$  etc.  $\Rightarrow$  hierarchy of moment equations

- ▶ a) truncation of the moment hierarchy (phenomenological closure)
- ▶ b) asymptotic schemes, small parameter, e.g.  $\epsilon = \lambda_{\text{mfp}}/L \ll 1$ ,

$$f(\mathbf{r}, \mathbf{v}, t) = f_0(\mathbf{r}, \mathbf{v}, t) + \epsilon f_1(\mathbf{r}, \mathbf{v}, t) + \dots$$

$$f_0(\mathbf{r}, \mathbf{v}, t) = n(\mathbf{r}, t) \left( \frac{m}{2\pi k_B T(\mathbf{r}, t)} \right)^{3/2} \exp \left[ -\frac{m(\mathbf{v} - \mathbf{u}(\mathbf{r}, t))^2}{2k_B T(\mathbf{r}, t)} \right],$$

$f_0$ : collision integrals vanish  $\Rightarrow$  no heat flux, no viscous stress

$f_1$  expanded into Laguerre/Sonine polynomials (Chapman-Enskog)

- ▶ **no similar mature approaches to quantum plasmas**, except for linear response theory (Zubarev, Röpke, ... generalized Gibbs ensemble)



# QHD for plasmas: problems and open questions

## Quantum hydrodynamics – status:

- ▶ straightforward for a single particle and for bosons
- ▶ fermions, plasmas: questionable assumptions in derivation<sup>a</sup>
- ▶ In many papers: incorrect coefficients in QHD equations<sup>b</sup>

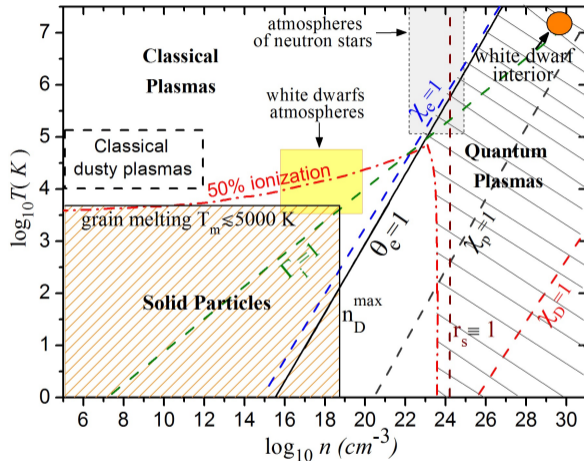
## questionable predictions:

- ▶ “novel attractive forces” between ions in quantum plasmas<sup>c</sup>
- ▶ “spinning quantum plasmas”,
- ▶ most recent example: “quantum dusty plasmas” (see figure)<sup>d</sup>

## questionable claims of relevance:

- ▶ semiconductors, metals
- ▶ white dwarfs, magnetars, neutron stars

Are they sure?? Can they prove it??  
how important are solitons in neutron stars??



<sup>a</sup>Manfredi, Haas, PRB 2001

<sup>b</sup>clarification: Michta, Graziani, Bonitz, Contrib. Plasma Phys. **55**, 437 (2015),  
Moldabekov *et al.*, Phys. Plasmas **25**, 031903 (2018), open access

<sup>c</sup>Shukla, Eliasson, PRL **108**, 165007 (2012); correction: Bonit *et al.*, PRE **87**, 033105 (2013), Moldabekov *et al.*, Phys. Plasmas (2015)

<sup>d</sup>Shukla, Ali, Phys. Plasmas **12**, 114502 (2005);

clarification: Bonitz, Moldabekov, Ramazanov, Phys. Plasmas **26**, 090601 (2019), perspectives article, Editors' pick, open access



**enforce high quality standards: manuscripts should<sup>6</sup>**

1. present an important **advance to plasma physics**;
2. contain a **convincing motivation** of the research in the introduction explaining the importance of the work;
3. present **applications to concrete plasmas** (in case of experiments) and—in case of theory—**clear predictions for real plasmas**, including their physical parameters;
4. not be purely formal studies of mathematical properties of equations, for example, in terms of dimensionless parameters, without application to real plasmas, as discussed in 3.

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<sup>6</sup>M. Bonitz, Editorial: Contrib. Plasma Phys. **59** (1), 8 (2019)

## Fluid description of quantum dynamics

- ▶ N interacting identical quantum particles, hamiltonian

$$\hat{H} = \sum_{i=1}^N \left( -\frac{\hbar^2}{2m} \nabla_i^2 + V(\mathbf{r}_i) \right) + \frac{1}{2} \sum_{i \neq j} w_{ij}(\mathbf{R}),$$

$\mathbf{R} = (\mathbf{r}_1, \sigma_1; \mathbf{r}_2, \sigma_2; \dots \mathbf{r}_N, \sigma_N)$ ,  $\mathbf{r}_i$  are the particle coordinates and  $\sigma_i$  their spin projections.

- ▶ for a pure state: dynamics governed by the **N-particle Schrödinger equation**

$$i\hbar \frac{\partial \Psi(\mathbf{R}, t)}{\partial t} = \hat{H} \Psi(\mathbf{R}, t), \quad \Psi(\mathbf{R}, t_0) = \Psi_0(\mathbf{R}),$$

normalization:  $\sum_{\sigma_1 \dots \sigma_N} \int d^{3N}R |\Psi(\mathbf{R}, t)|^2 = N$ ,

for particles with spin  $s$ :  $g_s = 2s + 1$  different spin projections

- ▶ ansatz with **real amplitude and phase** (Bohm, Madelung)<sup>7</sup>:

$$\Psi(\mathbf{R}, t) = A(\mathbf{R}, t) e^{\frac{i}{\hbar} S(\mathbf{R}, t)}, \quad A, S \in \mathcal{R}.$$

- ▶ can be used for semiclassical molecular dynamics, e.g. Gericke, Gregori *et al.* (2018)

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<sup>7</sup>E. Madelung, Z. Phys. (1927), D. Bohm, Phys. Rev. (1952)

## Quantum hydrodynamics for 1 particle

hydrodynamic fields from wave function (for pure state):

$$n(\mathbf{r}, t) = A^2(\mathbf{r}, t),$$

density,

$$\mathbf{p}(\mathbf{r}, t) = m\mathbf{v}(\mathbf{r}, t) = \nabla S(\mathbf{r}, t),$$

momentum,

exact “hydrodynamic” equations:

$$\frac{\partial n}{\partial t} + \nabla(\mathbf{v}n) = 0,$$

$$\frac{\partial \mathbf{p}}{\partial t} + \mathbf{v}\nabla\mathbf{p} = -\nabla(V + Q),$$

$$Q[n(\mathbf{r}, t)] = -\frac{\hbar^2}{2m} \frac{\nabla^2 n^{1/2}}{n^{1/2}},$$

“Bohm potential”,

### Modifications for N fermions:

- ▶ spin statistics (Pauli principle):  $N$  different orbitals  $\phi_i \Rightarrow$  different  $n_i, \mathbf{p}_i$
- ▶ interaction  $w_{ij}$  between particles
- ▶ statistical weight  $f_i$  of each orbital?
- ▶ coupling of orbitals? Definition of mean  $n$  and  $\mathbf{p}$ ?

## Statistical description of N-particle quantum system<sup>8</sup>

- ▶ mixed state described by **N-particle density operator**:

$$\hat{\rho}(t) = \sum_a \rho_a |\Psi^a(t)\rangle \langle \Psi^a(t)|, \quad \text{Tr} \hat{\rho}(t) = 1,$$

sum over projection operators on all solutions of the Schrödinger equation,  
 $\rho_a$ : real probabilities,  $0 \leq \rho_a \leq 1$ , with  $\sum_a \rho_a = 1$

- ▶ **one-particle reduced density operator**:

$$\hat{F}_1(t) \equiv N \text{Tr}_{2\dots N} \hat{\rho}(t), \quad \text{Tr}_1 \hat{F}_1 = N.$$

- ▶ equation of  $F_1$ : **quantum kinetic equation**,  $\hat{F}_{12} = \hat{F}_1 \hat{F}_2 + \hat{g}_{12}$ :

$$i\hbar \frac{\partial \hat{F}_1}{\partial t} - [\hat{H}_1, \hat{F}_1] = \text{Tr}_2 [\hat{w}_{12}, \hat{g}_{12}] \equiv \hat{l}_1 [\hat{g}_{12}], \quad \text{collision integral},$$

$$\hat{H}_1(t) = \hat{H}_1 + \hat{H}_1^H(t), \quad \hat{H}_1^H(t) \equiv \text{Tr}_2 \hat{w}_{12} \hat{F}_2(t), \quad \text{mean field},$$

---

<sup>8</sup>M. Bonitz, "Quantum Kinetic Theory", 2nd ed., Springer 2016

## Density matrix equation in Hartree approximation

- ▶ **mean field approximation** (no anti-symmetrization, no correlations):

$$\Psi(\mathbf{R}, t) \approx \phi_{\alpha_1}(\mathbf{r}_1) \cdot \phi_{\alpha_2}(\mathbf{r}_2) \cdots \phi_{\alpha_N}(\mathbf{r}_N), \quad \langle \phi_\alpha | \phi_\beta \rangle = \delta_{\alpha, \beta}$$

$\phi_{\alpha_k}(\mathbf{r}_k)$ : single-particle orbital occupied by particle  $k$ ,

for density operators: no antisymmetrization and  $\hat{l}_1 = 0 \Rightarrow \hat{F}_{12} \approx \hat{F}_1 \hat{F}_2$

- ▶ **coordinate representation**:  $\langle \mathbf{r}' | \hat{F}_1(t) | \mathbf{r}'' \rangle = f(\mathbf{r}', \mathbf{r}'', t)$ , density matrix

$$i\hbar \frac{\partial}{\partial t} f(\mathbf{r}', \mathbf{r}'', t) = -\frac{\hbar^2}{2m} \left( \nabla_{\mathbf{r}'}^2 - \nabla_{\mathbf{r}''}^2 \right) f(\mathbf{r}', \mathbf{r}'', t) + \left\{ U^{\text{eff}}(\mathbf{r}', t) - U^{\text{eff}}(\mathbf{r}'', t) \right\} f(\mathbf{r}', \mathbf{r}'', t)$$

$$U^{\text{eff}}(\mathbf{r}, t) = V(\mathbf{r}, t) + U^{\text{H}}(\mathbf{r}, t),$$

$$U^{\text{H}}(\mathbf{r}, t) = g_s \int d\bar{\mathbf{r}} w(r - \bar{r}) f(\bar{\mathbf{r}}, \bar{\mathbf{r}}, t),$$

diagonal element of the density matrix:  $f(\mathbf{r}, \mathbf{r}, t) = n(\mathbf{r}, t)$ , density

# Hartree approximation for fermions

- ▶ **mean field approximation** (no anti-symmetrization, no correlations):

$$\Psi(\mathbf{R}, t) \approx \phi_{\alpha_1}(\mathbf{r}_1) \cdot \phi_{\alpha_2}(\mathbf{r}_2) \cdots \phi_{\alpha_N}(\mathbf{r}_N), \quad \langle \phi_\alpha | \phi_\beta \rangle = \delta_{\alpha,\beta}$$

Thermodynamic equilibrium: orbitals  $\phi_1, \phi_2 \dots$  occupied with probability  $f_i = [e^{\beta(\epsilon_i - \mu)} + 1]^{-1}$ ,  $i = 1 \dots \infty$

- ▶ One-particle density operator and density matrix:

$$\hat{F}_1 = N \text{Tr}_{2 \dots N} \sum_{\alpha} \frac{1}{Z_G} e^{-\beta(E_{\alpha} - \mu N_{\alpha})} |\psi_{\alpha}\rangle \langle \psi_{\alpha}| = N \sum_{i=1}^{\infty} f_i |\phi_i\rangle \langle \phi_i|$$

$$f(\mathbf{r}', \mathbf{r}'', t) = N \sum_{i=1}^{\infty} f_i \phi_i(\mathbf{r}', t) \phi_i^*(\mathbf{r}'', t),$$

- ▶ Equation for  $f$  solved by set of nonlinear “Hartree-”Schrödinger (t-dependent Kohn-Sham) equations:

$$i\hbar \frac{\partial}{\partial t} \phi_i(\mathbf{r}, t) = \left\{ -\frac{\hbar^2}{2m} \nabla_{\mathbf{r}}^2 + V(\mathbf{r}, t) + U_F^H(\mathbf{r}, t) \right\} \phi_i(\mathbf{r}, t),$$

$$U_F^H[n(\mathbf{r}, t)] = \int d\mathbf{r}_2 w(\mathbf{r} - \mathbf{r}_2) [g_s n(\mathbf{r}_2, t) - n_0],$$

$U_F^H$ : Hartree mean field,  $n(\mathbf{r}, t) = \sum_{i=1}^{\infty} f_i |\phi_i(\mathbf{r}, t)|^2$ .  $n_0$ : background

# Microscopic QHD equations (MQHD)<sup>10</sup>

- ▶ in TDSE: introduce real amplitude ( $A_i = n_i^{1/2}$ ) and phase for each orbital:
- ▶ restore exchange-correlation effects formally exactly<sup>9</sup>: *either*
  - i) potential**  $V^{\text{xc}} = V^{\text{xc}}[n(\mathbf{r}, \tilde{t})]$ , from TD-DFT, *or, alternatively,*
  - ii) collision integral**  $I_i[g_{12}(\tilde{t})]$  from quantum kinetic theory

$$\begin{aligned}\frac{\partial n_i}{\partial t} + \nabla(\mathbf{v}_i n_i) &= 0, \\ \frac{\partial \mathbf{p}_i}{\partial t} + \mathbf{v}_i \text{div} \mathbf{p}_i &= -\nabla (V(\mathbf{r}, t) + U^{\text{H}} + Q_i + V^{\text{xc}} + I_i) , \\ U^{\text{H}}[n(\mathbf{r}, t)] &= \int d\mathbf{r}_2 w(\mathbf{r} - \mathbf{r}_2) [g_s n(\mathbf{r}_2, t) - n_0], \\ Q_i(\mathbf{r}, t) &= -\frac{\hbar^2}{2m} \frac{\nabla^2 \sqrt{n_i(\mathbf{r}, t)}}{\sqrt{n_i(\mathbf{r}, t)}} .\end{aligned}$$

- ▶  $n(\mathbf{r}, t) = \sum_{i=1}^{\infty} f_i \cdot |\phi_i(\mathbf{r}, t)|^2$ ,  $f_i$  statistical weight (Fermi function),  $g_s = 2s + 1$ .
- ▶  $\Rightarrow$  exact formulation of the many-body problem (with proper set of orbitals), **equivalent to TD-DFT**

---

<sup>9</sup>in general: memory effects (non-adiabatic), i.e.  $t_0 \leq \tilde{t} \leq t$

<sup>10</sup>Bonitz *et al.*, Phys. Plasmas **26**, 090601 (2019), agrees with Manfredi, Fields Inst. Comm. 2005



## Derivation of QHD equations from MQHD<sup>11</sup>

- ▶ consider mean field (Hartree / quantum Vlasov) approximation:  $V^{\text{xc}} \equiv I_j \equiv 0$
- ▶ average MQHD equations over  $i$  with weights  $f_i$
- ▶ define hydrodynamic (average) quantities:

$$\bar{n}(\mathbf{r}, t) = \frac{1}{N} \sum_{i=1}^{\infty} f_i n_i(\mathbf{r}, t),$$

$$\bar{\mathbf{p}}(\mathbf{r}, t) = \frac{1}{N} \sum_{i=1}^{\infty} f_i \mathbf{p}_i(\mathbf{r}, t),$$

$$\bar{Q}(\mathbf{r}, t) = -\frac{\hbar^2}{2mN} \sum_{i=1}^{\infty} f_i \frac{\nabla^2 \sqrt{n_i(\mathbf{r})}}{\sqrt{n_i(\mathbf{r})}} \neq \frac{\hbar^2}{2mN} \frac{\nabla^2 \sqrt{\bar{n}(\mathbf{r})}}{\sqrt{\bar{n}(\mathbf{r})}} \equiv Q_1[\bar{n}].$$

- ▶ final equality was *postulated* by Manfredi and Haas (PRB 2001, Fields Inst. Comm. 2005) reproducing *single-particle* QHD equations (with Hartree potential)
- ▶ When does it hold? What are the corrections?  $\Rightarrow$  perform strict derivation

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<sup>11</sup>Bonitz *et al.*, Phys. Plasmas 2019

## Derivation of QHD equations (contd.)<sup>12</sup>

- ▶ use average of product:  $\overline{a_i b_i} = \bar{a} \cdot \bar{b} + \overline{\delta a_i \delta b_i}$ ,
- ▶ result: **formally exact QHD equations** from orbital average of MQHD (TD-DFT):

$$\frac{\partial \bar{n}}{\partial t} + \frac{1}{m} \nabla \cdot (\bar{\mathbf{p}} \cdot \bar{\mathbf{n}}) = J_{np}^\Delta$$

$$\frac{\partial \bar{\mathbf{p}}}{\partial t} + \frac{1}{m} \bar{\mathbf{p}} \cdot \text{div} \bar{\mathbf{p}} = -\nabla \left( V(\mathbf{r}, t) + U_F^H[\bar{n}] + Q_1[\bar{n}] + Q^\Delta + V^{xc}[\bar{n}] \right) + J_{pp}^\Delta$$

- ▶ three correction terms (correlation functions):

$$J_{np}^\Delta = -\frac{1}{m} \nabla \cdot \overline{\delta \mathbf{p}_i \delta n_i}, \quad Q^\Delta \approx \frac{\hbar^2}{2m\bar{n}} \overline{\delta A_i \cdot \nabla^2 \delta A_i} + O\left(\left(\frac{\delta A_i}{A}\right)^2\right),$$

$$J_{pp}^\Delta = -\frac{1}{m} \overline{\delta \mathbf{p}_i \text{div} \delta \mathbf{p}_i} = \frac{1}{\bar{n}} \partial_\beta \bar{P}_{\alpha\beta} = \frac{1}{\bar{n}} \nabla \bar{P}_F + \frac{1}{\bar{n}} \partial_\gamma \bar{\sigma}_{\alpha\gamma}, \quad \gamma \neq \alpha, \quad \bar{\sigma} : \text{stress tensor},$$

- ▶ special case: **ideal Fermi gas** of dimension  $D = 1, 2, 3$ ,  $T = 0$ :  $\bar{\sigma}_{\alpha\gamma} \rightarrow 0$  and

$$J_{pp}^\Delta = \frac{1}{D} \nabla \cdot \overline{\delta E_{\text{kin}}} = \frac{m}{2(D+2)} \nabla v_F^2 = \frac{1}{2\bar{n}} \nabla \bar{P}^{\text{id}}.$$

# Test case plasma oscillations: linearization of MQHD-equations<sup>13</sup>

- ▶ stability conditions without excitation:

$$\begin{aligned}\partial_t n_{i0} &= \partial_t \mathbf{p}_{i0} = \bar{\mathbf{v}}_{i0} = 0 \\ \mathbf{v}_{i0} \operatorname{div} \mathbf{p}_{i0} &= -\nabla(U_0^H + Q_{i0})\end{aligned}$$

- ▶ monochromatic weak excitation:  $V_1^{\text{ext}}(\mathbf{r}, t) = \tilde{V}_1^{\text{ext}} e^{-i\hat{\omega}t + i\mathbf{q}\mathbf{r}}$ , where  $\hat{\omega} = \omega + i\epsilon$ ,  $\epsilon > 0$
- ▶ linearization and Fourier-Laplace transform:

$$\begin{aligned}n_{i0} &\rightarrow n_{i0} + n_{i1}; \mathbf{v}_{i0} \rightarrow \mathbf{v}_{i0} + \mathbf{v}_{i1}; \\ U_0^H &\rightarrow U_0^H + U_1^H; Q_{i0} \rightarrow Q_{i0} + Q_{i1}, \quad U_1^{\text{eff}} = V_1^{\text{ext}} + U_1^H\end{aligned}$$

- ▶ density response exactly coincides with random phase approximation (RPA)

$$\begin{aligned}\tilde{n}_1(\mathbf{q}, \hat{\omega}) &= \frac{1}{N} \sum_{i=1}^{\infty} f_i \tilde{n}_{i1}(\mathbf{q}, \hat{\omega}) = \tilde{U}_1^{\text{eff}}(\mathbf{q}, \hat{\omega}) \tilde{\Pi}_1^R(\mathbf{q}, \hat{\omega}) \\ \tilde{\Pi}_1^R(\mathbf{q}, \hat{\omega}) &= \frac{1}{N} \sum_{i=1}^{\infty} \frac{f_i}{(\hat{\omega} - \mathbf{q}\mathbf{v}_{i0})^2 - \frac{\hbar^2 \mathbf{q}^4}{4m^2}}.\end{aligned}$$

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<sup>13</sup>Bonitz *et al.*, Phys. Plasmas 2019, result agrees with Manfredi, Fields Inst. Comm. 2005

## Test case plasma oscillations: linearization of QHD equations<sup>15</sup>

- ▶ Hartree approximation, neglect correlation terms
- ▶ linearization and Fourier-Laplace transform:

$$\begin{aligned}\bar{n}_0 &\rightarrow \bar{n}_0 + \bar{n}_1; \quad \bar{\mathbf{v}}_0 \rightarrow \bar{\mathbf{v}}_0 + \bar{\mathbf{v}}_1; \quad U_0^H \rightarrow U_0^H + U_1^H; \\ Q[\bar{n}] &\rightarrow -\frac{\hbar^2}{2m} \frac{\nabla^2 \sqrt{\bar{n}_0}}{\sqrt{\bar{n}_0}} + \frac{\hbar^2}{4m} \nabla^2 \sqrt{\bar{n}_1}, \\ \bar{P}_0^{\text{id}} &\rightarrow \frac{2}{D+2} \bar{n}_0 E_F(\bar{n}_0) + \frac{2}{D} \frac{\bar{n}_1}{\bar{n}_0} E_F(\bar{n}_0),\end{aligned}$$

- ▶ plasmon dispersion – QHD vs. MQHD (RPA): agreement only in 1D, for  $\omega \geq \omega_{pl}$

$$\begin{aligned}\omega_{\text{QHD}}^2(q) &= \omega_{pl}^2 + \frac{1}{D} v_F^2 q^2 + \frac{\hbar^2}{4m^2} q^4, \\ \omega_{\text{MQHD}}^2(q) &= \omega_{pl}^2 + \frac{3}{D+2} v_F^2 q^2 + (1 - \delta_{2,D}) \frac{\hbar^2}{4m^2} q^4,\end{aligned}$$

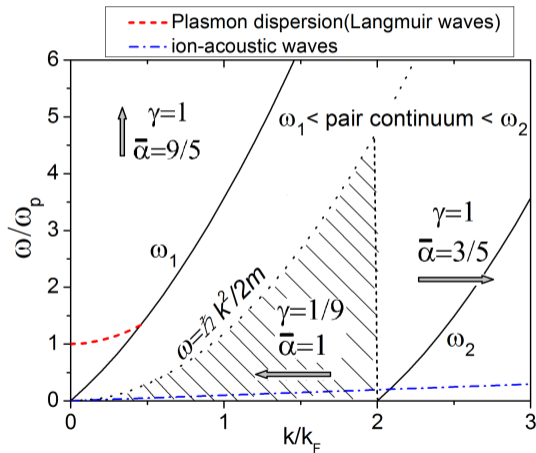
- ▶ for  $\omega < \omega_{pl}$ : Bohm potential  $Q$  is factor 9 too big  $\Rightarrow$  incorrect acoustic modes and statically screened potential<sup>14</sup>

<sup>14</sup>Michta *et al.*, CPP **55**, 437 (2015), Moldabekov *et al.*, Phys. Plasmas **25**, 031903 (2018)

<sup>15</sup>Bonitz *et al.*, Phys. Plasmas **26**, 090601 (2019)

## Correction of coefficients in QHD equations<sup>16</sup>

- ▶ Comparison to RPA reveals  $\omega$ - and  $k$ -dependence of pressure ( $\bar{\alpha}$ ) and Bohm potential ( $\gamma$ ).  
In addition: dependence on temperature, dimensionality



# Summary

- ▶ *ab initio* **QMC** simulations provide complete thermodynamic data for warm dense uniform electron gas<sup>17</sup>
- ▶ accurate functional  $f_{xc}(r_s, \Theta, \xi)$  input for finite-T LDA-DFT, implemented in [Libxc](#) (LDA\_XC\_GDSMFB)
- ▶ *ab initio* data for inhomogeneous EG<sup>18</sup>  $\Rightarrow$  accurate parametrization of static local field correction<sup>19</sup>  $G(q)$
- ▶ first *ab initio* data for the dynamic structure factor  $S(q, \omega)$  of warm dense electrons<sup>20</sup>

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- ▶ Electronic correlations and correlation build up (thermalization, dynamical screening, Auger processes etc.) are captured by (Nonequilibrium) **Green functions**. Highly efficient new computational techniques available<sup>22</sup>

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# Outlook: simulating WDM out of equilibrium—combination of methods promising

- ▶ **QHD** derived via averaging over all orbitals  $\phi_\alpha(\mathbf{r}) \Rightarrow$  no resolution of microscopic lengths scales,  $L_A \sim a_B, \lambda_{TF}$ , and associated times  $\Rightarrow$  advantage: extendable to large length and time scales

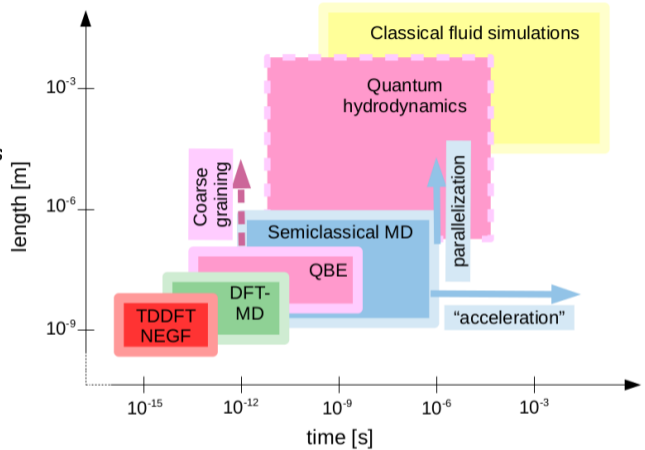


Figure: modified from M. Bonitz *et al.*, *Front. Chem. Science Engin.* (2019); QBE: quantum Boltzmann equation, NEGF: Nonequilibrium Green functions

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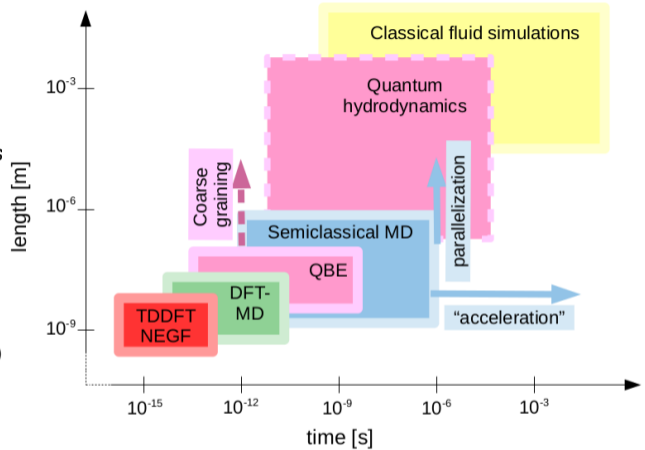


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- ▶ **DFT-MD** and **TD-DFT** capture atomic scales and inhomogeneity effects. But: miss electronic correlations (inaccurate band gap, ionization energies, no Auger processes etc.)
- ▶ **Green functions, quantum kinetic theory** capture electronic correlations, correlation build up and electron thermalization

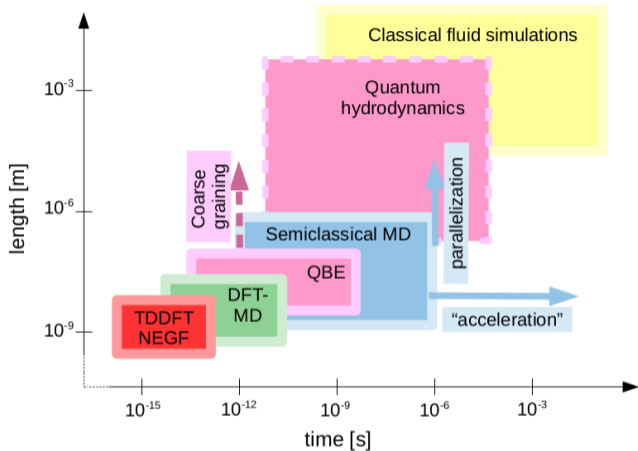


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## Outlook: Exciting developments in Warm Dense Matter are ahead of us

**Experimental progress requires ambitious theory and simulations  
Looking forward to great contributions from Indian plasma physics!**

### ▶ Extreme and exotic state of matter:

- High temperature:  $T \sim 10^4 - 10^8$  K
- Solid state density:  $n \sim 10^{21} - 10^{27}$  cm<sup>-3</sup>

### ▶ Characteristic parameters:

- ▶ Density parameter  $r_s = \bar{r}/a_B \sim 1$
- ▶ Degeneracy temperature  $\theta = k_B T/E_F \sim 1$
- ▶  $\Theta > 1$ : quantum plasma,  
 $\Theta < 1$ : classical plasma
- ▶ Classical coupling parameter  $\Gamma = e^2/r_s k_B T \sim 1$

