Towards *ab initio* simulation of quantum plasmas and warm dense matter

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October 2019





Introduction: warm dense matter

Warm dense matter (WDM):

- Nearly classical ions
- Degenerate non-ideal electrons
- Coupling parameter:

$$r_{s}=rac{\overline{r}}{a_{
m B}}\sim 0.1\dots 10$$

Degeneracy parameter:

$$heta = T/T_{\sf F} \sim 0.1 \dots 10$$



Figure: From T. Dornheim, S. Groth, and M. Bonitz, *Physics Reports* **744**, 1-86 (2018)

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 Temperature, degeneracy and coupling effects equally important
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log₁₀ n / cm⁻³

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Perturbation theory and ground-state approaches fail

Theoretical problems and simulation methods for WDM

1. WDM in equilibrium

- multiple components free electrons, ions, atoms, and (highly excited) solid
- \Rightarrow accessible only by DFT+MD

Problems of DFT:

- assumes electrons in ground state
- poor electronic correlations, binding energies, band gaps etc.
- ► \Rightarrow accurate static input: QMC

2. WDM out of equilibrium



Figure: modified from M. Bonitz et al., Front. Chem. Science Engin. (2019) QBE: quantum Boltzmann equation, NEGF: Nonequilibrium Green functions

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- time-dependent DFT (same problems as DFT), quantum kinetic theory, Nonequilibrium Green functions
- complementary strengths and limitations need combination of methods



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Importance of the uniform electron gas (UEG)

Model system of Coulomb interacting quantum electrons in a uniform positive background

Ground state (T = 0):

- Simple model for conduction electrons in metals
- Exchange-correlation (XC) energy:

$$e_{\mathrm{xc}}(r_s) = e_{\mathrm{tot}}(r_s) - e_0(r_s)$$

- \rightarrow Input for density functional theory (DFT) simulations (in LDA and GGA)
- \rightarrow Parametrization¹ of $e_{xc}(r_s)$ from ground state quantum Monte Carlo data²

¹ J.P. Perdew and A. Zunger, PRB 23, 5048 (1981) ² D.M. Ceperley and B. Alder, PRL 45, 566 (1980) ³ N.D. Mermin, Phys. Rev 137, A1441 (1965) ⁴ A.Y. Potekhin and G. Chabrier, A&A 550, A43 (2013)

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Warm dense matter ($T \sim T_F$):

► **Thermal DFT**³: minimize free energy F = E - TS→ Requires parametrization of XC **free** energy of UEG:

$$f_{xc}(r_s, \theta) = f_{tot}(r_s, \theta) - f_0(r_s, \theta)$$

- f_{xc}(r_s, θ) direct input for EOS models of astrophysical objects⁴
- f_{xc}(r_s, θ) contains complete thermodynamic information of UEG

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Many parametrizations for f_{xc} based on different approximate approaches:

- Semi-analytical approaches by Ebeling¹
- Dielectric methods, e.g. Singwi-Tosi-Land-Sjölander² (STLS) and Vashista-Singwi³ (VS)
- Quantum-classical mappings, e.g. Perrot and Dharma-wardana⁴ (PDW)
- Most recent: Fit by Karasiev⁵ et al. (KSDT) to Restricted Path Integral Monte Carlo (RPIMC) data⁶



¹ W. Ebeling and H. Lehmann, Ann. Phys. 45, (1988) ² S. Ichimaru, H. Iyetomi, and S. Tanaka, Phys. Rep. 149, (1987)

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⁵ V.V. Karasiev et al., PRL **112**, (2014)

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Accuracy of existing parametrizations for $f_{xc}(r_s, \theta)$ unclear



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Standard PIMC in warm dense regime severely hampered by fermion sign problem:



² T. Schoof *et al.*, Phys. Rev. Lett **115**, 130402 (2015) ³ T. Schoof *et al.*, Contrib. Plasma Phys. **51**, 687 (2011) E.W. Brown et al., PBL 110, 146405 (2013) ⁴ T. Dornheim et al., New J. Phys. 17, 073017 (2015)

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- > Standard PIMC in warm dense regime severely hampered by *fermion sign problem*:
 - First results¹ by E. Brown, D. Ceperley et al. (2013) based on fixed node approximation (RPIMC)
 - Induces systematic errors of unknown magnitude
 - **RPIMC** limited to $r_s \gtrsim 1$



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Our approach:

Avoid fermion sign problem by combining two exact and complementary QMC methods:



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1. Configuration PIMC (CPIMC)^{2,3}

 \rightarrow Excels at high density $\textit{r}_{s} \lesssim$ 1 and strong degeneracy



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Ab initio simulations over broad range of parameters

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PRIMC limited to $r_s \ge 1$



²T. Dornheim *et al.*, Phys. Rev. B **93**, 205134 (2016)

¹S. Groth *et al.*, Phys. Rev. B **93**, 085102 (2016) ³F.D. Malone *et al.*, Phys. Rev. Lett. **117**, 115701 (2016)

- **RPIMC** limited to $r_s \ge 1$
- CPIMC excels at high density



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- **RPIMC** limited to $r_s \ge 1$
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- **RPIMC** limited to $r_s \ge 1$
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Combination¹ yields exact results over entire density range down to $\theta \sim 0.5$

- Also applies to the unpolarized UEG²
- confirmed by independent DMQMC simulations³



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1. Accurate results for f_{xc} for finite N for paramagnetic and ferromagnetic cases^{1,2}

¹T. Schoof *et al.*, PRL **115**, 130402 (2015) ²S. Groth *et al.*, PRB **93**, 085102 (2016) ⁵T. Dornheim *et al.*, Phys. Plasmas **24**, 056303 (2017)

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- 2. Accurate finite size corrections and extrapolation to thermodynamic limit³, $N \rightarrow \infty$
- 3. Connect our data to ground state data, cover entire temperature range⁴

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- 4. New data for intermediate spin polarizations^{4,6}, 3-dimensional analytical parametrization⁷, $f_{xc}(r_s, \Theta, \xi) = f_{xc}$ -functional implemented in Libxc (LDA_XC_GDSMFB)

Comments and outlook:

- first unbiased tests of many earlier models and fits^{5,6}: STLS, VS, Ichimaru, Dharma wardana, Ebeling, Green functions (Kraeft, Vorberger, Rehr...) etc.: benchmark data allow for model improvement
- ▶ earlier fit (KSDT) by Karasiev et al.: tested and subsequently corrected, now good agreement with GDSMFB

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- ▶ earlier fit (KSDT) by Karasiev et al.: tested and subsequently corrected, now good agreement with GDSMFB
- derivatives $\partial f_{xc}/\partial n$ and $\partial f_{xc}/\partial T$ potentially inaccurate; \Rightarrow fits to separate ab initio data needed

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Benchmarks for existing and future models and fits:

S. Groth et al., PRL (2017)



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S. Groth et al., PRL (2017)



Impact of finite-T functional on DFT-MD simulations: dense hydrogen¹



GDSMFB: present finite-T functional, **PZ**: Perdew, Zunger (1981), **PBE**: Perdew, Burke, Enzerhof (1996) **Hu-PIMC**: Hu, Militzer, PRB 2011; **Wang-DFT**: Orbital-free MD, Phys. Plasmas (2013) **Karasiev** *et al.*, PRE (2016)

¹Kushal Ramakrishna, unpublished

Impact of finite-T functional: dense hydrogen, <u>electronic</u> pressure²



significantly larger differences in specific heat, susceptibilities

²Kushal Ramakrishna, unpublished

Impact of finite-T functional on DFT-MD simulations: dense carbon³, $r_s \approx 1.5$ total pressure (left) vs. electronic pressure, $E_F \approx 23$ eV $\approx 267,000$ K



baseline=FTXC: present finite-T functional (GDSMFB), compared to T = 0 **PBE** and **LDA Danel**: Danel, Kazandjian, Piron, PRE 2018;

Benedict: Benedict, Driver, Hamel, Militzer, Qi, Correa, Saul, and Schwegler, Phys. Rev. B (2014)

³Shen Zhang, unpublished (highly efficient high-T approach)

► Key quantity: dynamic structure factor

$$S(\mathbf{q},\omega) := rac{1}{2\pi} \int_{-\infty}^{\infty} \mathrm{d}t \; \underbrace{\langle \hat{n}_{\mathbf{q}}(t) \hat{n}_{-\mathbf{q}}(0)
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- \rightarrow Directly measured in scattering experiments
- Chihara decomposition applies for non-collective scattering:

$$\begin{split} & S(\mathbf{q},\omega) = S_{\text{b-b}}(\mathbf{q},\omega) + S_{\text{b-f}}(\mathbf{q},\omega) + S_{\text{l-f}}(\mathbf{q},\omega) \\ & \rightarrow S_{\text{l-f}}(\mathbf{q},\omega) \sim S^{\text{UEG}}(\mathbf{q},\omega) \end{split}$$

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 Practical example: Fit model for S(q, ω; T_e) to spectrum to determine electron temperature T_e



Figure: Scattering spectrum of isochorically heated graphite at LCLS. Taken from D. Kraus *et al.*, *Plasma Phys. Control. Fusion* (2019)

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Problem:

 $F(\mathbf{q}, t)$ requires **real time-dependent simulations** \rightarrow with PIMC have to use analytic continuation, reconstruct F(q, it) and 4 frequency moments, but: insufficient information



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Incorporating additional information on $S(\mathbf{q}, \omega)$ via dielectric formulation

Fluctuation-dissipation theorem:

$$S(\mathbf{q},\omega) = -rac{{
m Im}\chi(\mathbf{q},\omega)}{\pi n(1-e^{-eta\omega})}$$

Dynamic structure factor of the UEG: $\overline{(\theta = 1, r_s = 10, N = 33, q = 0.63q_F)}$ 40RPA 35GIFT ML30 25 $S\cdot \omega_P$ 201510 M $\mathbf{5}$ 0 0.51.5 ω/ω_P
Fluctuation-dissipation theorem:

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• Express response function χ via ideal response function χ_0 and dynamic local field correction *G*:

$$\chi(\mathbf{q},\omega) = \frac{\chi_0(\mathbf{q},\omega)}{1 - v_q [1 - G(\mathbf{q},\omega)]\chi_0(\mathbf{q},\omega)}$$

• Random phase approximation (RPA): $G \equiv 0$

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Make ansatz and optimize $G(q, \omega)$ instead of $S(q, \omega)$

Advantages:

- Limits $G(\mathbf{q}, 0)$ and $G(\mathbf{q}, \infty)$ known from PIMC simulation
- Other exact properties of G can be incorporated

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 ω/ω_P

0.5

1.5

Fluctuation-dissipation theorem:

 $S(\mathbf{q}, \omega) = -\frac{\text{Im}\chi(\mathbf{q}, \omega)}{\text{GIFT and ML spectra not in agreement}}$ with exact properties of $G(\mathbf{q}, \omega)$ \rightarrow to be discarded as unphysical

$$\chi(\mathbf{q},\omega) = \frac{1 - v_q [1 - G(\mathbf{q},\omega)] \chi_0(\mathbf{q},\omega)}{1 - v_q [1 - G(\mathbf{q},\omega)] \chi_0(\mathbf{q},\omega)}$$

• Random phase approximation (RPA): $G \equiv 0$

Stochastic sampling of $G(\mathbf{q}, \omega)$ accurately determines $S(\mathbf{q}, \omega)$

Advantages:

- Limits $G(\mathbf{q}, 0)$ and $G(\mathbf{q}, \infty)$ known from PIMC simulation
- Other exact properties of G can be incorporated

Dynamic structure factor of the UEG: $\overline{(\theta = 1, r_{s} = 10, N = 33, q = 0.63q_{\rm F})}$ 40QMC 3530 25dm. 20S 151050.51.5

 ω/ω_P

Correlation effects in the dispersion relation: $\theta = 1$, $r_s = 2$

T. Dornheim, S. Groth, J. Vorberger, and M. Bonitz, Phys. Rev. Lett. 121, 255001 (2018)



 Slight correlation induced redshift for intermediate q (at small r_s)

Correlation effects in the dispersion relation: $\theta = 1$, $r_s = 6$

T. Dornheim, S. Groth, J. Vorberger, and M. Bonitz, Phys. Rev. Lett. 121, 255001 (2018)

- Peak position and FWHM: Dynamic structure factor of the UEG: 12QMC RPA 3.5103 8 2.5 ω/ω_P q/q_F 6 $\mathbf{2}$ 4 1.5QMC $\mathbf{2}$ 1 RPA 0 0.5 -2.57.50 5 10 0 2 3 ω/ω_P q/q_F
- Slight correlation induced redshift for intermediate q (at small r_s)
- Pronounced redshift and broadening with increasing r_s

Correlation effects in the dispersion relation: $\theta = 1$, $r_s = 10$

T. Dornheim, S. Groth, J. Vorberger, and M. Bonitz, Phys. Rev. Lett. 121, 255001 (2018)

Peak position and FWHM: Dynamic structure factor of the UEG: 124 QMC RPA 3.510 3 8 2.5 ω/ω_P q/q_F 6 $\mathbf{2}$ 4 1.5QMC $\mathbf{2}$ 1 RPA 0 0.52.57.55 100 3 ω/ω_P q/q_F

 Slight correlation induced redshift for intermediate q (at small r_s)

- Pronounced redshift and broadening with increasing r_s
- Negative plasmon dispersion for large r_s around q = 2q_F predicted for dense hydrogen

The Static Local Field Correction: Ab initio PIMC Simulations

 PIMC gives direct access to imaginary-time density-density correlation function:

$$F(\mathbf{q}, au) = rac{1}{N} raket{
ho}(\mathbf{q}, au)
ho(-\mathbf{q},0)$$



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 PIMC gives direct access to imaginary-time density-density correlation function:

$$F(\mathbf{q}, \tau) = rac{1}{N} \langle
ho(\mathbf{q}, \tau)
ho(-\mathbf{q}, 0) \rangle$$

• $F(\mathbf{q}, \tau)$ is directly connected to <u>static</u> density response $\chi(\mathbf{q}) = \chi(\mathbf{q}, \omega = 0)$:

$$\chi(\mathbf{q}) = -n \int_0^eta \mathrm{d} au \ F(\mathbf{q}, au)$$

 \rightarrow Full q-dependence from a single simulation of the unperturbed UEG

Source: S. Groth, **TD**, and J. Vorberger, *Phys. Rev. B* **99**, 235122 (2019)



The Static Local Field Correction: Ab initio PIMC Simulations

 PIMC gives direct access to imaginary-time density-density correlation function:

$$F(\mathbf{q}, \tau) = \frac{1}{N} \langle \rho(\mathbf{q}, \tau) \rho(-\mathbf{q}, 0) \rangle$$

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$$\chi(\mathbf{q}) = -n \int_0^eta \mathrm{d} au \ F(\mathbf{q}, au)$$

 \rightarrow Full q-dependence from a single simulation of the unperturbed UEG

• G(q) can be obtained as the deviation from $\chi_0(q)$:

$$G(\mathbf{q}) = 1 - rac{1}{v_q} \left(rac{1}{\chi_0(\mathbf{q},0)} - rac{1}{\chi(\mathbf{q})}
ight)$$



Extensive set of new PIMC data

• QMC data available at discrete grid $(q; \theta, r_s)$



Extensive set of new PIMC data

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- <u>Problem:</u> Applications (DFT, hydrodynamics, ...) typically require continuous representation



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- Complicated, non-trivial behavior of G(q; r_s, θ), only few analytical limits are known

Source: T. Dornheim et al., J. Chem. Phys. (2019) хх MCS × CDOP this work X × × × 10⁶K 10⁵κ × $0^4 K$ × XXX × × ×

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Φ 2

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- Solution: Neural net as flexible function approximator



Extensive set of new PIMC data

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- <u>Problem:</u> Applications (DFT, hydrodynamics, ...) typically require continuous representation
- Complicated, non-trivial behavior of G(q; r_s, θ), only few analytical limits are known
- Solution: Neural net as flexible function approximator
- Successful validation against independent data!
- Basis for transport quantities, screened ion potential Benchmarks for models and simulations







Part II: Warm dense matter out of equilibrium

increasing experimental relevance:

femtosecond pump-probe experiments with lasers and free electron lasers: XFEL (Hamburg), LCLS (Stanford), China e.g. A. Ng, S. Glenzer

Theoretical approaches

- 1. Quantum kinetic theory⁴, nonequilibrium Green functions⁵: Gericke *et al.*, Vorberger *et al.*, Bornath *et al.* new simulations currently being developed in Kiel (WDM conference 2019)
- 2. Time-dependent DFT (TD-DFT), e.g. A. Baczewski et al.
- 3. (Quantum) Hydrodynamics

⁴M. Bonitz, "Quantum Kinetic Theory", 2nd ed., Springer 2016

⁵D. Kremp et al., "Quantum Statistics of Nonideal Plasmas", Springer 2005

Nonequilibrium Green functions approach to the dynamical structure factor

VOLUME 84, NUMBER 8 PHYSICAL REVIEW LETTERS

21 February 2000

Real-Time Kadanoff-Baym Approach to Plasma Oscillations in a Correlated Electron Gas

N.-H. Kwong* and M. Bonitz

Fachbereich Physik, Universität Rostock, Universitätsplatz 3, D-18051 Rostock, Germany (Received 30 August 1999)

A nonequilibrium Green's functions approach to the collective response of correlated Coulomb systems at finite temperatures is presented. It is shown that solving Kadanoff-Baym-type equations of motion for the two-time correlation functions including the external perturbing field allows one to compute the plasmon spectrum with collision effects in a systematic and consistent way. The scheme has a "built-in" sum-rule preservation and is simpler to implement numerically than the equivalent equilibrium approach based on the Bethe-Salpeter equation.

PACS numbers: 73.20.Mf, 05.30.-d

- Relation to equilbrium approaches: M. Bonitz, "Quantum Kinetic Theory", 2nd ed., Springer (2016)
- Similar approach in TD-DFT (Bertsch, Yabana 1995)

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FIG. 1. Density fluctuation of a strongly correlated electron gas for two wave numbers. For comparison, the uncorrelated response for one wave number (dotted line) and the exciting field (dashes) are shown, too. k_F denotes the Fermi momentum, Ry = 13.6 eV.

FIG. 2. Dynamic structure factor (12) for the correlated electron gas of Fig. 1 (same line styles). Inset shows S for $q_0 = 0.62k_r$ and contains two other approximations to the correlations corresponding to retaining the first diagram in Eq. (17) and first plus second diagrams, respectively.

Quantum hydrodynamics (QHD) for plasmas: problems and open questions

Examples of quantum fluid theories

- Fermi liquid theory (Landau, Abrikosov...): metals, quasiparticles non-Fermi liquids (Luttinger)
- quantum spin liquids (magnetic materials); superfluid theory (Bogolyubov)
- kinetic theory approach: moments of the momentum distribution function:

 $n \sim \langle v^0 \rangle$, $\mathbf{u} \sim \langle \mathbf{v}^1 \rangle$, $P \sim \langle \mathbf{vv} \rangle$ etc. \Rightarrow hierarchy of moment equations

- > a) truncation of the moment hierarchy (phenomenological closure)
- ▶ b) asymptotic schemes, small parameter, e.g. $\epsilon = \lambda_{
 m mfp}/L \ll$ 1,

$$f(\mathbf{r},\mathbf{v},t) = f_0(\mathbf{r},\mathbf{v},t) + \epsilon f_1(\mathbf{r},\mathbf{v},t) + \ldots$$

$$f_0(\mathbf{r},\mathbf{v},t) = n(\mathbf{r},t) \left(\frac{m}{2\pi k_B T(\mathbf{r},t)}\right)^{3/2} \exp\left[\frac{m(\mathbf{v}-\mathbf{u}(\mathbf{r},t))^2}{2k_B T(\mathbf{r},t)}\right],$$

 f_0 : collision integrals vanish \Rightarrow no heat flux, no viscous stress

f1 expanded into Laguerre/Sonine polynomials (Chapman-Enskog)

no similar mature approaches to quantum plasmas, except for linear response theory (Zubarev, Röpke, ... generalized Gibbs ensemble)

QHD for plasmas: problems and open questions

Quantum hydrodynamics - status:

- straightforward for a single particle and for bosons
- fermions, plasmas: questionable assumptions in derivation^a
- In many papers: incorrect coefficients in QHD equations^b

questionable predictions:

- "novel attractive forces" between ions in quantum plasmas^c
- "spinning quantum plasmas",
- most recent example: "quantum dusty plasmas" (see figure)^d

questionable claims of relevance:

- semiconductors, metals
- white dwarfs, magnetars, neutron stars

Are they sure?? Can they prove it?? how important are solitons in neutron starts??

^b clarification: Michta, Graziani, Bonitz, Contrib. Plasma Phys. **55**, 437 (2015), Moldabekov et al., Phys. Plasmas **25**, 031903 (2018), open access



clarification: Bonitz, Moldabekov, Ramazanov, Phys. Plasmas 26, 090601 (2019), perspectives article, Editors' pick, open access



^aManfredi, Haas, PRB 2001

QHD for plasmas: advancing quantum plasma research



enforce high quality standards: manuscripts should⁶

- 1. present an important advance to plasma physics;
- 2. contain a **convincing motivation** of the research in the introduction explaining the importance of the work;
- 3. present applications to concrete plasmas (in case of experiments) and—in case of theory—clear predictions for real plasmas, including their physical parameters;
- 4. not be purely formal studies of mathematical properties of equations, for example, in terms of dimensionless parameters, without application to real plasmas, as discussed in 3.

⁶M. Bonitz, Editorial: Contrib. Plasma Phys. 59 (1), 8 (2019)

Fluid description of quantum dynamics

N interacting identical quantum particles, hamiltonian

$$\hat{\mathcal{H}} = \sum_{i=1}^{N} \left(-\frac{\hbar^2}{2m} \nabla_i^2 + V(\mathbf{r}_i) \right) + \frac{1}{2} \sum_{i \neq j} w_{ij}(\mathbf{R}),$$

 $\mathbf{R} = (\mathbf{r}_1, \sigma_1; \mathbf{r}_2, \sigma_2; \dots \mathbf{r}_N, \sigma_N), \mathbf{r}_i$ are the particle coordinates and σ_i their spin projections.

► for a pure state: dynamics governed by the N-particle Schrödinger equation

$$i\hbar \frac{\partial \Psi(\mathbf{R},t)}{\partial t} = \hat{H} \Psi(\mathbf{R},t), \quad \Psi(\mathbf{R},t_0) = \Psi_0(\mathbf{R}),$$

normalization: $\sum_{\sigma_1...\sigma_N} \int d^{3N} R |\Psi(\mathbf{R}, t)|^2 = N$, for particles with spin *s*: $g_s = 2s + 1$ different spin projections

ansatz with real amplitude and phase (Bohm, Madelung)⁷:

$$\Psi(\mathbf{R},t) = A(\mathbf{R},t) e^{\frac{i}{\hbar}S(\mathbf{R},t)}, \qquad A, S \in \mathcal{R}.$$

▶ can be used for semiclassical molecular dynamics, e.g. Gericke, Gregori et al. (2018)

⁷E. Madelung, Z. Phys. (1927), D. Bohm, Phys. Rev. (1952)

Quantum hydrodynamics for 1 particle

hydrodynamic fields from wave function (for pure state):

$$n(\mathbf{r}, t) = A^{2}(\mathbf{r}, t), \qquad \text{density},$$

$$\mathbf{p}(\mathbf{r}, t) = m\mathbf{v}(\mathbf{r}, t) = \nabla S(\mathbf{r}, t), \qquad \text{momentum},$$

exact "hydrodynamic" equations:

$$\begin{aligned} \frac{\partial n}{\partial t} + \nabla(\mathbf{v}n) &= 0, \\ \frac{\partial \mathbf{p}}{\partial t} + \mathbf{v}\nabla \mathbf{p} &= -\nabla(V + Q), \\ \mathcal{Q}[n(\mathbf{r}, t)] &= -\frac{\hbar^2}{2m} \frac{\nabla^2 n^{1/2}}{n^{1/2}}, \end{aligned}$$
 "Bohm potential",

Modifications for N fermions:

- ▶ spin statistics (Pauli principle): *N* different orbitals $\phi_i \Rightarrow$ different n_i , \mathbf{p}_i
- interaction w_{ij} between particles
- statistical weight f_i of each orbital?
- coupling of orbitals? Definition of mean n and p?

Statistical description of N-particle quantum system⁸

mixed state described by N-particle density operator:

$$\hat{\rho}(t) = \sum_{a} p_{a} |\Psi^{a}(t)\rangle \langle \Psi^{a}(t)|, \quad \operatorname{Tr} \hat{\rho}(t) = 1,$$

sum over projection operators on all solutions of the Schrödinger equation, p_a : real probabilities, $0 \le p_a \le 1$, with $\sum_a p_a = 1$

one-particle reduced density operator:

$$\hat{F}_1(t) \equiv N \operatorname{Tr}_{2...N} \hat{\rho}(t), \quad \operatorname{Tr}_1 \hat{F}_1 = N.$$

• equation of F_1 : quantum kinetic equation, $\hat{F}_{12} = \hat{F}_1\hat{F}_2 + \hat{g}_{12}$:

$$i\hbar \frac{\partial \hat{F}_1}{\partial t} - [\hat{H}_1, \hat{F}_1] = \operatorname{Tr}_2[\hat{w}_{12}, \hat{g}_{12}] \equiv \hat{l}_1[\hat{g}_{12}], \quad \text{collision integral},$$
$$\hat{H}_1(t) = \hat{H}_1 + \hat{H}_1^{\mathrm{H}}(t), \quad \hat{H}_1^{\mathrm{H}}(t) \equiv \operatorname{Tr}_2 \hat{w}_{12} \hat{F}_2(t), \quad \text{mean field},$$

⁸M. Bonitz, "Quantum Kinetic Theory", 2nd ed., Springer 2016

Density matrix equation in Hartree approximation

mean field approximation (no anti-symmetrization, no correlations):

 $\Psi(\mathbf{R},t) \approx \phi_{\alpha_1}(\mathbf{r}_1) \cdot \phi_{\alpha_2}(\mathbf{r}_2) \cdots \phi_{\alpha_N}(\mathbf{r}_N), \quad \langle \phi_\alpha | \phi_\beta \rangle = \delta_{\alpha,\beta}$

 $\phi_{\alpha_k}(\mathbf{r}_k)$: single-particle orbital occupied by particle k,

for density operators: no antisymmetrization and $\hat{l}_1 = 0 \Rightarrow \hat{F}_{12} \approx \hat{F}_1 \hat{F}_2$ • coordinate representation: $\langle \mathbf{r}' | \hat{F}_1(t) | \mathbf{r}'' \rangle = f(\mathbf{r}', \mathbf{r}'', t)$, density matrix

$$\begin{split} i\hbar \frac{\partial}{\partial t} f(\mathbf{r}',\mathbf{r}'',t) &= -\frac{\hbar^2}{2m} \left(\nabla_{r'}^2 - \nabla_{r''}^2 \right) f(\mathbf{r}',\mathbf{r}'',t) \\ &+ \left\{ U^{\text{eff}}(\mathbf{r}',t) - U^{\text{eff}}(\mathbf{r}'',t) \right\} f(\mathbf{r}',\mathbf{r}'',t) \\ U^{\text{eff}}(\mathbf{r},t) &= V(\mathbf{r},t) + U^{\text{H}}(\mathbf{r},t), \\ U^{\text{H}}(\mathbf{r},t) &= g_{\text{s}} \int d\bar{r} w(r-\bar{r}) f(\bar{r},\bar{r},t), \end{split}$$

diagonal element of the density matrix: $f(\mathbf{r}, \mathbf{r}, t) = n(\mathbf{r}, t)$, density

Hartree approximation for fermions

• mean field approximation (no anti-symmetrization, no correlations):

$$\Psi(\mathbf{R},t) \approx \phi_{\alpha_1}(\mathbf{r}_1) \cdot \phi_{\alpha_2}(\mathbf{r}_2) \cdots \phi_{\alpha_N}(\mathbf{r}_N), \quad \langle \phi_\alpha | \phi_\beta \rangle = \delta_{\alpha,\beta}$$

Thermodynamic equilibrium: orbitals $\phi_1, \phi_2 \dots$ occupied with probability $f_i = [e^{\beta(\epsilon_i - \mu)} + 1]^{-1}, i = 1 \dots \infty$ • One-particle density operator and density matrix:

$$\hat{F}_{1} = N \operatorname{Tr}_{2...N} \sum_{\alpha} \frac{1}{Z_{G}} e^{-\beta(E_{\alpha} - \mu N_{\alpha})} |\psi_{\alpha}\rangle \langle\psi_{\alpha}| = N \sum_{i=1}^{\infty} f_{i} |\phi_{i}\rangle \langle\phi_{i}$$
$$f(\mathbf{r}', \mathbf{r}'', t) = N \sum_{i=1}^{\infty} f_{i} \phi_{i}(\mathbf{r}', t) \phi_{i}^{*}(\mathbf{r}'', t) ,$$

Equation for f solved by set of nonlinear "Hartree-"Schrödinger (t-dependent Kohn-Sham) equations:

$$i\hbar\frac{\partial}{\partial t}\phi_i(\mathbf{r},t) = \left\{-\frac{\hbar^2}{2m}\nabla_{\mathbf{r}}^2 + V(\mathbf{r},t) + U_F^{\rm H}(\mathbf{r},t)\right\}\phi_i(\mathbf{r},t),$$
$$U_F^{\rm H}[n(\mathbf{r},t)] = \int d\mathbf{r}_2 w(\mathbf{r}-\mathbf{r}_2)[g_s n(\mathbf{r}_2,t) - n_0],$$

 $U_F^{
m H}$: Hartree mean field, $n({f r},t)=\sum_{i=1}^\infty f_i\,|\phi_i({f r},t)|^2$. n_0 : background

Microscopic QHD equations (MQHD)¹⁰

- ▶ in TDSE: introduce real amplitude ($A_i = n_i^{1/2}$) and phase for each orbital:
- ► restore exchange-correlation effects formally exactly⁹: *either* i) potential $V^{xc} = V^{xc}[n(\mathbf{r}, \tilde{t})]$, from TD-DFT, *or, alternatively,* ii) collision integral $l_i[g_{12}(\tilde{t})]$ from quantum kinetic theory

$$\begin{aligned} \frac{\partial n_i}{\partial t} + \nabla(\mathbf{v}_i n_i) &= 0, \\ \frac{\partial \mathbf{p}_i}{\partial t} + \mathbf{v}_i \operatorname{div} \mathbf{p}_i &= -\nabla \left(V(\mathbf{r}, t) + U^{\mathrm{H}} + Q_i + V^{\mathrm{xc}} + I_i \right), \\ U^{\mathrm{H}}[n(\mathbf{r}, t)] &= \int d\mathbf{r}_2 \, w(\mathbf{r} - \mathbf{r}_2) [g_s n(\mathbf{r}_2, t) - n_0], \\ Q_i(\mathbf{r}, t) &= -\frac{\hbar^2}{2m} \frac{\nabla^2 \sqrt{n_i(\mathbf{r}, t)}}{\sqrt{n_i(\mathbf{r}, t)}}. \end{aligned}$$

• $n(\mathbf{r}, t) = \sum_{i=1}^{\infty} f_i \cdot |\phi_i(\mathbf{r}, t)|^2$, f_i statistical weight (Fermi function), $g_s = 2s + 1$.

► ⇒ exact formulation of the many-body problem (with proper set of orbitals), equivalent to TD-DFT

⁹in general: memory effects (non-adiabatic), i.e. $t_0 \leq \tilde{t} \leq t$

¹⁰Bonitz et al., Phys. Plasmas 26, 090601 (2019), agrees with Manfredi, Fields Inst. Comm. 2005

Derivation of QHD equations from MQHD¹¹

- ▶ consider mean field (Hartree / quantum Vlasov) approximation: $V^{xc} \equiv I_i \equiv 0$
- average MQHD equations over i with weights f_i
- define hydrodynamic (average) quantities:

$$\overline{n}(\mathbf{r},t) = \frac{1}{N} \sum_{i=1}^{\infty} f_i n_i(\mathbf{r},t),$$

$$\overline{\mathbf{p}}(\mathbf{r},t) = \frac{1}{N} \sum_{i=1}^{\infty} f_i \mathbf{p}_i(\mathbf{r},t),$$

$$\overline{Q}(\mathbf{r},t) = -\frac{\hbar^2}{2mN} \sum_{i=1}^{\infty} f_i \frac{\nabla^2 \sqrt{n_i(\mathbf{r})}}{\sqrt{n_i(\mathbf{r})}} \neq \frac{\hbar^2}{2mN} \frac{\nabla^2 \sqrt{\overline{n}(\mathbf{r})}}{\sqrt{\overline{n}(\mathbf{r})}} \equiv Q_1[\overline{n}].$$

- final equality was *postulated* by Manfredi and Haas (PRB 2001, Fields Inst. Comm. 2005) reproducing single-particle QHD equations (with Hartree potential)
- ▶ When does it hold? What are the corrections? ⇒ perform strict derivation

¹¹Bonitz et al., Phys. Plasmas 2019

Derivation of QHD equations (contd.)¹²

- use average of product: $\overline{a_i b_i} = \overline{a} \cdot \overline{b} + \overline{\delta a_i \delta b_i}$,
- result: formally exact QHD equations from orbital average of MQHD (TD-DFT):

$$\begin{aligned} \frac{\partial \overline{n}}{\partial t} &+ \frac{1}{m} \nabla (\overline{\mathbf{p}} \cdot \overline{n}) = J_{np}^{\Delta} \\ \frac{\partial \overline{\mathbf{p}}}{\partial t} &+ \frac{1}{m} \overline{\mathbf{p}} \cdot \operatorname{div} \overline{\mathbf{p}} = -\nabla \left(V(\mathbf{r}, t) + U_{F}^{\mathrm{H}}[\overline{n}] + Q_{1}[\overline{n}] + Q^{\Delta} + V^{\mathrm{xc}}[\overline{n}] \right) + J_{pp}^{\Delta} \end{aligned}$$

three correction terms (correlation functions):

$$\begin{split} J^{\Delta}_{np} &= -\frac{1}{m} \nabla \,\overline{\delta \mathbf{p}_i \delta n_i}, \quad Q^{\Delta} \approx \frac{\hbar^2}{2m\overline{n}} \overline{\delta A_i \cdot \nabla^2 \delta A_i} + O\left(\left(\frac{\delta A_i}{\overline{A}}\right)^2\right), \\ J^{\Delta}_{pp} &= -\frac{1}{m} \overline{\delta \mathbf{p}_i \mathrm{div} \, \delta \mathbf{p}_i} = \frac{1}{\overline{n}} \partial_{\beta} \overline{P}_{\alpha\beta} = \frac{1}{\overline{n}} \nabla \overline{P}_F + \frac{1}{\overline{n}} \partial_{\gamma} \overline{\sigma}_{\alpha\gamma}, \quad \gamma \neq \alpha, \quad \overline{\sigma}: \text{stress tensor}, \end{split}$$

▶ special case: ideal Fermi gas of dimension D = 1, 2, 3, T = 0: $\overline{\sigma}_{\alpha\gamma} \rightarrow 0$ and

$$J^{\Delta}_{
hop} = rac{1}{D}
abla \overline{\delta E_{
m kin}} = rac{m}{2(D+2)}
abla v^2_F = rac{1}{2\overline{n}}
abla \overline{P}^{
m id} \, .$$

12 Bonitz et al., Phys. Plasmas 2019

Test case plasma oscillations: linearization of MQHD-equations¹³

stability conditions without excitation:

$$\partial_t n_{i0} = \partial_t \mathbf{p}_{i0} = \overline{\mathbf{v}}_{i0} = \mathbf{0}$$

 $\mathbf{v}_{i0} \operatorname{div} \mathbf{p}_{i0} = -\nabla (U_0^{\mathrm{H}} + Q_{i0})$

• monochromatic weak excitation: $V_1^{\text{ext}}(\mathbf{r}, t) = \tilde{V}_1^{\text{ext}} e^{-i\hat{\omega}t + i\mathbf{qr}}$, where $\hat{\omega} = \omega + i\epsilon$, $\epsilon > 0$

Inearization and Fourier-Laplace transform:

$$\begin{array}{l} n_{i0} \to n_{i0} + n_{i1} ; \mathbf{v}_{i0} \to \mathbf{v}_{i0} + \mathbf{v}_{i1} ; \\ U_0^{\rm H} \to U_0^{\rm H} + U_1^{\rm H} ; Q_{i0} \to Q_{i0} + Q_{i1} , \quad U_1^{\rm eff} = V_1^{\rm ext} + U_1^{\rm H} \end{array}$$

density response exactly coincides with random phase approximation (RPA)

$$\tilde{n}_{1}(\mathbf{q},\hat{\omega}) = \frac{1}{N} \sum_{i=1}^{\infty} f_{i} \tilde{n}_{i1}(\mathbf{q},\hat{\omega}) = \tilde{U}_{1}^{\text{eff}}(\mathbf{q},\hat{\omega}) \tilde{\Pi}_{1}^{R}(\mathbf{q},\hat{\omega})$$
$$\tilde{\Pi}_{1}^{R}(\mathbf{q},\hat{\omega}) = \frac{1}{N} \sum_{i=1}^{\infty} \frac{f_{i}}{(\hat{\omega} - \mathbf{q}\mathbf{v}_{i0})^{2} - \frac{\hbar^{2}q^{4}}{4m^{2}}}.$$

¹³Bonitz et al., Phys. Plasmas 2019, result agrees with Manfredi, Fields Inst. Comm. 2005

Test case plasma oscillations: linearization of QHD equations¹⁵

- Hartree approximation, neglect correlation terms
- Inearization and Fourier-Laplace transform:

$$\begin{split} \overline{n}_0 &\rightarrow \overline{n}_0 + \overline{n}_1 ; \ \overline{\mathbf{v}}_0 \rightarrow \overline{\mathbf{v}}_0 + \overline{\mathbf{v}}_1 ; \ U_0^{\rm H} \rightarrow U_0^{\rm H} + U_1^{\rm H} ; \\ \mathcal{Q}[\overline{n}] &\rightarrow -\frac{\hbar^2}{2m} \frac{\nabla^2 \sqrt{\overline{n}_0}}{\sqrt{\overline{n}_0}} + \frac{\hbar^2}{4m} \nabla^2 \sqrt{\overline{n}_1} , \\ \overline{P}_0^{\rm id} &\rightarrow \frac{2}{D+2} \overline{n}_0 \mathcal{E}_F(\overline{n}_0) + \frac{2}{D} \frac{\overline{n}_1}{\overline{n}_0} \mathcal{E}_F(\overline{n}_0) , \end{split}$$

▶ plasmon dispersion – QHD vs. MQHD (RPA): agreement only in 1D, for $\omega \ge \omega_{pl}$

$$egin{aligned} &\omega_{ ext{QHD}}^2(q) = \omega_{
ho l}^2 + rac{1}{D} v_F^2 q^2 + rac{\hbar^2}{4m^2} q^4, \ &\omega_{ ext{MQHD}}^2(q) = \omega_{
ho l}^2 + rac{3}{D+2} v_F^2 q^2 + (1-\delta_{2,D}) rac{\hbar^2}{4m^2} q^4, \end{aligned}$$

• for $\omega < \omega_{pl}$: Bohm potential Q is factor 9 too big \Rightarrow incorrect acoustic modes and statically screened potential¹⁴

¹⁴Michta *et al.*, CPP **55**, 437 (2015), Moldabekov *et al.*, Phys. Plasmas **25**, 031903 (2018)

¹⁵Bonitz et al., Phys. Plasmas 26, 090601 (2019)

Correction of coefficients in QHD equations¹⁶

• Comparison to RPA reveals ω - and k-dependence of pressure ($\bar{\alpha}$) and Bohm potential (γ). In addition: dependence on temperature, dimensionality



 $^{^{16}}$ Figure for 3D case, T = 0, from: Moldabekov, Bonitz, and Ramazanov, Phys. Plasmas **25**, 031903 (2018)

Summary

- ► ab initio QMC simulations provide complete thermodynamic data for warm dense uniform electron gas¹⁷
- accurate functional $f_{xc}(r_s, \Theta, \xi)$ input for finite-T LDA-DFT, implemented in Libxc (LDA_XC_GDSMFB)
- ab initio data for inhomogeneous EG¹⁸ \Rightarrow accurate parametrization of static local field correction¹⁹ G(q)
- first *ab initio* data for the dynamic structure factor $S(q, \omega)$ of warm dense electrons²⁰

¹⁷T. Dornheim *et al.*, Phys. Reports (2018)

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- Electronic correlations and correlation build up (thermalization, dynamical screening, Auger processes etc.) are captured by (Nonequilibrium) Green functions. Highly efficient new computational techniques available²²

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Outlook: simulating WDM out of equilibrium-combination of methods promising



Figure: modified from M. Bonitz *et al., Front. Chem. Science Engin.* (2019); QBE: quantum Boltzmann equation, NEGF: Nonequilibrium Green functions

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- DFT-MD and TD-DFT capture atomic scales and inhomogeneity effects. But: miss electronic correlations (inaccurate band gap, ionization energies, no Auger processes etc.)



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- Green functions capture electronic correlations, correlation build up and electron thermalization



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